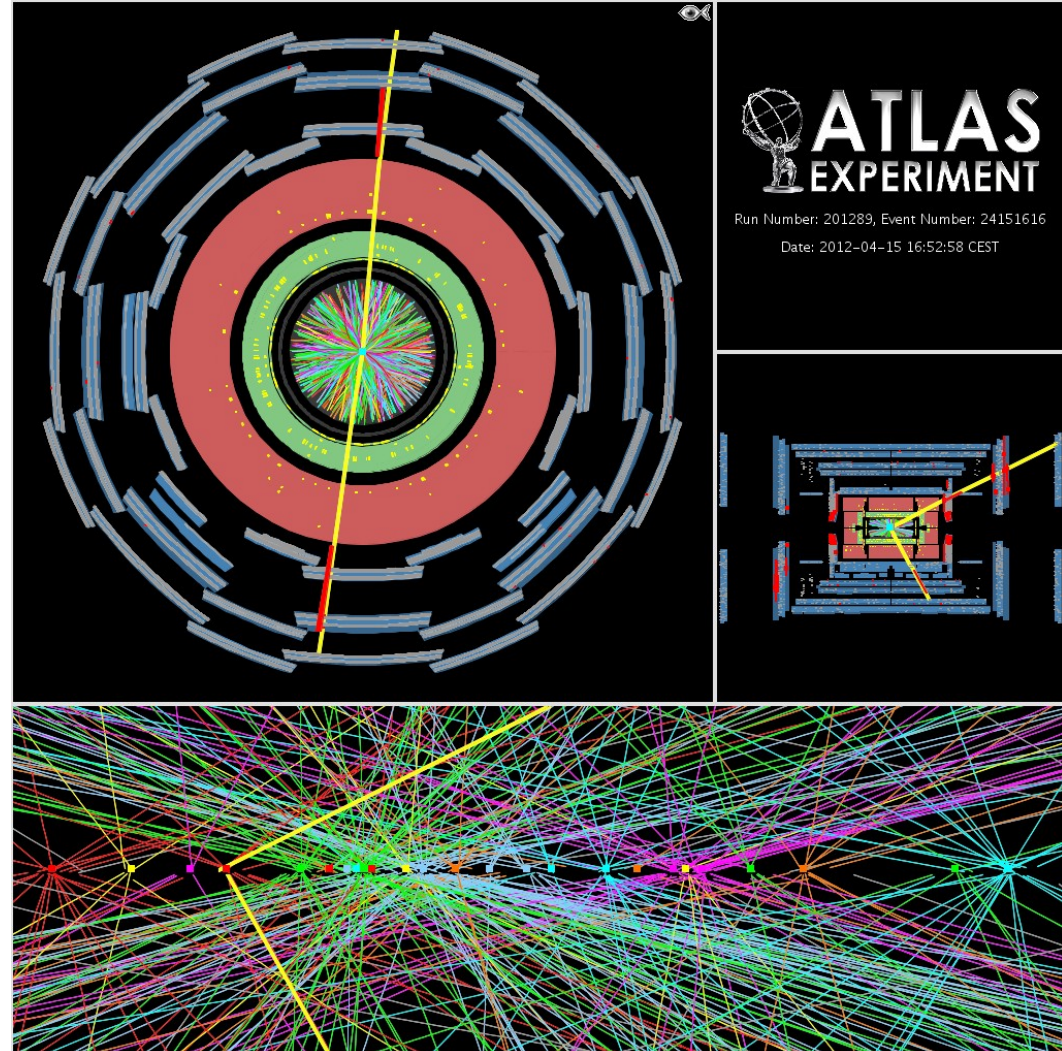


Recherche directe de nouvelle physique au LHC

Henri Bachacou
IRFU

École de Gif
Gif-sur-Yvette
20-21 sept. 2012



Recherche directe de nouvelle physique au LHC

- Focus on ATLAS and CMS direct searches
- Apologies for the strong ATLAS bias...
- Not a comprehensive overview: I chose a few themes
- Much more information available here:
 - <https://twiki.cern.ch/twiki/bin/view/AtlasPublic>
 - <https://twiki.cern.ch/twiki/bin/view/CMSPublic/PhysicsResults>
- See other lectures on related topics:
 - SM and BSM theories
 - Heavy Flavour
 - SM and MSSM Higgs

Outline

■ Introduction

- The Standard Model and its limitations
- The main Beyond-Standard-Model candidates to cope with the SM limitations
- General methodology at the LHC
- The LHC and the ATLAS and CMS detectors

■ Supersymmetry *

- “standard” SUSY: MSUGRA/CMSSM
- “compressed” SUSY
- “natural” SUSY: dedicated stop/sbottom/gaugino searches

■ Long-lived particles (not only SUSY)

■ Extra-dimensions and heavy resonances

■ 4th generation and heavy quarks

■ Model-independent searches

■ Outlook and conclusions

* Thanks to P. Pralavorio for help with SUSY

today

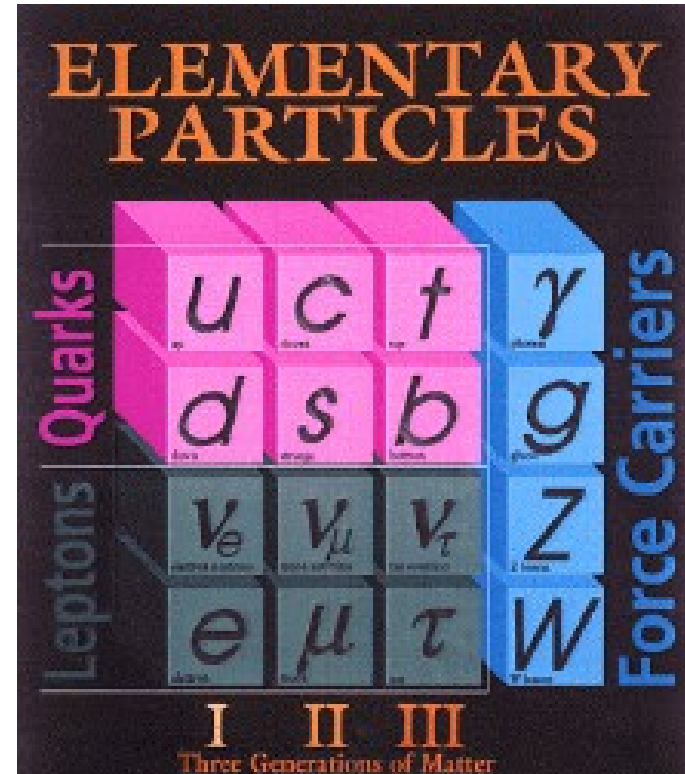
tomorrow

Introduction

- The Standard Model and its limitations
- The main Beyond-Standard-Model candidates to cope with the SM limitations
- General methodology at the LHC
- The LHC and the ATLAS and CMS detectors

The Standard Model in 1 slide: 3 forces, 3 generations, 1 Higgs boson

- Strong, Weak, and Electromagnetic forces described by gauge theory:
→ $SU(3)_{\text{QCD}} \times SU(2)_L \times U(1)$
- Force carriers: photon, gluons, W and Z bosons
- $SU(2)_L \times U(1)$ spontaneously broken by Higgs mechanism to give mass to W and Z bosons
- CKM matrix describes mixing of quarks
- **Incredible success of the SM: explains (almost) all phenomena of particle physics from 0 to 1 TeV !**

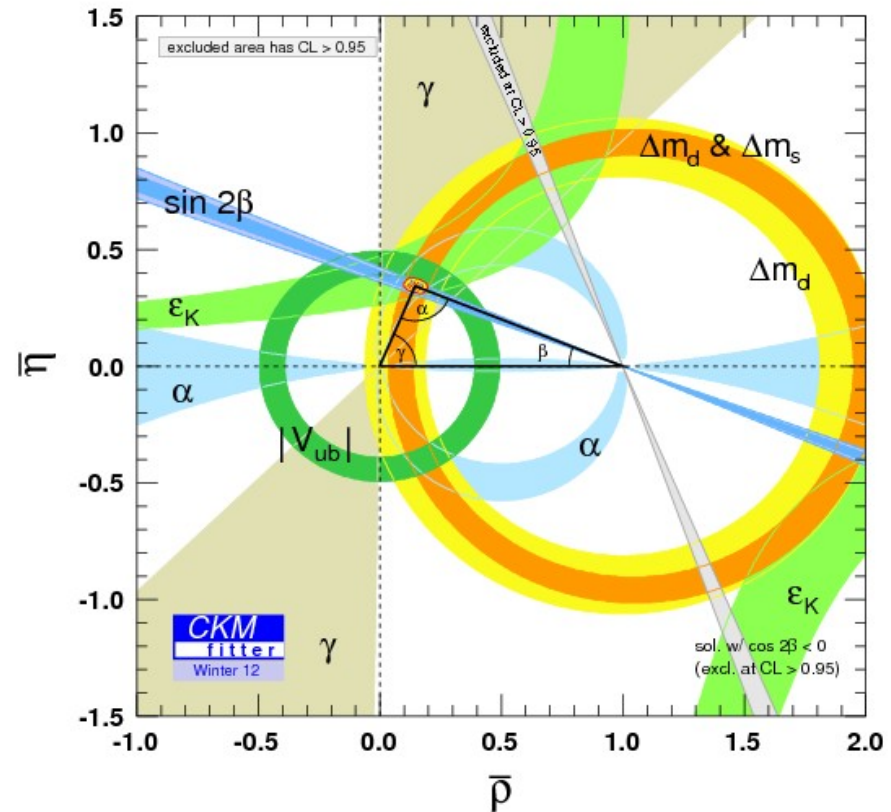
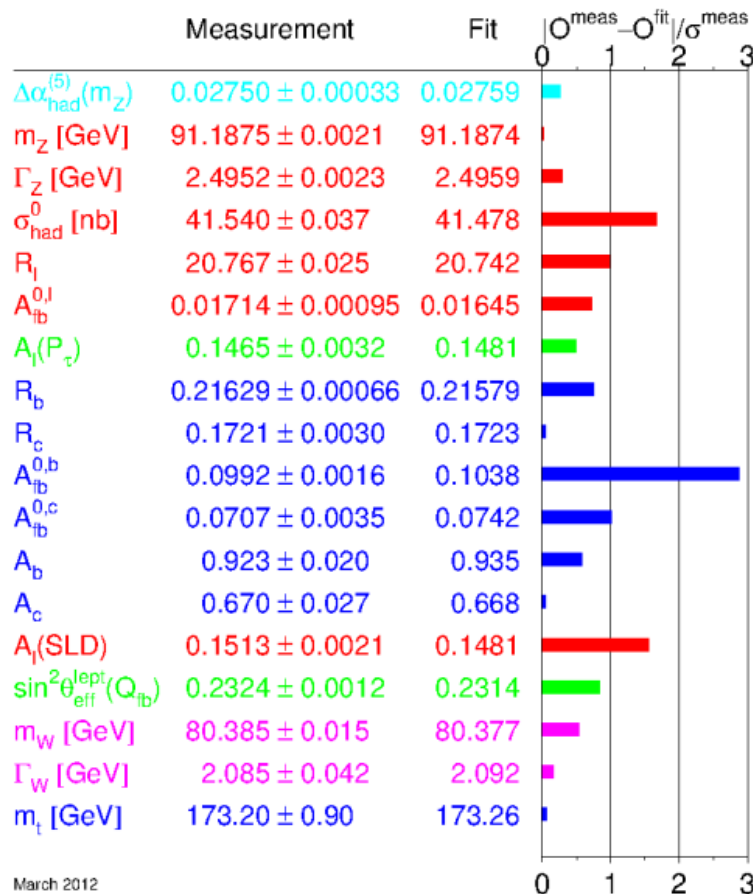


$$V(\Phi^+\Phi) = \mu^2 \Phi^+\Phi + \frac{\lambda}{2} (\Phi^+\Phi)^2$$

$$v = \sqrt{\frac{-\mu^2}{\lambda}} = 246 \text{ GeV}$$

Success of the Standard Model: EW and CP-violation Precision Measurements

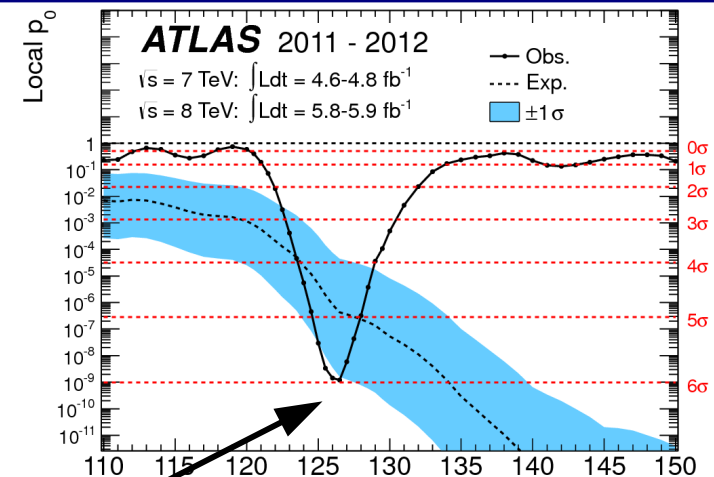
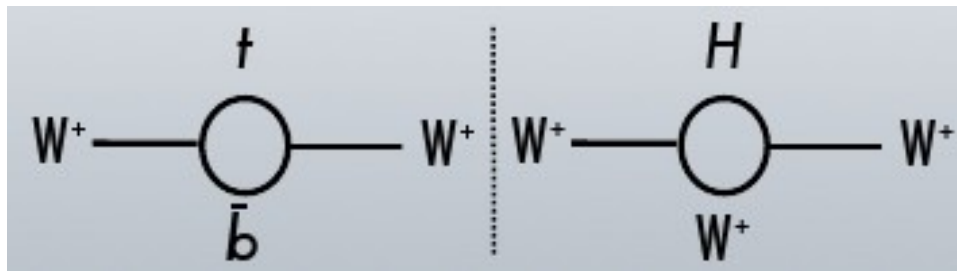
- Electroweak measurements at LEP, SLD, and Tevatron
- CP-violation at B-factories, K-factories



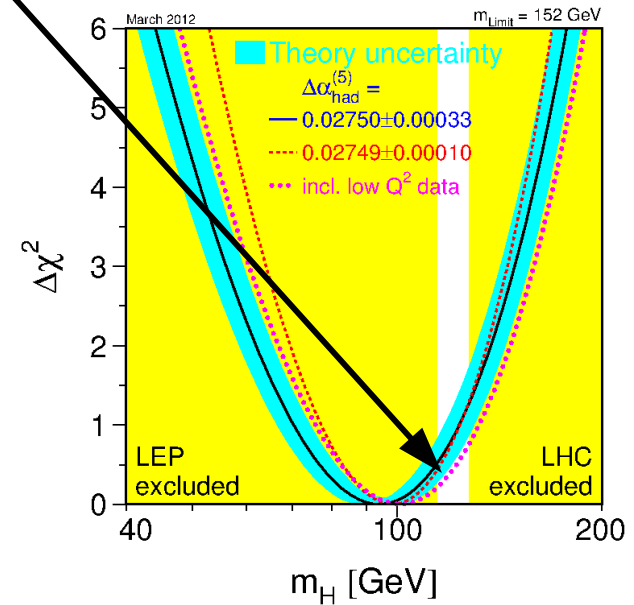
March 2012

Latest Success of the Standard Model: The Higgs Boson !

- Precision measurements at LEP, SLD, and Tevatron (Z pole, W mass, top mass) → constraint on Higgs mass
- A light Higgs is preferred
- LHC Higgs(-like) discovery at ~ 126 GeV is compatible with precision measurements



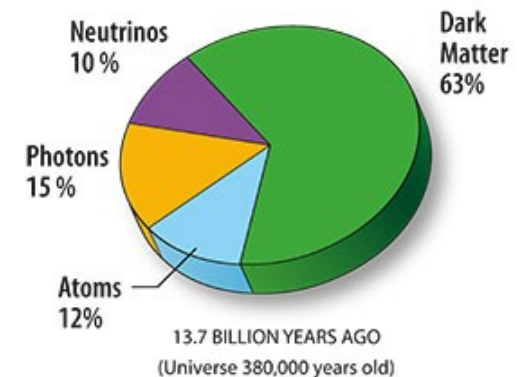
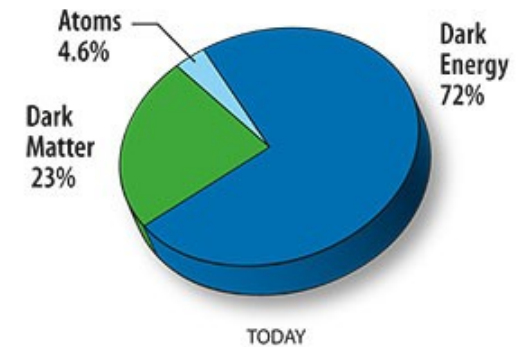
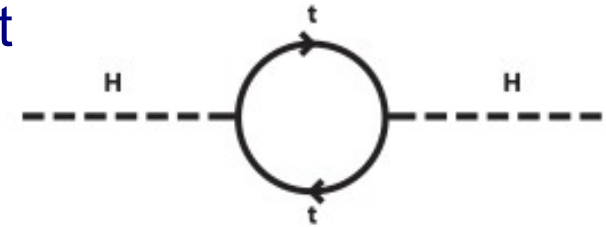
126 GeV



Why look “beyond” the Standard Model?

Some limitations of the Standard Model

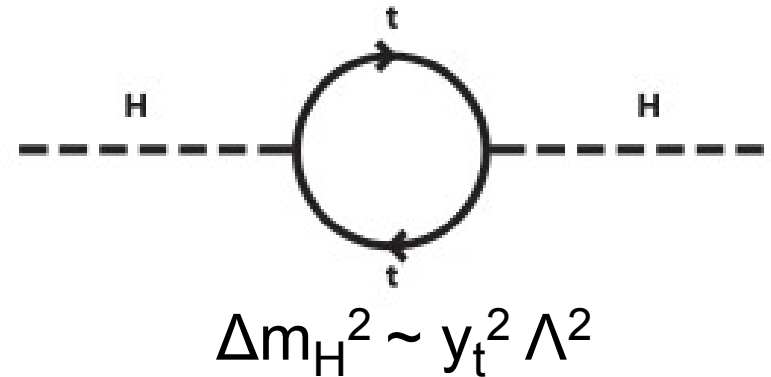
- The Standard Model is an effective theory that must break down at a certain scale
 - Hierarchy: quadratic divergence of the Higgs mass, extremely fine-tuned
 - What is the underlying nature of EWSB?
- Dark Matter and Dark Energy
 - cannot be explained by SM
- Neutrinos have mass
 - where are the right-handed neutrinos?
- How to include gravitation?
- How to produce enough CP-violation to explain the Universe matter-antimatter asymmetry? 4th generation of fermions?
- BSM models try to answer limitations of the SM



Why look “beyond” the Standard Model?

The Hierarchy Problem

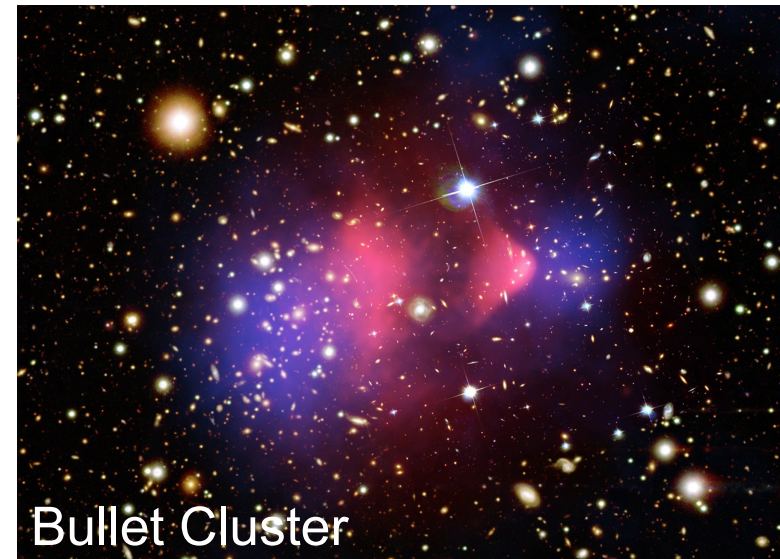
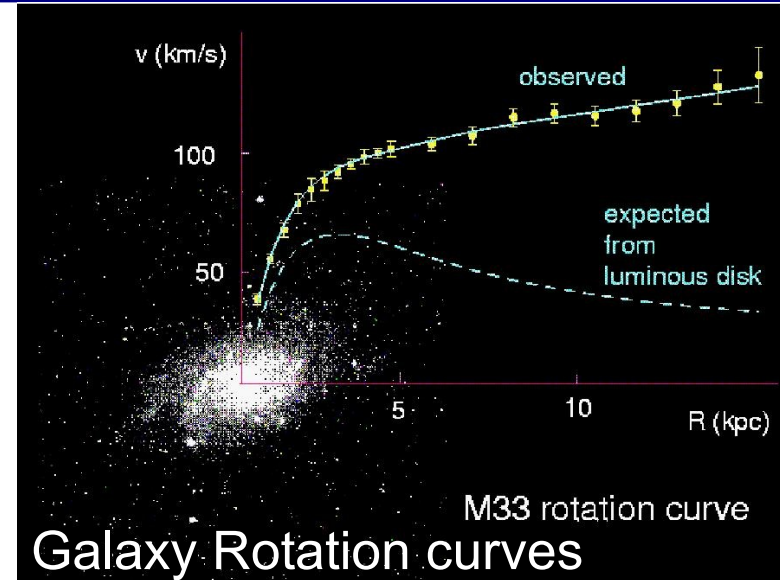
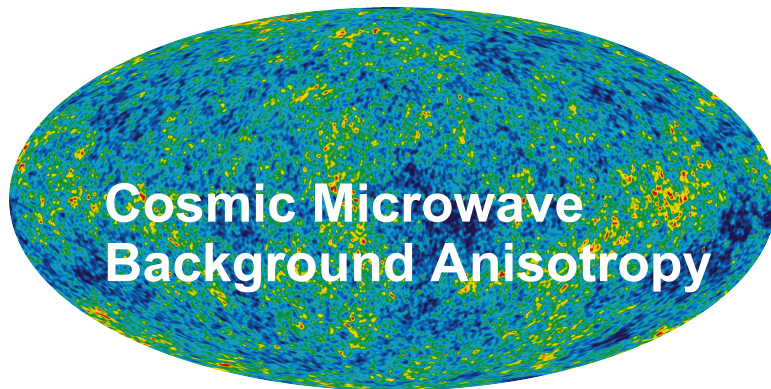
- Loop corrections to Higgs mass lead to quadratic divergences
- Hierarchy problem:
Electroweak scale ~ 100 GeV
Plank scale $\sim 10^{19}$ GeV
17 orders of magnitude!
- Fine tuning:
bare mass (very large)
– radiative corrections (very large)
= 126 GeV
- Can new physics cancel the quadratic divergences?



Why look “beyond” the Standard Model?

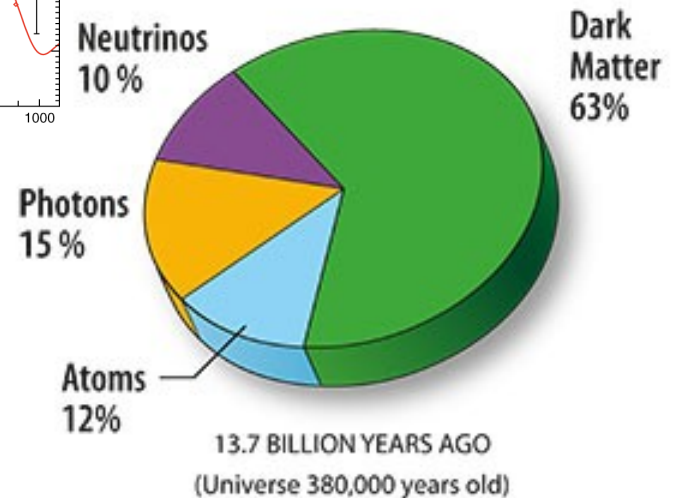
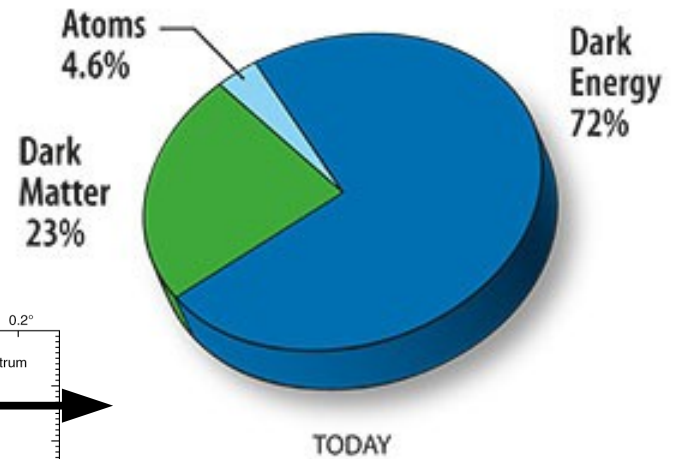
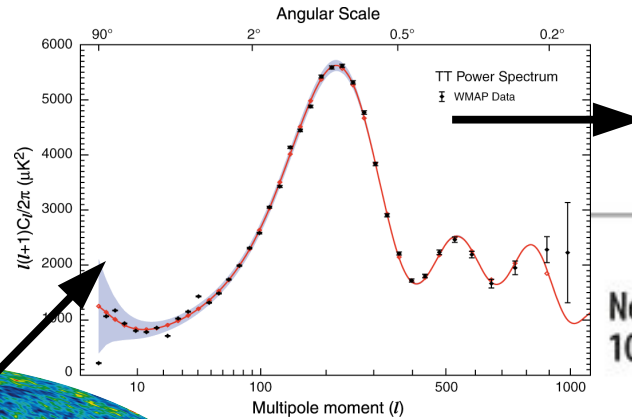
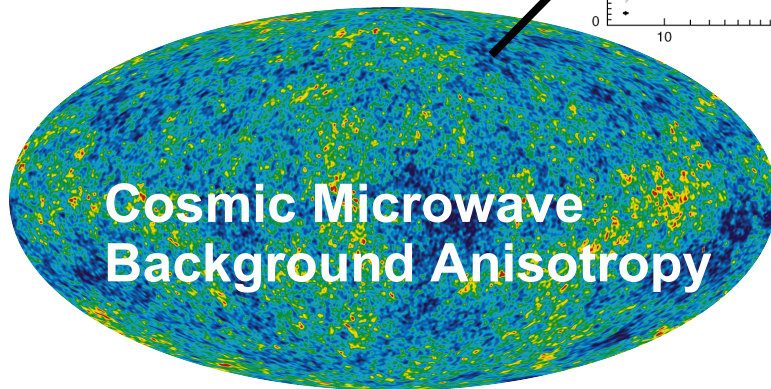
Dark Matter

- Dark Matter has been observed in several ways:
 - Galaxy rotation curves
 - Gravity lensing
 - CMB anisotropy
- BSM phenomena have already been observed!



Why look “beyond” the Standard Model? Dark Matter

- ~ 80% of matter is Dark Matter
- **Weakly Interacting Massive Particles?**
(WIMPS)

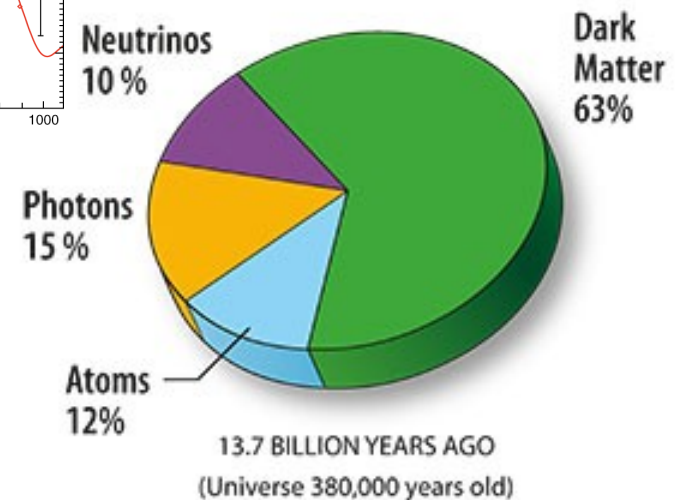
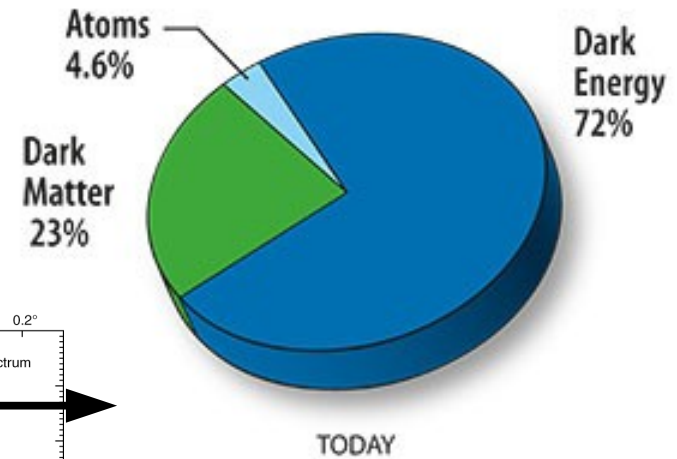
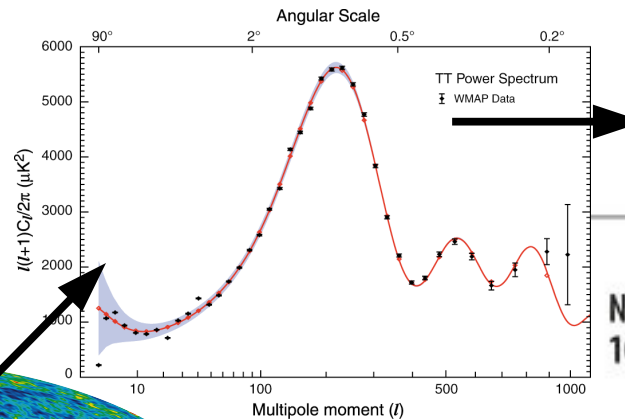
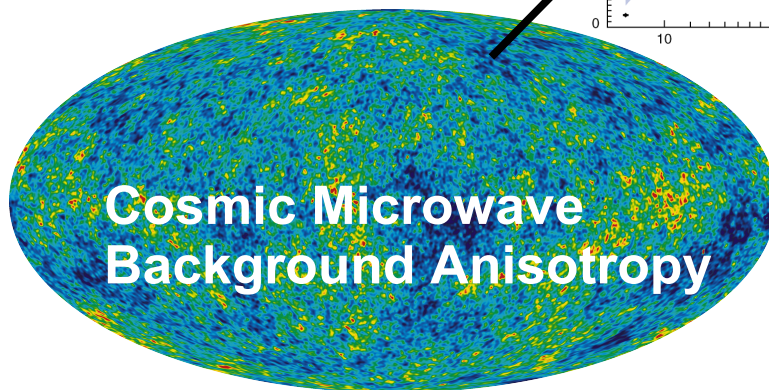


Why look “beyond” the Standard Model?

Dark Energy

- ~ 70% of energy is Dark Energy

???



Why look “beyond” the Standard Model?

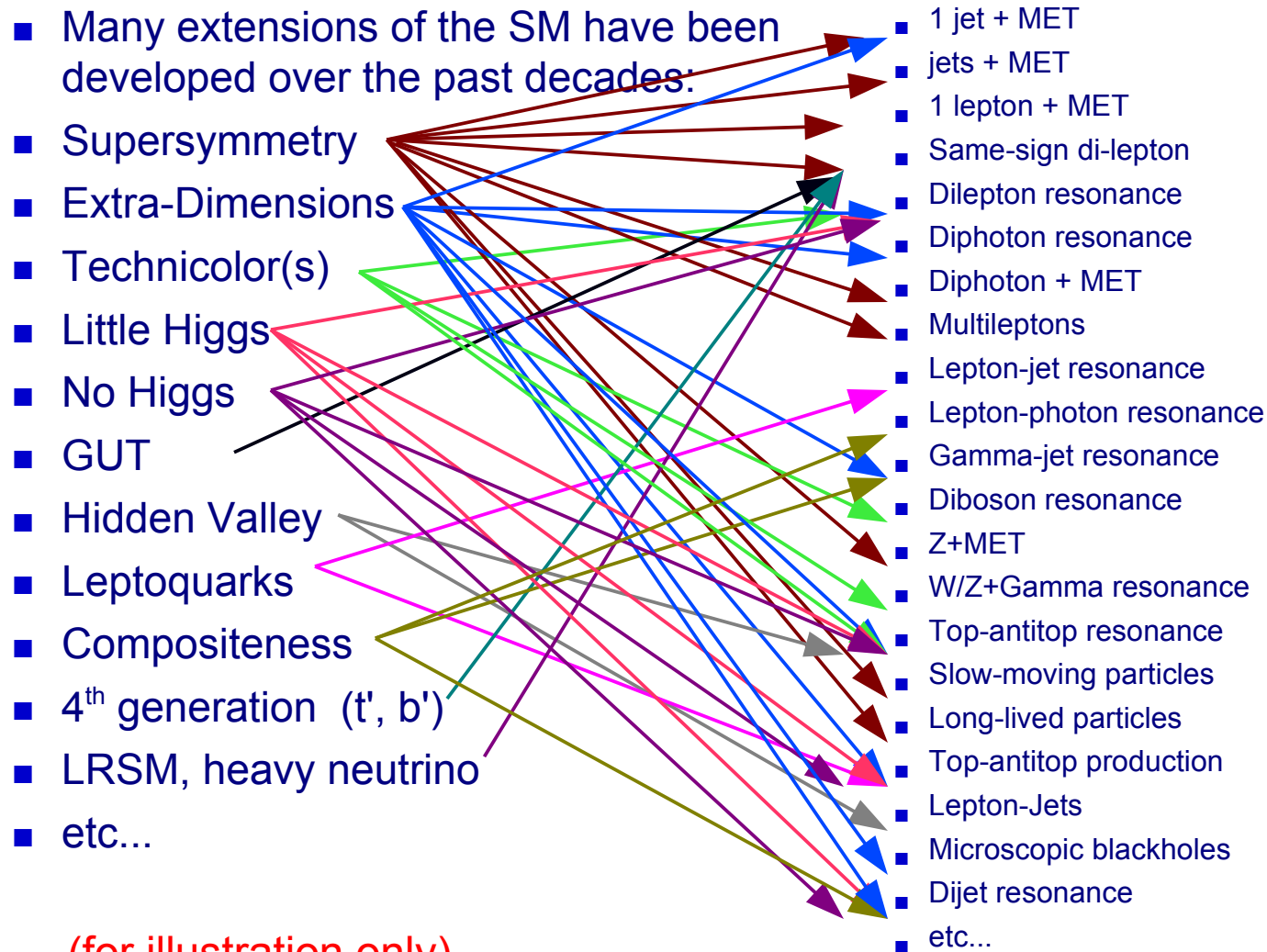
The main candidates

- Many BSM models developed to answer SM limitations:
- **Supersymmetry**: addresses Hierarchy Problem, offers good Dark Matter candidates
- **Technicolor**: explains underlying nature of EW SB through strong interaction *
- **Extra-dimensions**: addresses Hierarchy Problem by bringing the Plank scale down to TeV (i.e. explains why gravity is so weak)
- **4th generation models**: new source of CP-violation capable of explaining Matter-Antimatter asymmetry *
- Etc...

* but have a hard time explaining the recent observation at 126 GeV

How do we look for BSM at the LHC?

A very long list of models x signatures



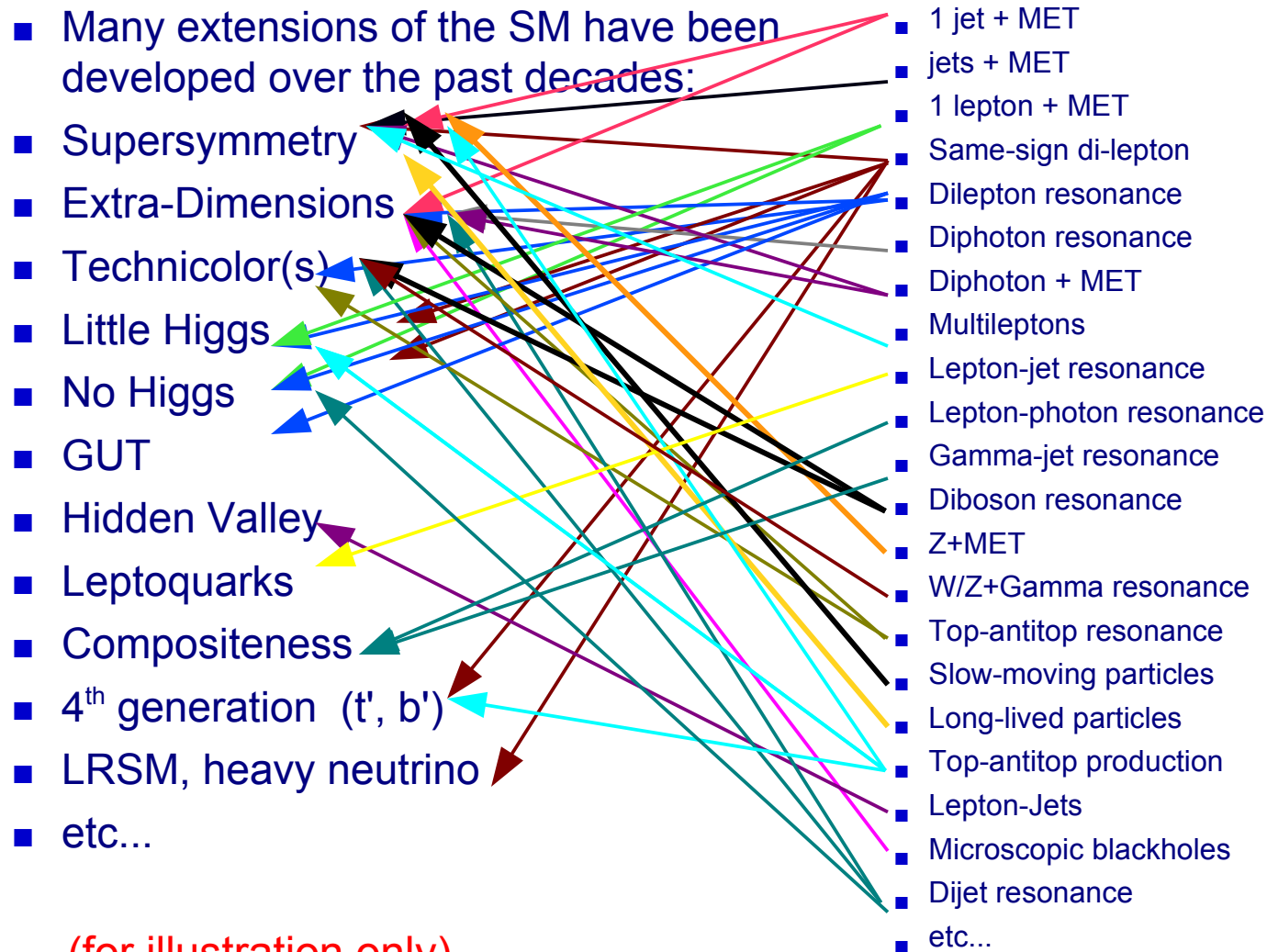
(for illustration only)

Many models leading to many signatures!

Must proceed methodically (?)

How do we look for BSM at the LHC?

A very long list of models x signatures



(for illustration only)

A complex 2D problem

Experimentally, a **signature standpoint** makes a lot of sense:

- Practical
- Less model-dependent
- Important to cover every possible signature

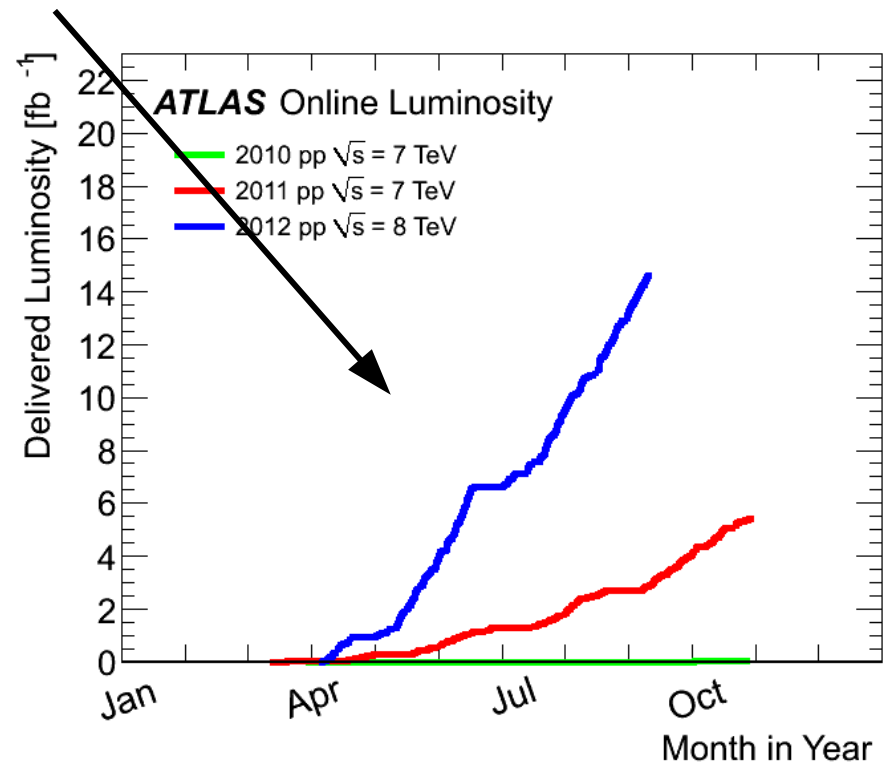
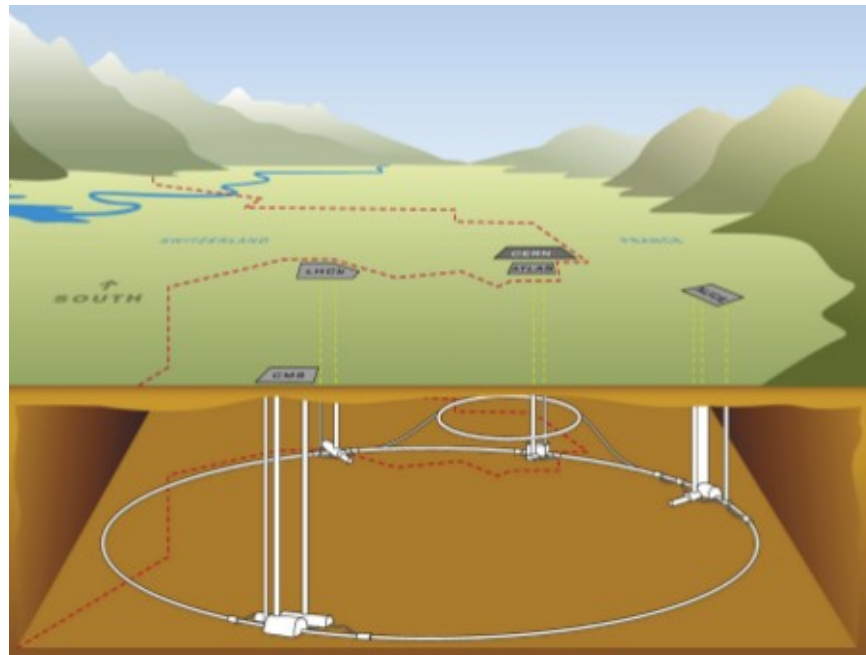
How do we look for BSM at the LHC?

Looking for models or for signatures?

- Several types of searches:
- **Very model-specific:**
 - Ex: monopole search
 - Look for highly-ionizing particle that bends in the wrong direction
- **Very model-independent:**
 - Ex: inclusive search for same-sign leptons
- **Somewhere In-between (most analyses):**
 - motivated by a particular model, but tries to remain inclusive
 - Ex: SUSY analysis looking for final state with jets and missing ET

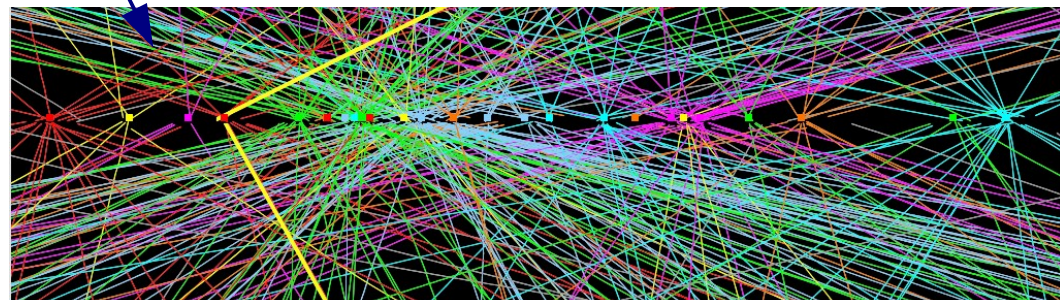
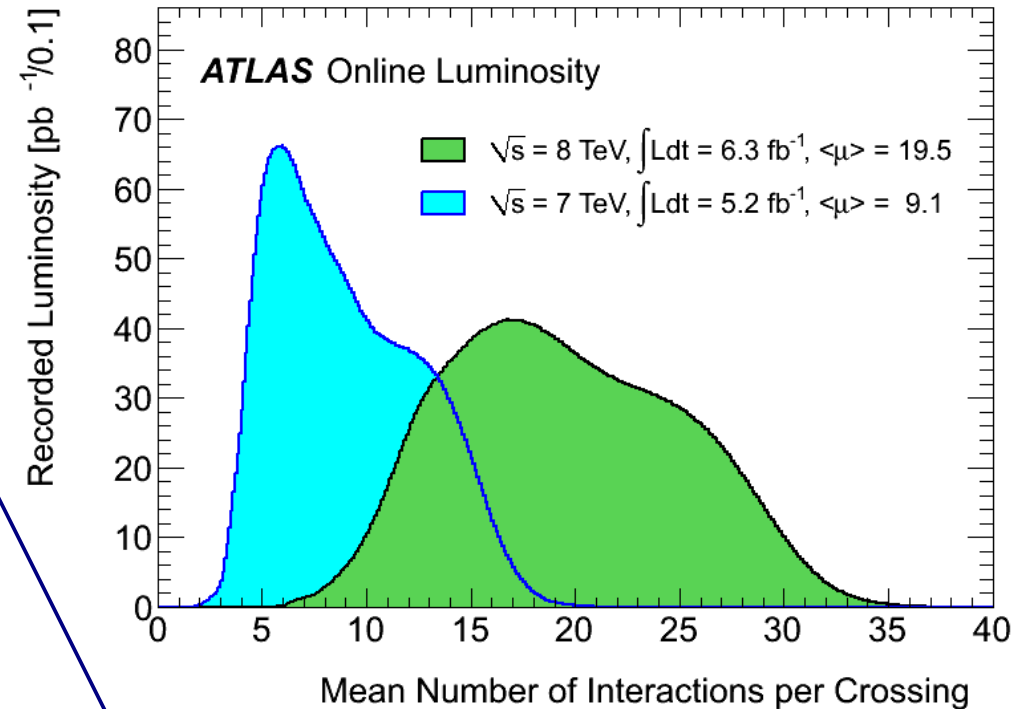
The Large Hadron Collider (LHC)

- LHC has performed extremely well:
 - 5.3 fb^{-1} at $\sqrt{s} = 7 \text{ TeV}$ delivered in 2011
 - Already $\sim 15 \text{ fb}^{-1}$ at $\sqrt{s} = 8 \text{ TeV}$ delivered in 2012, expect close to 30 fb^{-1} by the end of the the year!
- Past 2 years: significant gain every 6 months: difficult to keep up!

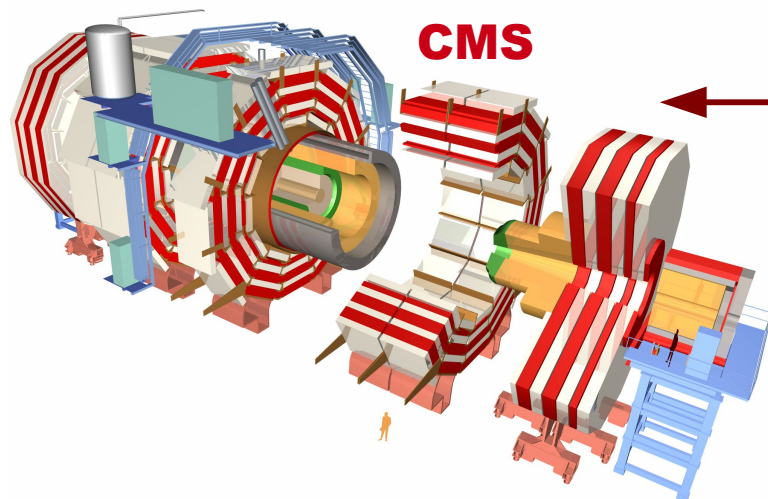


Pile-up: several proton-proton interactions in each bunch-crossing

- $Z \rightarrow \mu\mu$ event... and 24 other collisions from the same bunch-crossing (a.k.a. “in-time pile-up”)
- Due to small bunch spacing (50 ns), also effect from collisions from previous and next crossings (a.k.a. “out-of-time pile-up”)

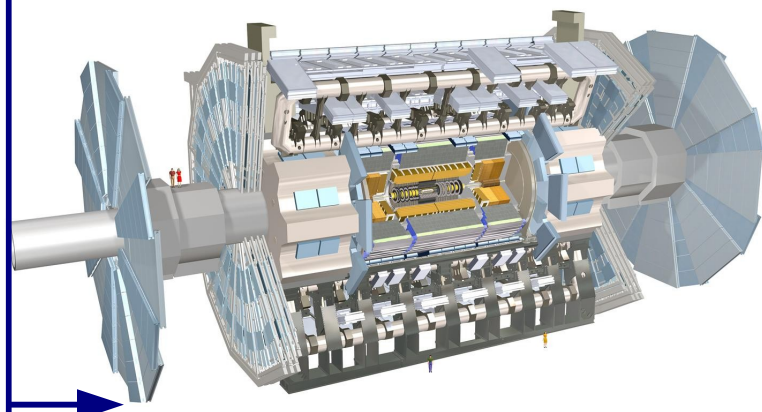


The ATLAS and CMS Detectors: same goals, different choices



- 3.8T solenoid containing calorimeters
- Silicon tracker: $\sigma(p_T)/p_T \sim 15\%$ at 1TeV
- EM cal: homogeneous Lead-Tungstate crystal, $\sigma_E/E \sim 3\%/\sqrt{E[\text{GeV}]} \oplus 0.5\%$
- HAD cal: Brass-scint., $\geq 7\lambda_0$
 $\sigma_E/E \sim 100\%/\sqrt{E[\text{GeV}]} \oplus 5\%$
- Iron return yoke muon spectrometer

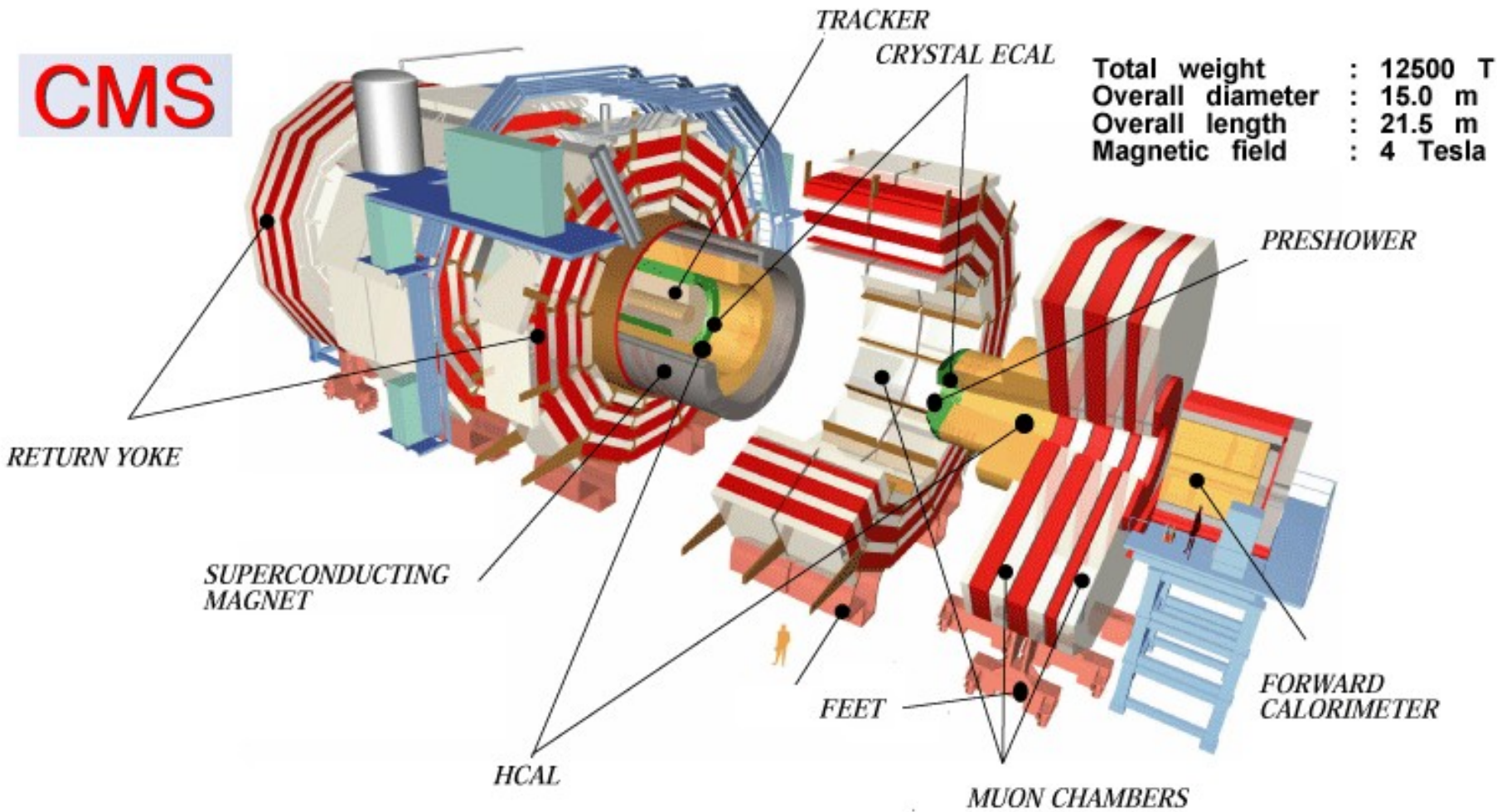
- 2T solenoid inside calorimeters
- Silicon+TRT tracker + electron ID
- EM cal: Longitudinally segmented Lead-Ar: $\sigma_E/E \sim 10\%/\sqrt{E[\text{GeV}]} \oplus 0.7\%$
- HAD cal: Fe-scint + Cu-Ar, $\geq 11\lambda_0$
 $\sigma_E/E \sim 50\%/\sqrt{E[\text{GeV}]} \oplus 3\%$
- Air-toroid muon sp.: $\int \sqrt{B \cdot dl} = 1$ to 7 T.m



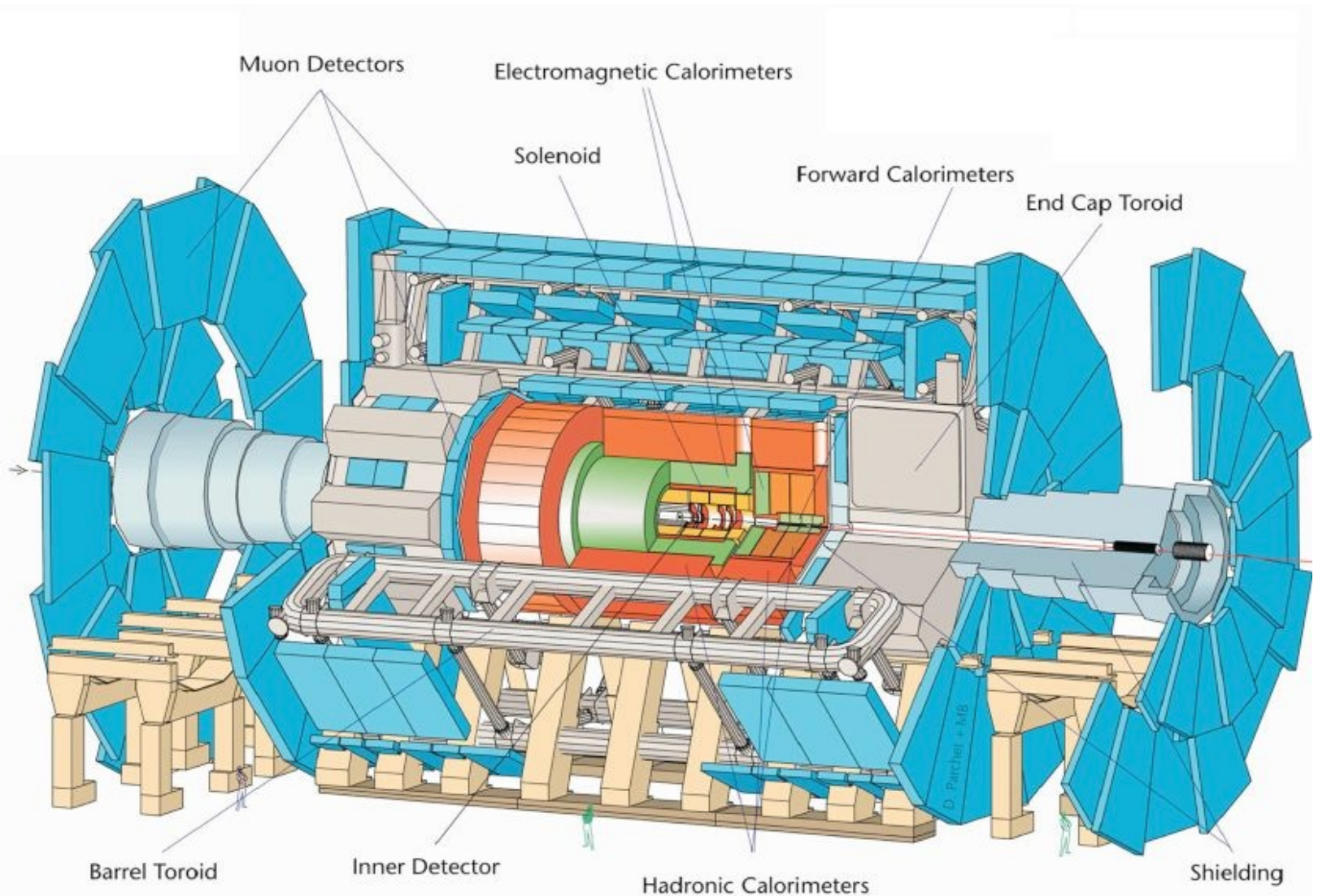
ATLAS

The CMS Detector

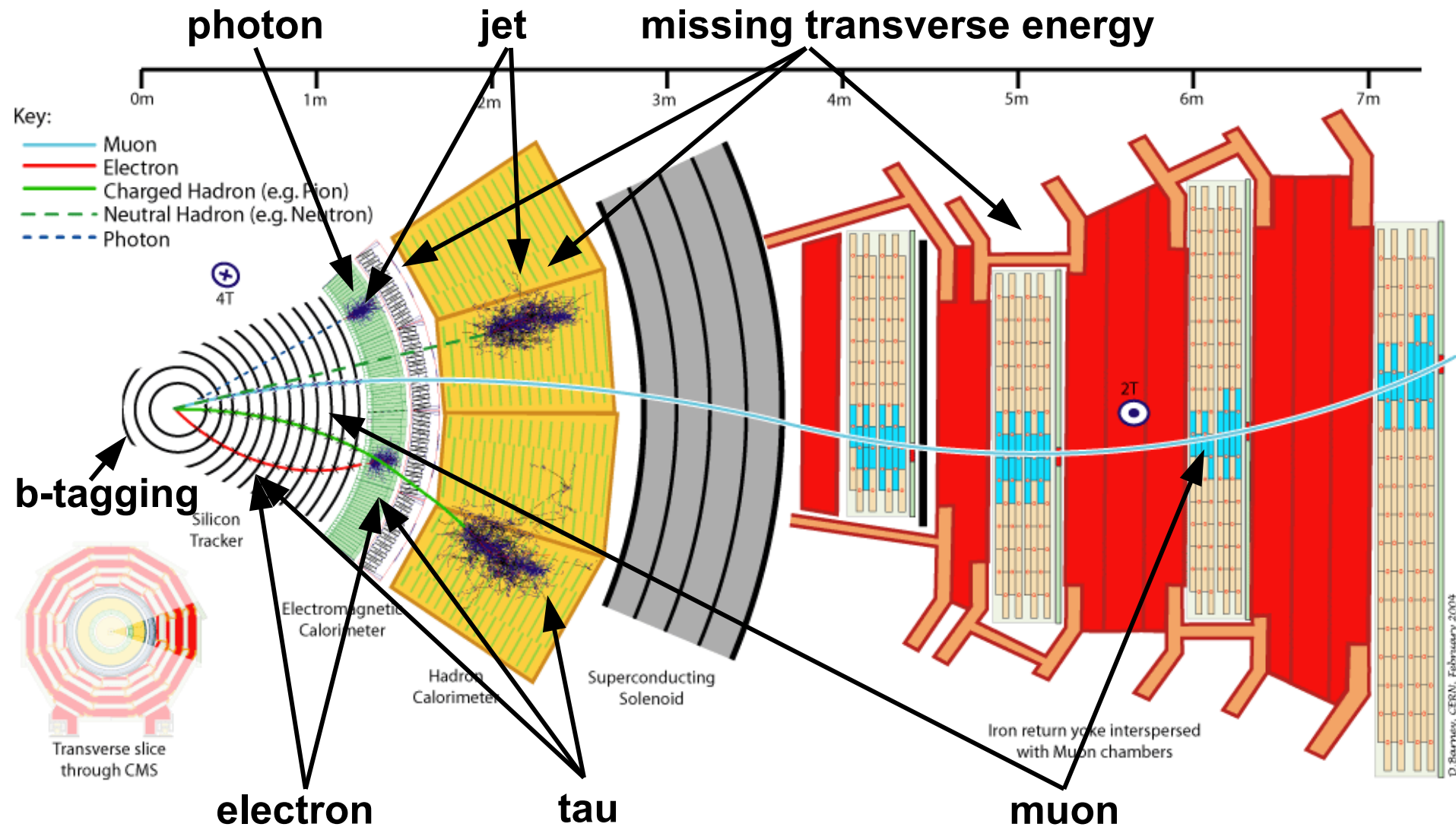
CMS



The ATLAS Detector



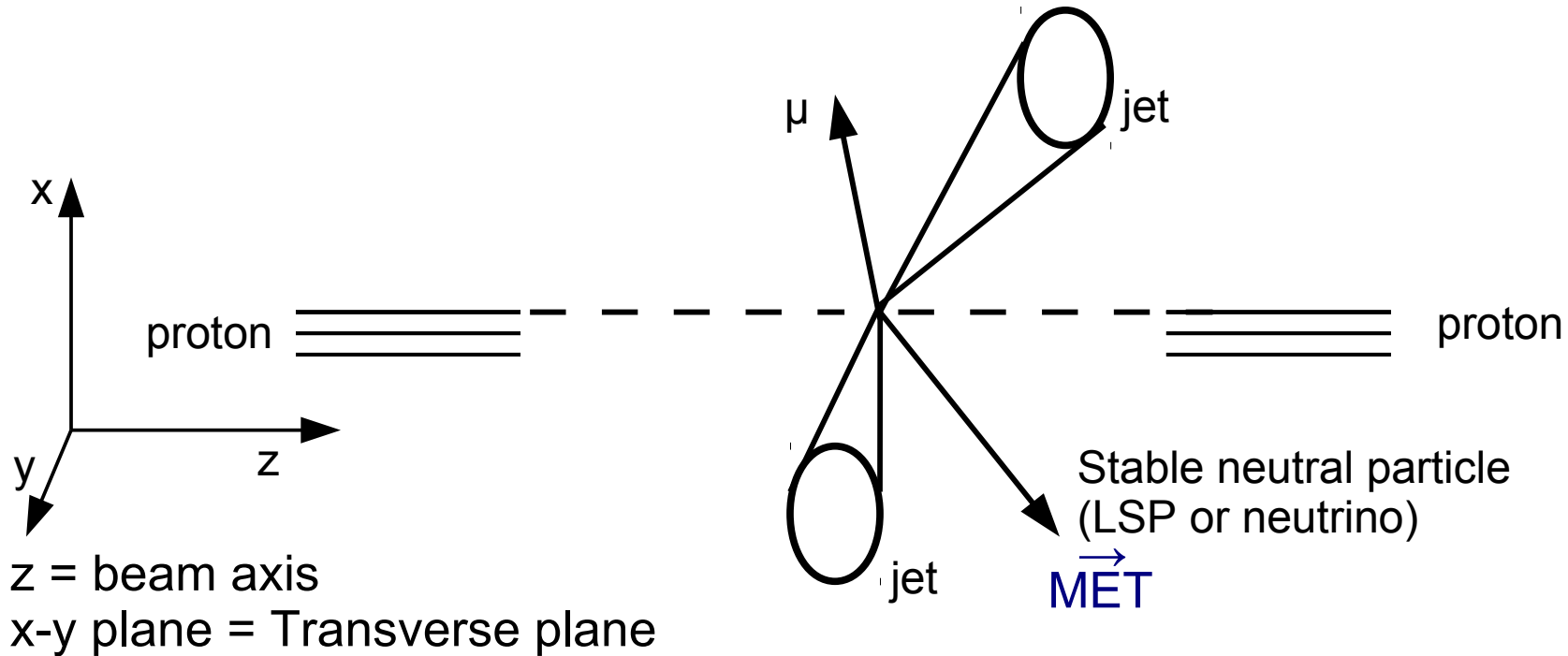
A word about Object Reconstruction



D. Barnier, CERN, February 2004

A word about Object Reconstruction

Missing Transverse Energy = $-\sum \vec{p}_T(\text{calo}) - \sum \vec{p}_T(\text{muon}) = p_T(\text{undetected})$
 (should be called missing transverse momentum)

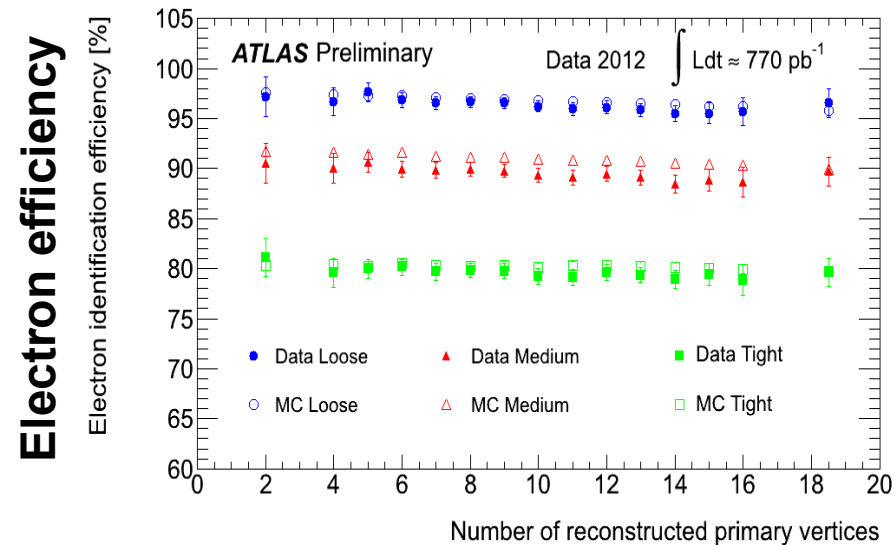
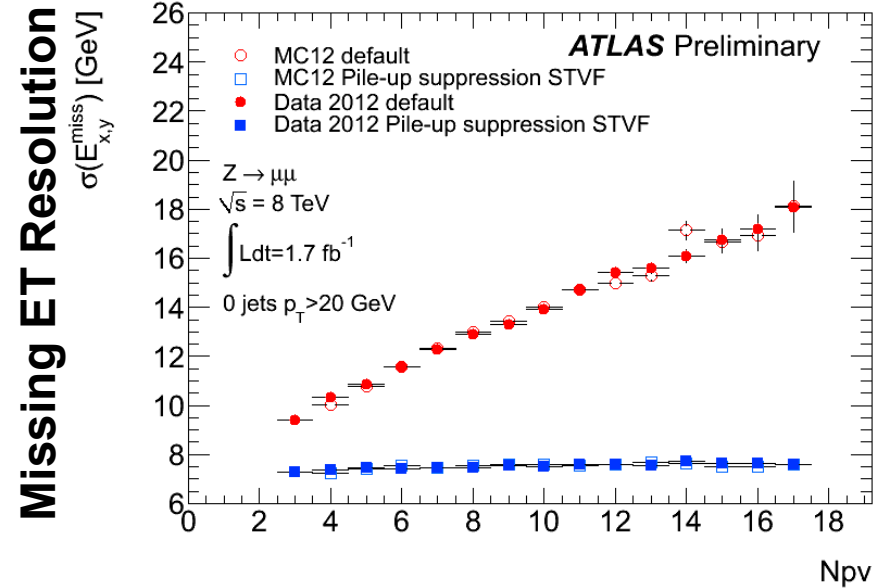
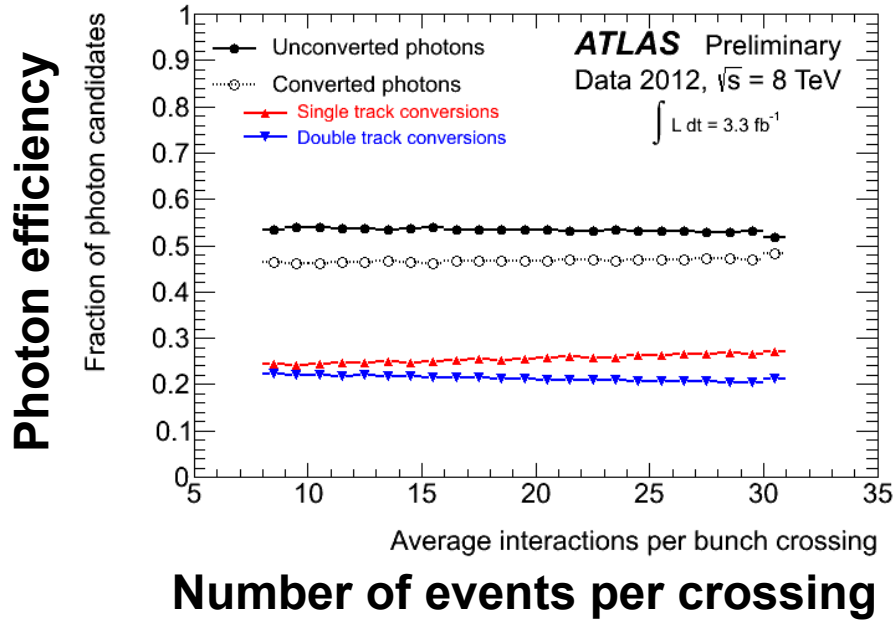


$p_z(\text{proton}) = 4 \text{ TeV}$
 $p_z(\text{parton}) = ? \text{ TeV}$
 $p_T(\text{parton}) = 0 \text{ TeV}$

Momentum conservation:
 $\sum \vec{p}_T(\text{final state}) = \vec{p}_T(\text{partons}) = \vec{0}$

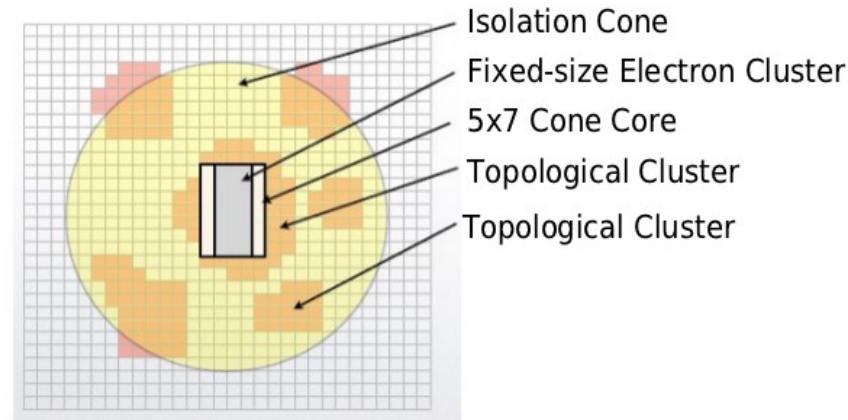
A word about Object Reconstruction

- A lot of work put into coping with pile-up

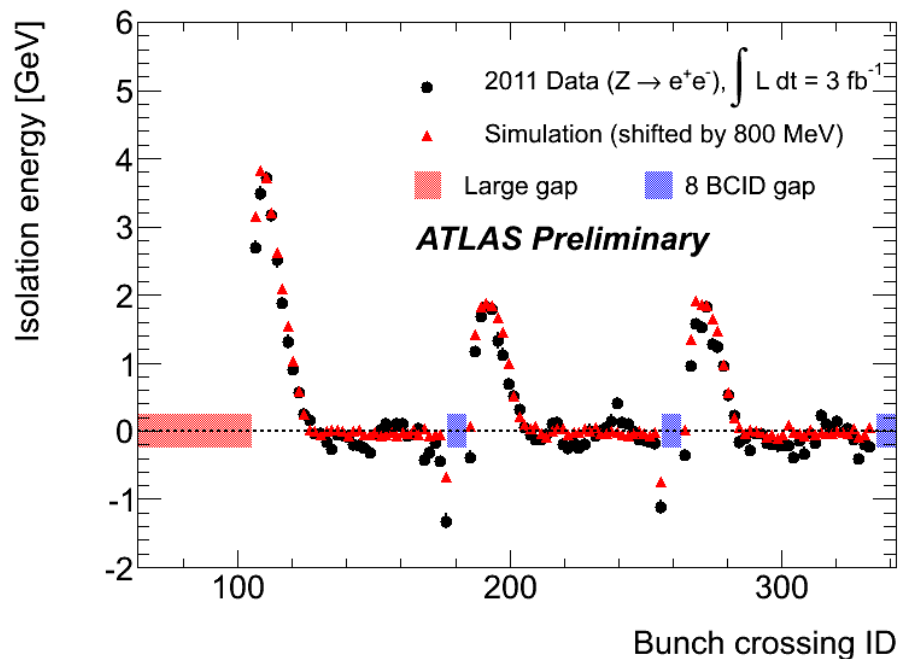


Pile-up and Calorimeter Isolation

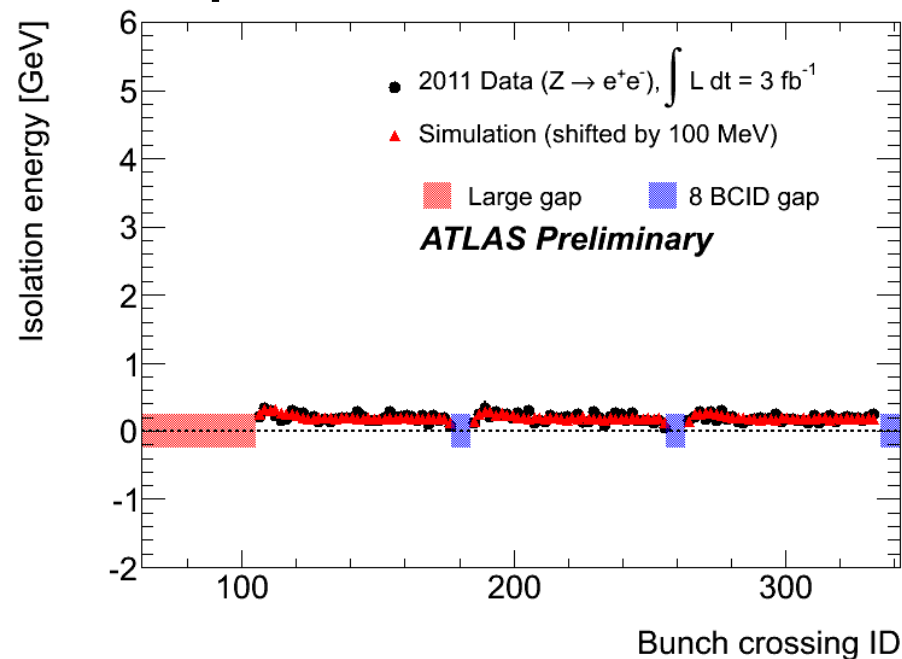
- “Isolation”: amount of energy surrounding a lepton
- Crucial variable to reject fake leptons



Cell-based isolation:

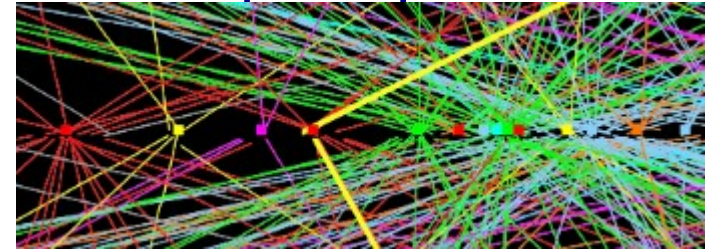


Topo-Cluster isolation:

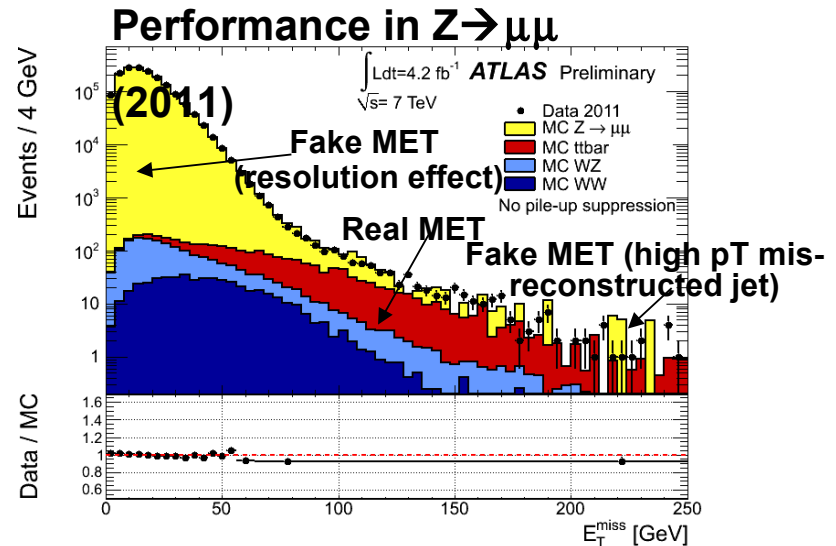
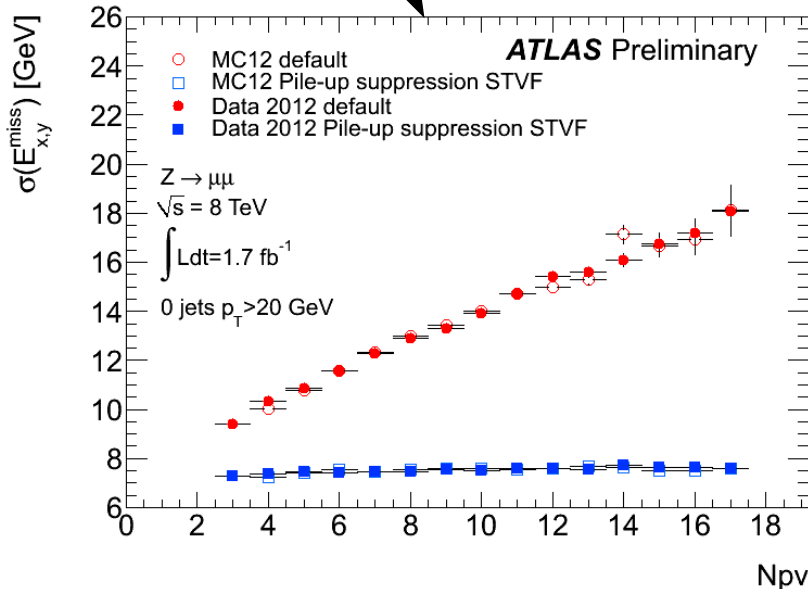


Missing ET and pile-up

- Use fraction of tracks from the Primary Vertex associated to a jet (a.k.a Jet Vertex Fraction) to correct for pile-up:
- Similarly, correct energy not associated with jets (a.k.a. Soft Term Vertex Fraction)

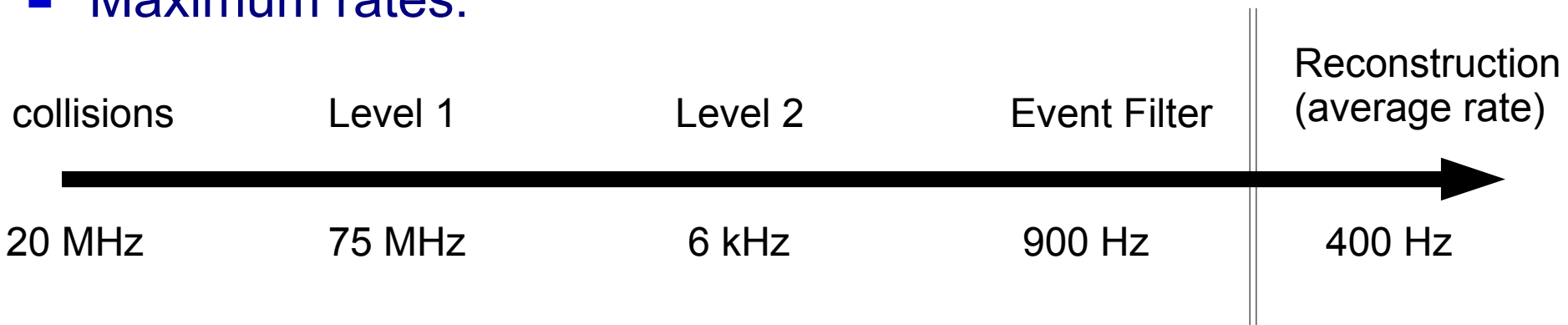
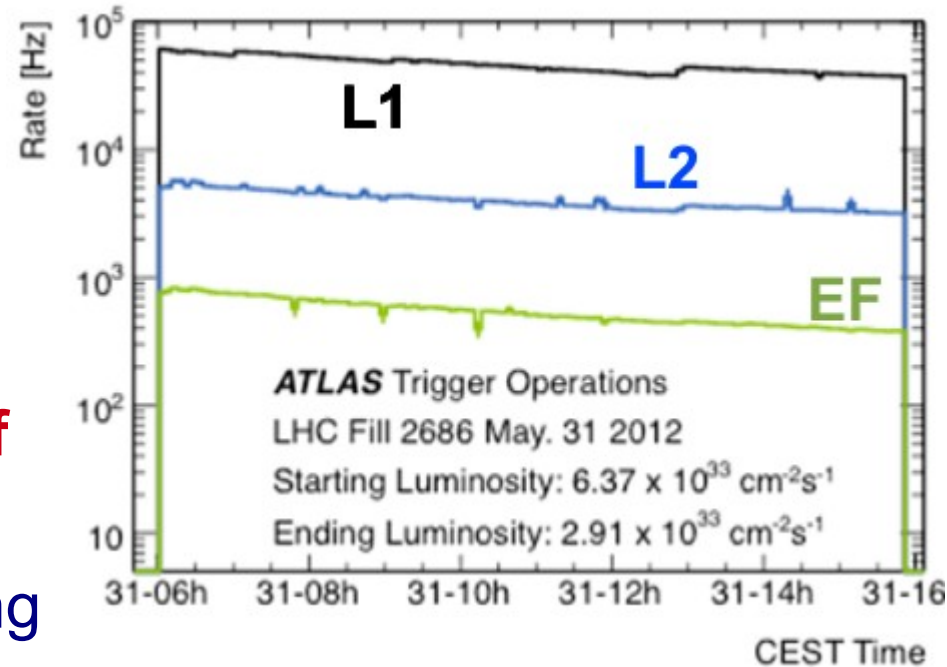


$$\frac{\sum_{\text{track, PV}} p_T}{\sum_{\text{track}} p_T}$$

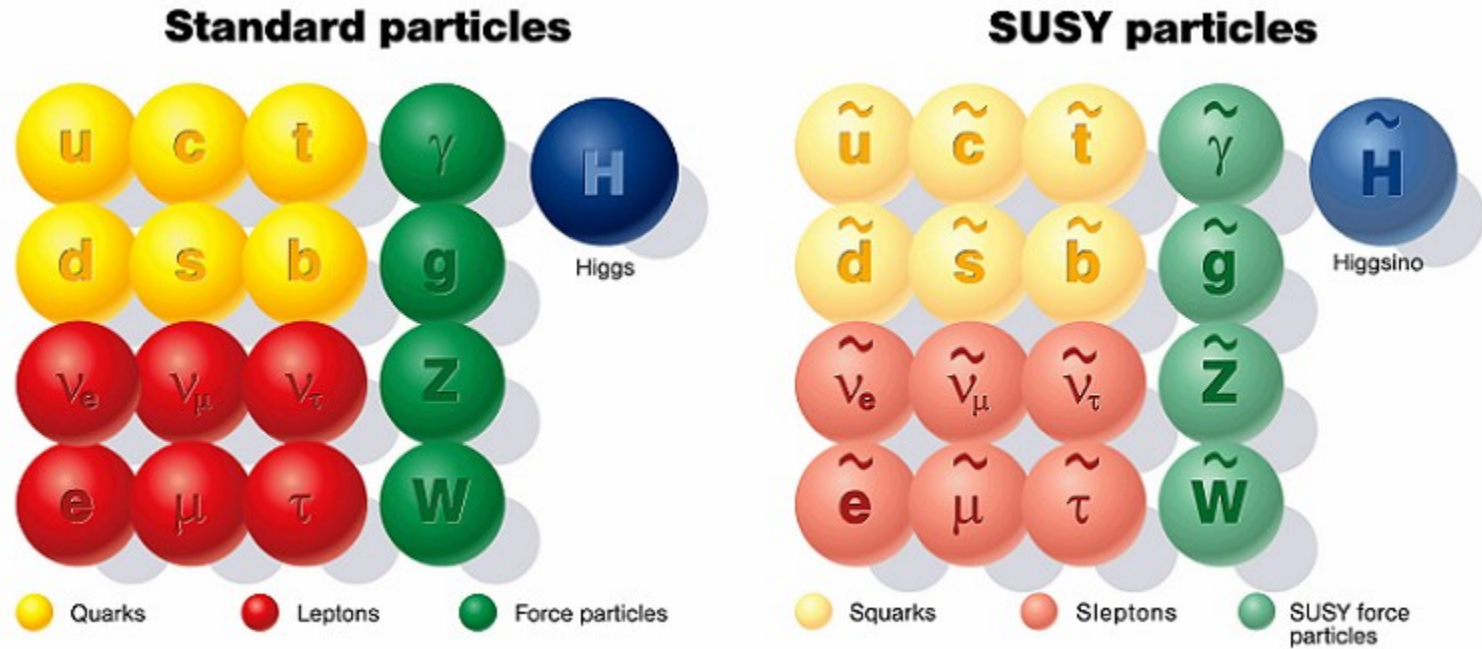


A word about Trigger

- Bunch spacing = 50 ns → collision rate = 20 MHz
- Recording events on disc at 400 Hz only
→ **We record only ~ 1 out of 100000 events!**
- In ATLAS, 3 levels of triggering
- Maximum rates:



Supersymmetry

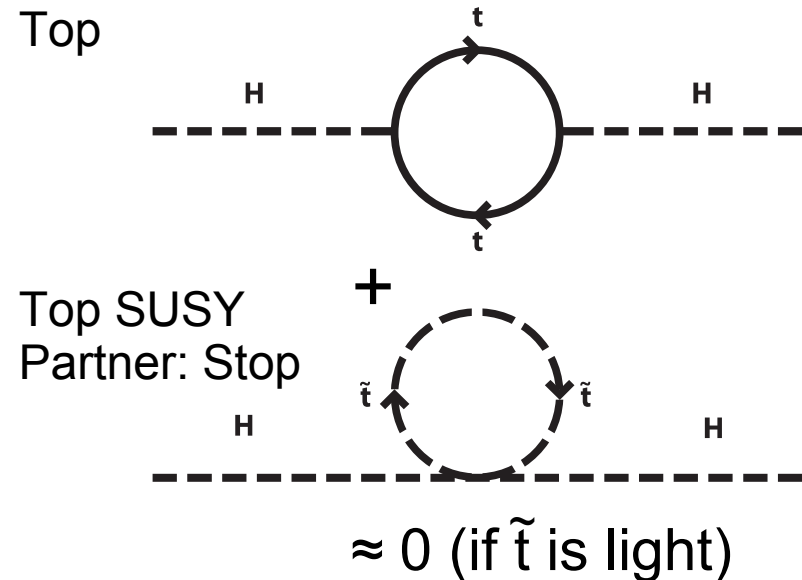


Supersymmetry

- Extension of the Poincaré algebra
- Fermion \leftrightarrow Boson symmetry
- Predicts new spectrum of supersymmetric particles

The pro's:

- Theoretically pleasing
 - Ingredient to string theory
- Solves many problems of the SM. Ex: stabilizes Higgs sector
- Excellent Dark Matter candidate



Supersymmetry

- Extension of the Poincaré algebra
- Fermion \leftrightarrow Boson symmetry
- Predicts new spectrum of supersymmetric particles

The con's:

- It has not been seen yet:

→ Must be seriously broken: $m(\text{selectron}) \gg m(\text{electron})$

→ To avoid lepton/baryon # violation, must add R-parity: $R = (-1)^{3(B-L) + 2s}$

Side effect: **Lightest SUSY Particle (LSP) is stable** → **Excellent Dark Matter candidate**

- Many free parameters

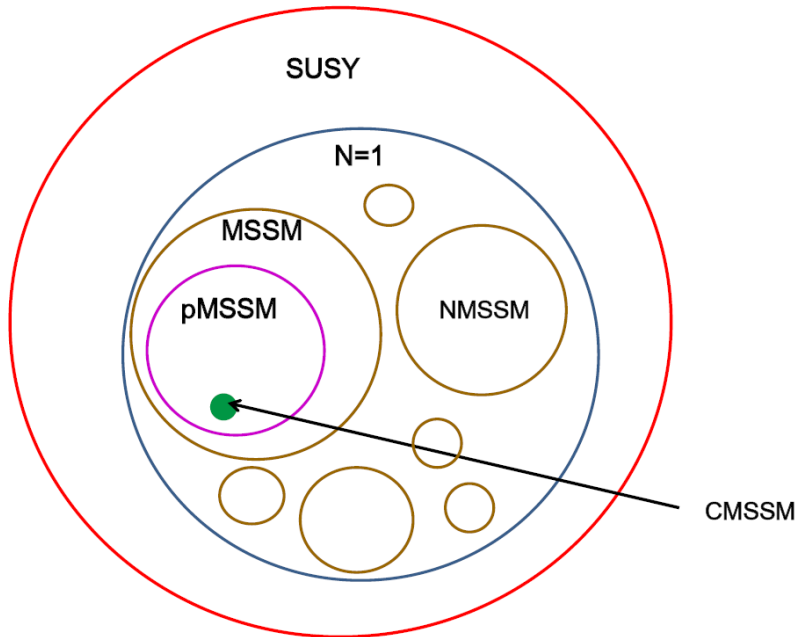


spin

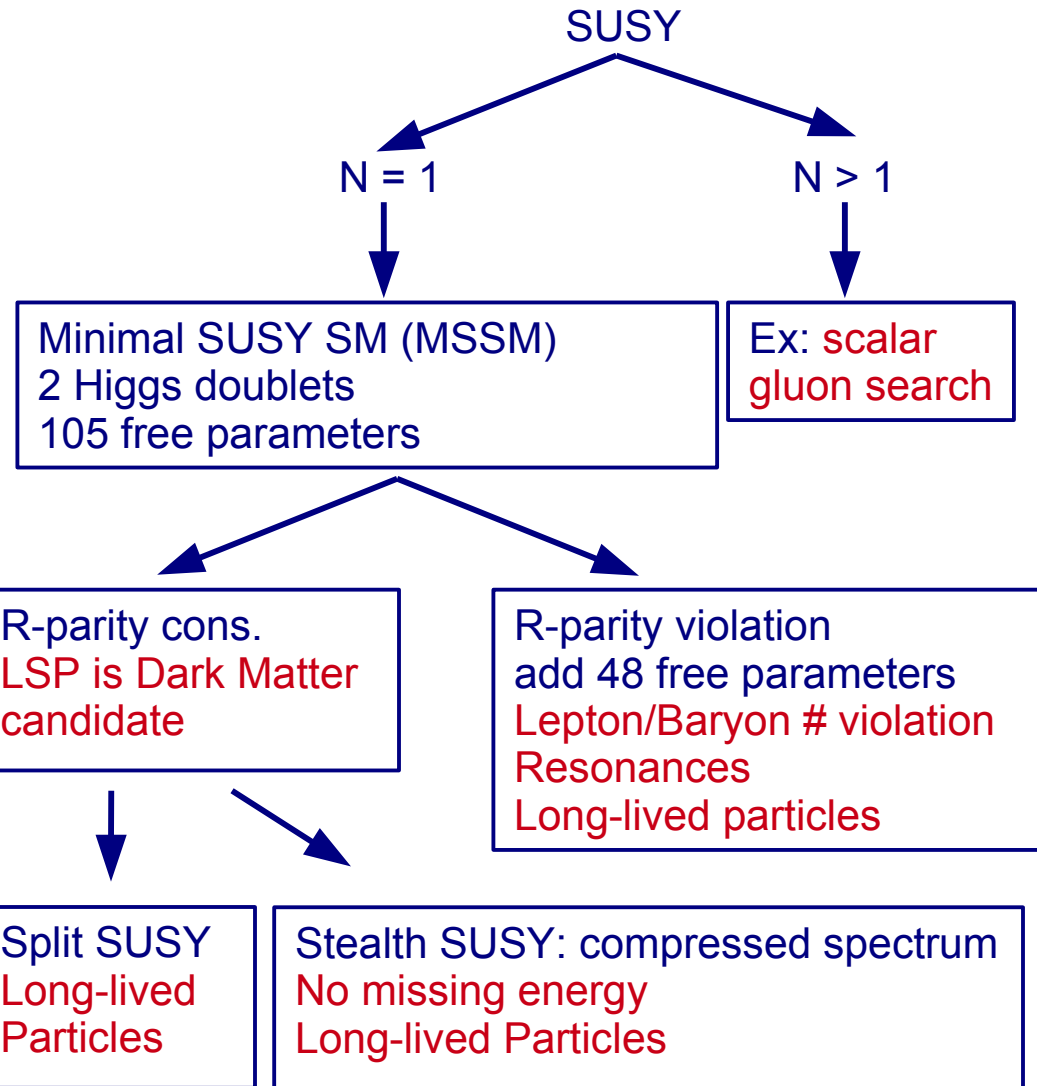


Which Supersymmetry?

SUSY Phenomenology is Very Diverse



T. Rizzo (SLAC Summer Institute, 01-Aug-12)



The MSSM and the Constrained MSSM

- MSSM: N=1 and 2 Higgs Doublets. 105 free parameters
- CMSSM / mSUGRA: down to 5 parameters
 - Strong assumptions, esp. on universality of the 3 generations

MSUGRA parameters:

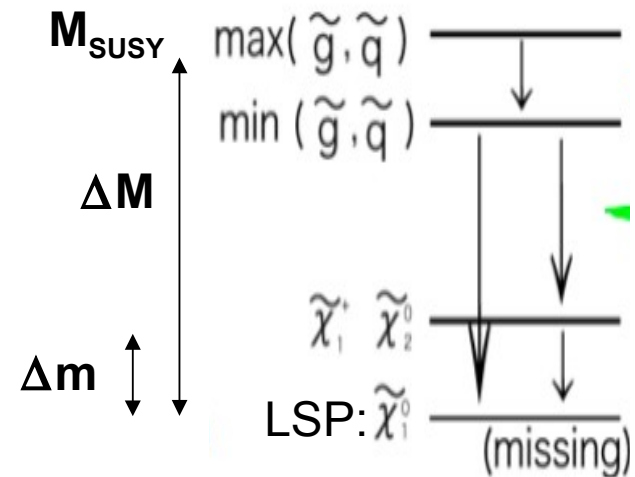
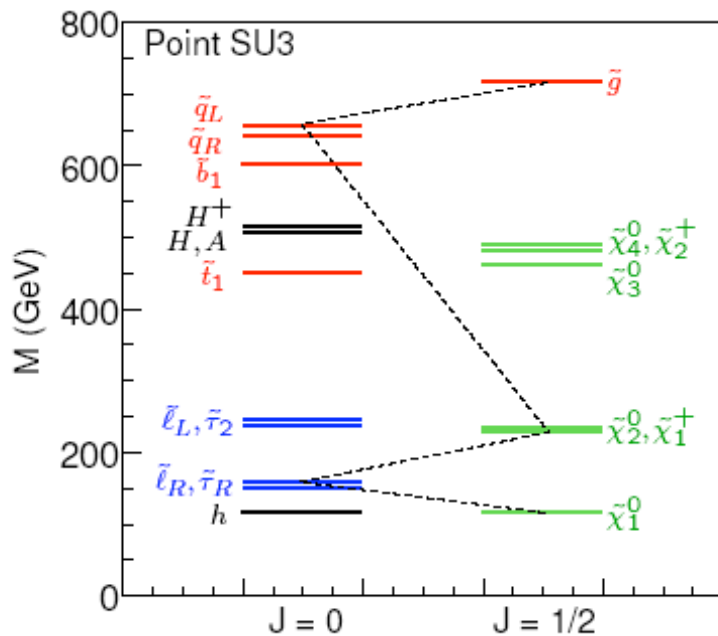
- m_0 : mass of scalars (squarks and sleptons) at GUT scale
- $m_{1/2}$: mass of fermions (gluinos and gauginos) at GUT scale
- A_0 : trilinear couplings
- $\tan \beta = \langle H_u \rangle / \langle H_d \rangle$
- Sign of μ

MSSM: 29 sparticles + 5 Higgs

Names	Spin	P_R	Gauge Eigenstates	Mass Eigenstates
Higgs bosons	0	+1	$H_u^0 H_d^0 H_u^+ H_d^-$	$h^0 H^0 A^0 H^\pm$
squarks	0	-1	$\tilde{u}_L \tilde{u}_R \tilde{d}_L \tilde{d}_R$	(same)
			$\tilde{s}_L \tilde{s}_R \tilde{c}_L \tilde{c}_R$	(same)
			$\tilde{t}_L \tilde{t}_R \tilde{b}_L \tilde{b}_R$	$\tilde{t}_1 \tilde{t}_2 \tilde{b}_1 \tilde{b}_2$
sleptons	0	-1	$\tilde{e}_L \tilde{e}_R \tilde{\nu}_e$	(same)
			$\tilde{\mu}_L \tilde{\mu}_R \tilde{\nu}_\mu$ (Bino) (Wino) (Higgsino)	(same)
			$\tilde{\tau}_L \tilde{\tau}_R \tilde{\nu}_\tau$ (Wino) (Higgsino)	$\tilde{\tau}_1 \tilde{\tau}_2 \tilde{\nu}_\tau$
neutralinos	1/2	-1	$\tilde{B}^0 \tilde{W}^0 \tilde{H}_u^0 \tilde{H}_d^0$	$\tilde{N}_1 \tilde{N}_2 \tilde{N}_3 \tilde{N}_4$
charginos	1/2	-1	$\tilde{W}^\pm \tilde{H}_u^\pm \tilde{H}_d^\mp$	$\tilde{C}_1^\pm \tilde{C}_2^\pm$
gluino	1/2	-1	\tilde{g}	(same)
goldstino (gravitino)	1/2 (3/2)	-1	\tilde{G}	(same)

Supersymmetry: Cascades at the LHC

- Cascade depends on SUSY mass spectrum
- Missing ET caused by LSP escaping detector
 - Larger Missing ET if ΔM is large
 - “compressed” SUSY: small ΔM and \sim no missing ET



Missing ET $\sim \Delta M$
 Total event energy $\sim 2M_{\text{SUSY}}$

Looking for squarks and gluinos

Jets + Missing E_T (a.k.a. “0-lepton”)

$$\tilde{q} \rightarrow q\tilde{\chi}_1^0$$
$$\tilde{g} \rightarrow qq\tilde{\chi}_1^0$$

- “Workhorse” analysis of SUSY+MET searches
- Several analyses developed by ATLAS and CMS. Use the ATLAS analysis as an example:
- Select events with **2 to 6 jets**
- **Veto leptons** and events with > 6 jets (left to dedicated high-multiplicity analysis)
- **Trigger: Jet $p_T > 75$ GeV and Missing $E_T > 55$ GeV**
 $\varepsilon > 98\%$ above turn-on
- **Discriminant variables:**
 - **H_T** = sum of jet p_T (including jets with $p_T > 40$ GeV and $|\eta| < 2.8$)
 - **m_{eff}** = H_T + Missing E_T
- **Optimize cut on m_{eff} and Missing ET for each jet multiplicity**

SUSY: Jets + Missing E_T

$$\tilde{q} \rightarrow q\tilde{\chi}_1^0$$

$$\tilde{g} \rightarrow qq\tilde{\chi}_1^0$$

ATLAS 7 TeV analysis selection: 11 signal regions
(re-optimized a bit for 8 TeV)

Cut-and-count in each signal region

above trigger
turn-on

Jet p_T

ETmiss
and
 m_{eff}

Requirement	Channel					
	A	A'	B	C	D	E
$E_T^{\text{miss}} [\text{GeV}] >$	160					
$p_T(j_1) [\text{GeV}] >$	130					
$p_T(j_2) [\text{GeV}] >$	60					
$p_T(j_3) [\text{GeV}] >$	-	-	60	60	60	60
$p_T(j_4) [\text{GeV}] >$	-	-	-	60	60	60
$p_T(j_5) [\text{GeV}] >$	-	-	-	-	40	40
$p_T(j_6) [\text{GeV}] >$	-	-	-	-	-	40
$\Delta\phi(\text{jet}, E_T^{\text{miss}})_{\text{min}} >$	0.4 ($i = \{1, 2, (3)\}$)			0.4 ($i = \{1, 2, 3\}$), 0.2 ($p_T > 40$ GeV jets)		
$E_T^{\text{miss}} / m_{\text{eff}}(Nj) >$	0.3 (2j)	0.4 (2j)	0.25 (3j)	0.25 (4j)	0.2 (5j)	0.15 (6j)
$m_{\text{eff}}(\text{incl.}) [\text{GeV}] >$	1900/1400/-	-/1200/-	1900/-/-	1500/1200/900	1500/-/-	1400/1200/900

2 jets

3

4

5

6

SUSY: Jets + Missing E_T

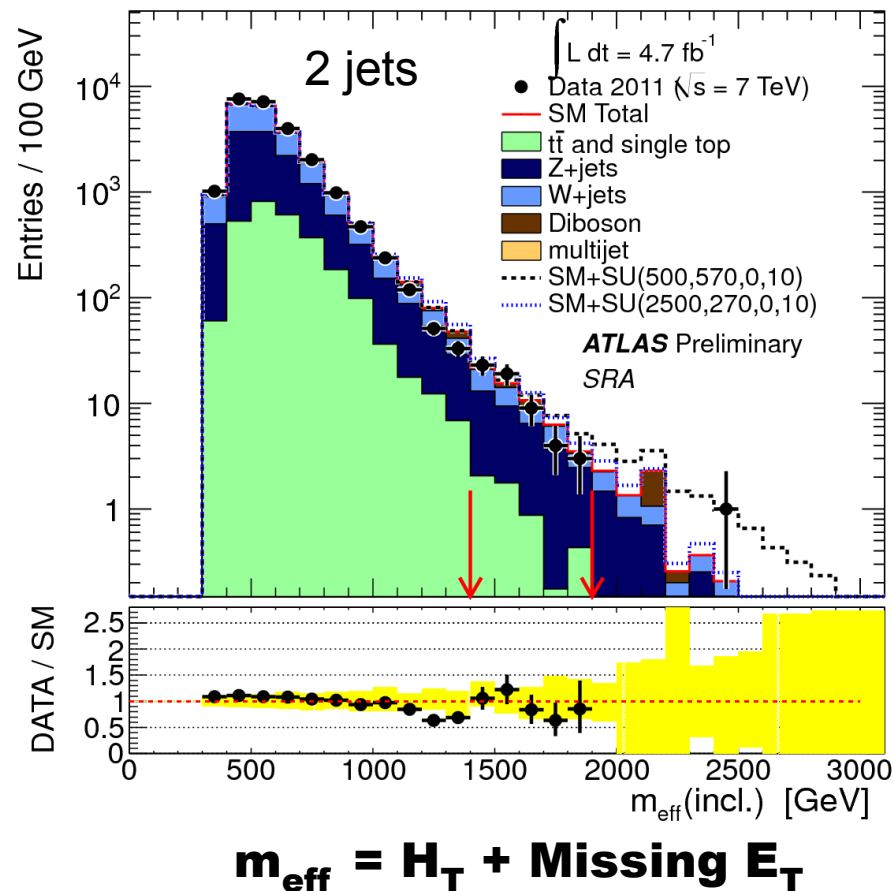
$$\tilde{q} \rightarrow q\tilde{\chi}_1^0$$

$$\tilde{g} \rightarrow qq\tilde{\chi}_1^0$$

- Low jet-multiplicity:
sensitive to squark production
- High jet-multiplicity:
sensitive to gluino production

$$\tilde{q} \rightarrow q\tilde{\chi}_1^0$$

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SUSY: Jets + Missing E_T

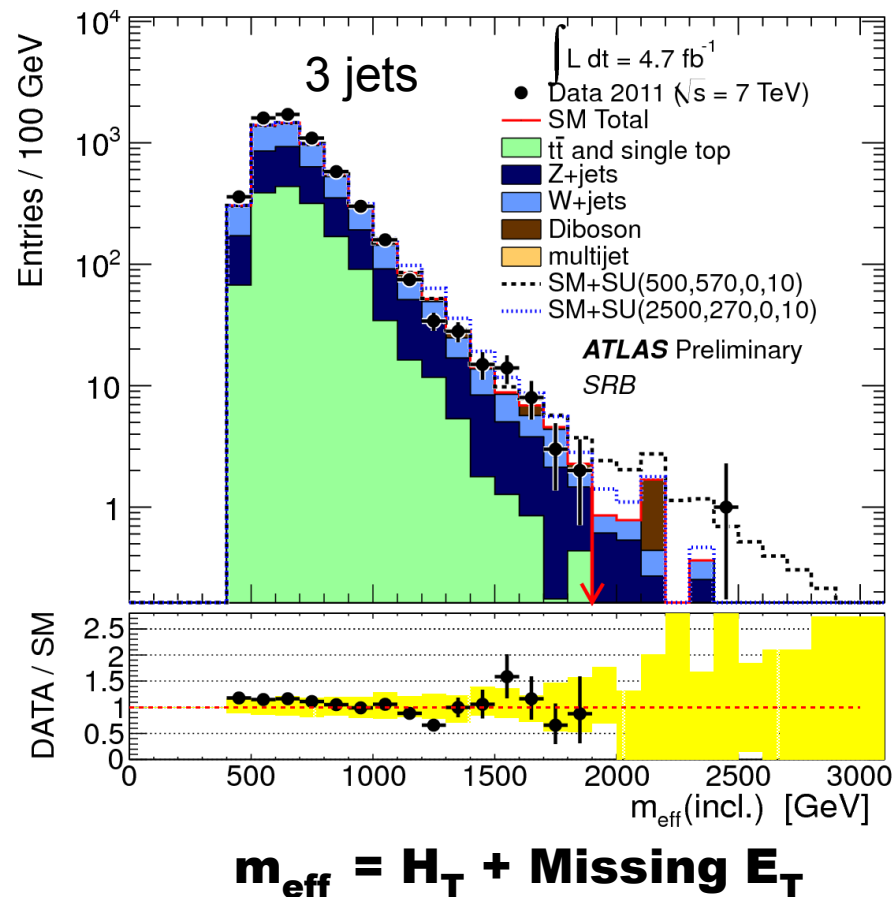
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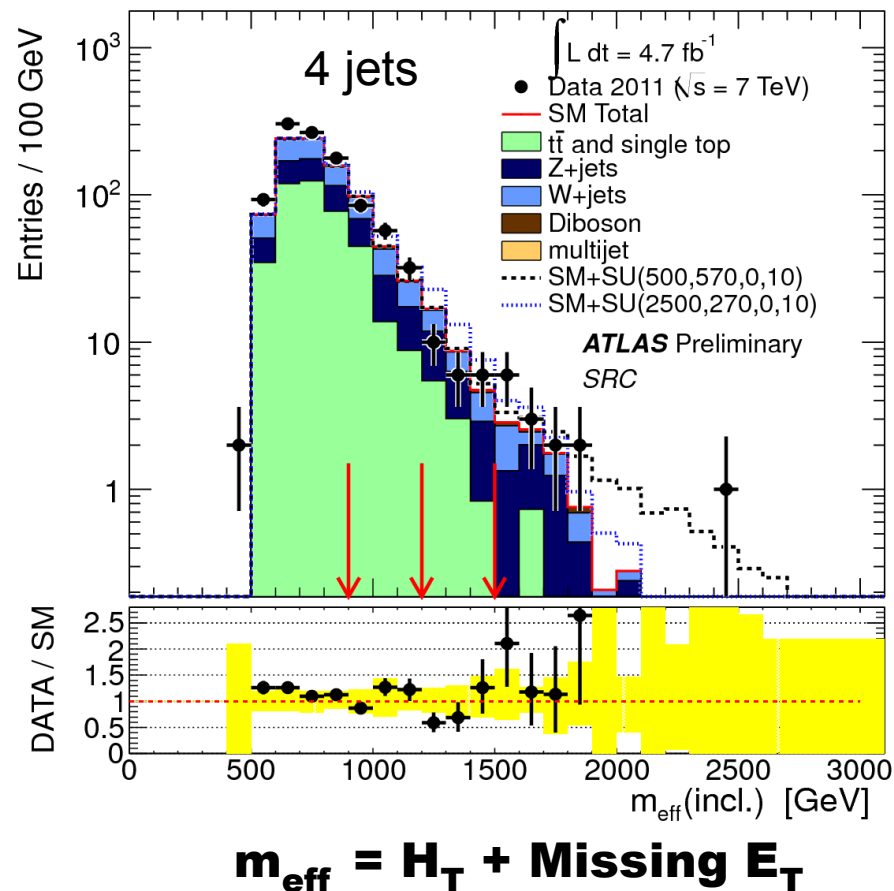
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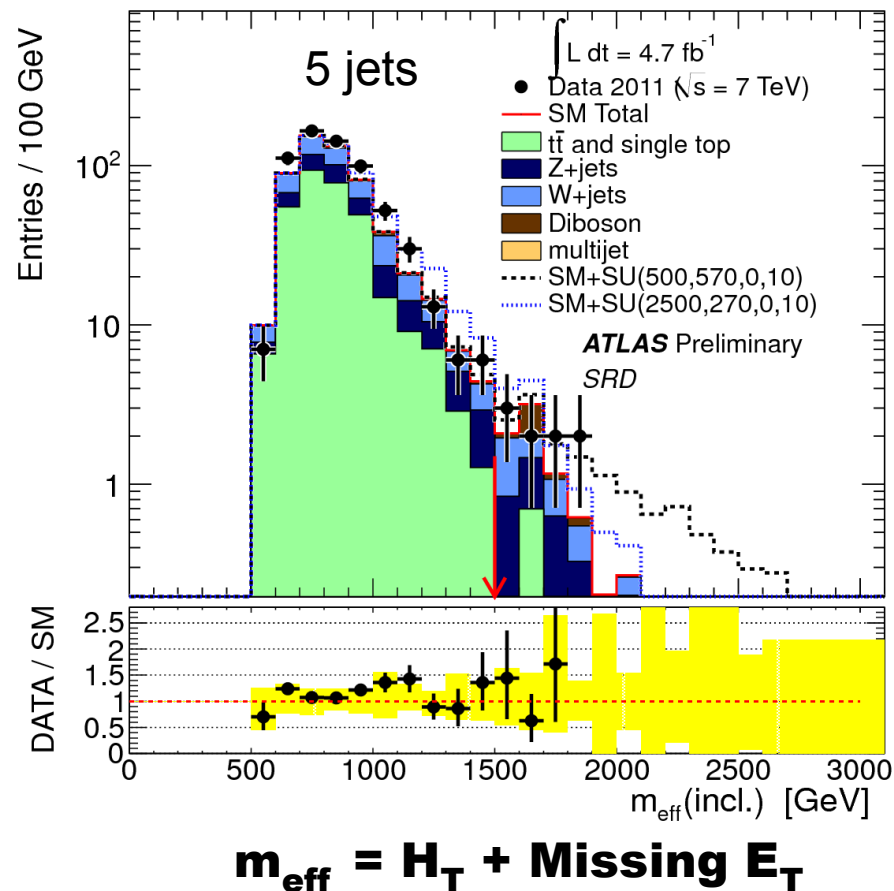
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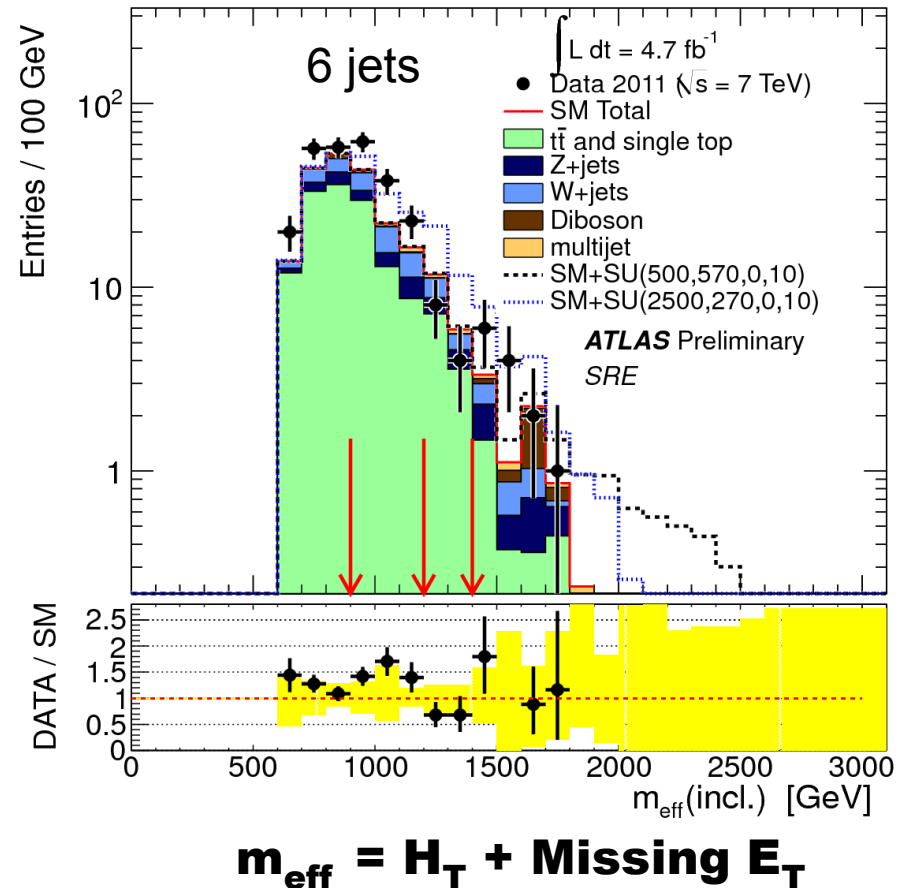
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$$\tilde{g} \rightarrow qq\tilde{\chi}_1^0$$



SUSY: Jets + Missing E_T

$$\tilde{q} \rightarrow q\tilde{\chi}_1^0$$

$$\tilde{g} \rightarrow qq\tilde{\chi}_1^0$$

- Background estimation is (mostly) data-driven
- Define control regions with enhanced background:

CR	SR background	CR process	CR selection
CRY	$Z(\rightarrow \nu\nu)+\text{jets}$	$\gamma+\text{jets}$	Isolated photon
CRQ	QCD jets	QCD jets	Reversed $\Delta\phi(\text{jet}, \mathbf{E}_T^{\text{miss}})_{\text{min}}$ and $E_T^{\text{miss}}/m_{\text{eff}}(Nj)$ cuts
CRW	$W(\rightarrow \ell\nu)+\text{jets}$	$W(\rightarrow \ell\nu)+\text{jets}$	$30 \text{ GeV} < m_T(\ell, E_T^{\text{miss}}) < 100 \text{ GeV}$, b -veto
CRT	$t\bar{t}$ and single- t	$t\bar{t} \rightarrow bbqq'\ell\nu$	$30 \text{ GeV} < m_T(\ell, E_T^{\text{miss}}) < 100 \text{ GeV}$, b -tag

- Calibrate backgrounds to control regions:

$$N(\text{SR, scaled}) = N(\text{CR, obs}) \times \left[\frac{N(\text{SR, unscaled})}{N(\text{CR, unscaled})} \right]$$

↑
from Monte-Carlo

SUSY: Jets + Missing E_T

$$\tilde{q} \rightarrow q\tilde{\chi}_1^0$$

$$\tilde{g} \rightarrow qq\tilde{\chi}_1^0$$

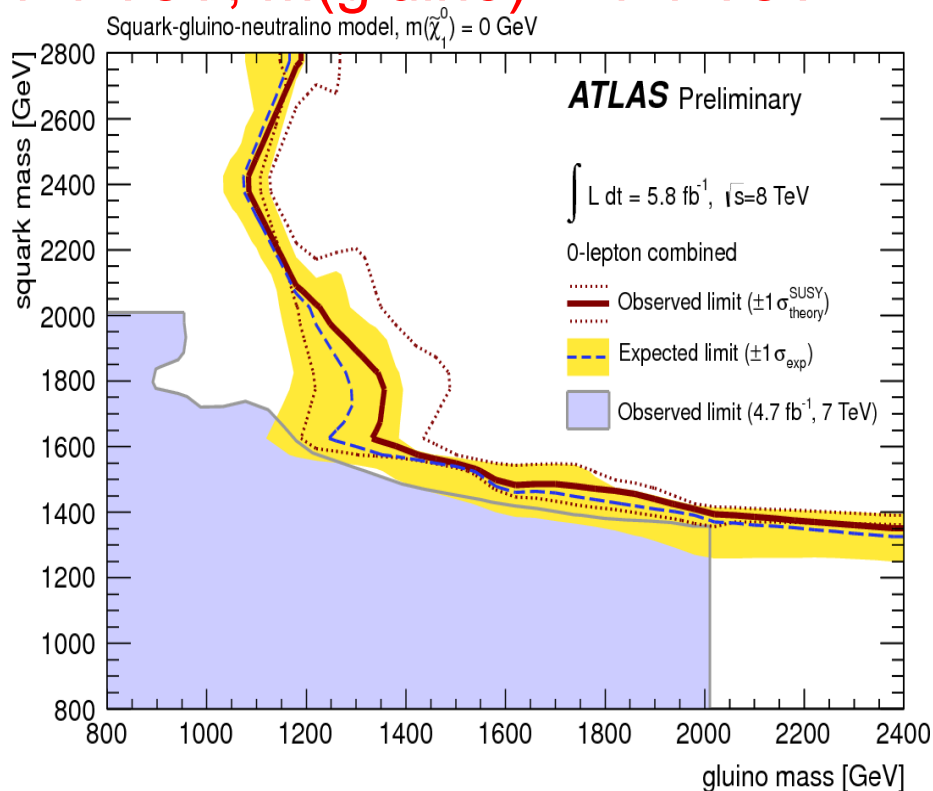
- Background estimation is (mostly) data-driven
- Result for the 8 TeV “tight selection” Signal Regions:

Signal Region	A-tight	B-tight	C-tight	D-tight	E-tight
MC expected events					
Diboson	3.3	0.2	0.0	0.8	2.6
W+jets	6.6	5.6	2.1	3.4	3.3
Z/ γ^* +jets	7.4	4.5	1.9	1.3	1.3
$t\bar{t}$ + single top	1.0	1.1	0.6	1.8	2.7
Fitted background events					
Diboson	3.3 ± 3.1	0.2 ± 1.4	–	0.8 ± 0.4	2.6 ± 2.0
Multi-jets	–	–	–	0.4 ± 0.5	0.1 ± 0.2
W+jets	3 ± 4	2.7 ± 3.4	0.3 ± 0.5	–	0.8 ± 1.3
Z/ γ^* +jets	6.8 ± 2.2	5.1 ± 1.7	2.0 ± 1.1	2.5 ± 1.1	1.2 ± 0.7
$t\bar{t}$ + single top	0.8 ± 0.8	0.8 ± 0.9	0.6 ± 0.5	2.6 ± 1.6	5.1 ± 3.3
Total bkg	14 ± 5	8.7 ± 3.4	2.8 ± 1.2	6.3 ± 2.1	10 ± 4
Observed	10	7	1	5	9
p_0	0.499	0.500	0.499	0.500	0.499
UL on N_{BSM}	8.9	7.3	3.3	6.0	9.3
UL on σ_{BSM} (fb)	1.53	1.26	0.57	1.03	1.60

SUSY: Jets + Missing E_T

$$\tilde{q} \rightarrow q\tilde{\chi}_1^0$$
$$\tilde{g} \rightarrow qq\tilde{\chi}_1^0$$

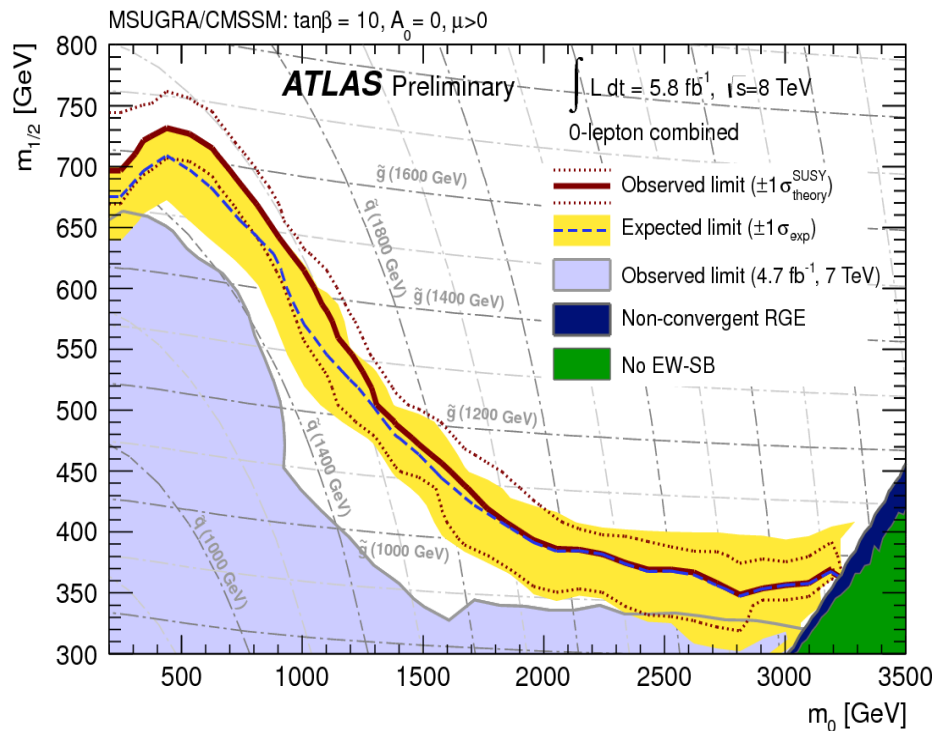
- Limit on simplified model: consider only squark and gluino pair production decaying to squarks, gluinos, quarks, and Lightest SUSY Particle (LSP)
- $m(\text{squark}) > 1.4 \text{ TeV}$, $m(\text{gluino}) > 1.1 \text{ TeV}$



SUSY: Jets + Missing E_T

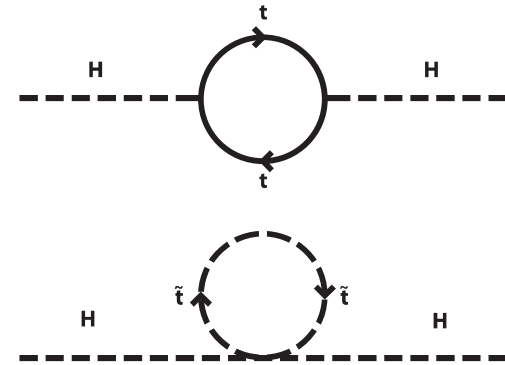
$$\tilde{q} \rightarrow q\tilde{\chi}_1^0$$
$$\tilde{g} \rightarrow qq\tilde{\chi}_1^0$$

- Limits on MSUGRA/CMSSM:
 $m(\text{squark}) = m(\text{gluino}) > 1.5 \text{ TeV}$
- MSUGRA limits in $m_0 - m_{1/2}$ plane:

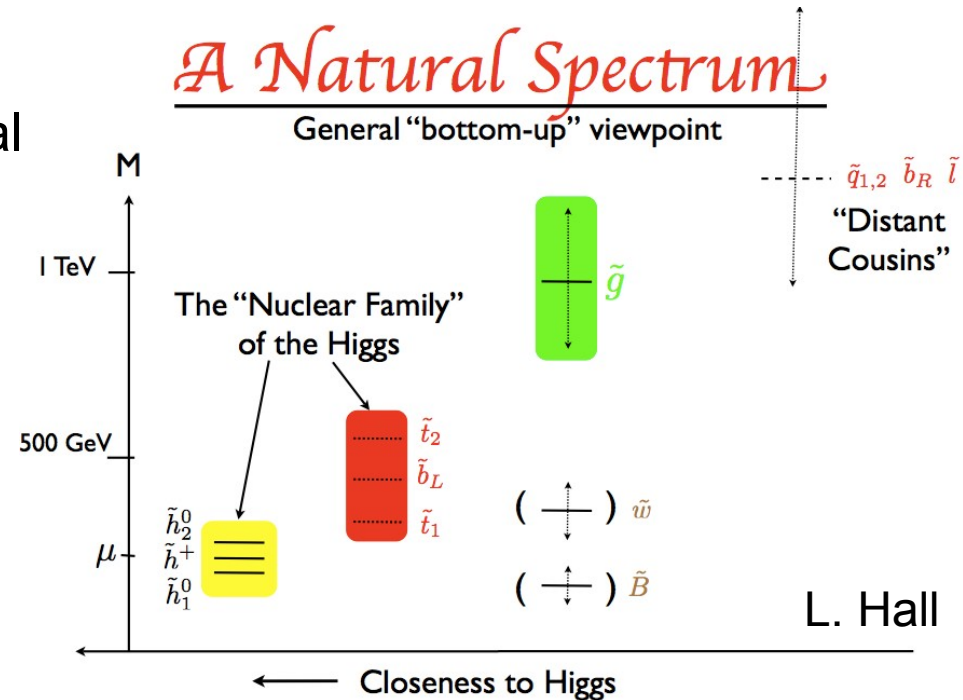


“Natural” SUSY: a lighter 3rd generation?

- SUSY solves Hierarchy Problem only if 3rd generation is “light” i.e. $m(\text{stop}) \sim m(\text{top})$
 - Other gen. can be heavier
- What if 3rd generation lighter than others?
 - Remove constraint of universality of mSUGRA/CMSSM
- Look specifically for stop and sbottom:
 - Direct production
 - Through gluino decays



A Natural Spectrum



“Natural” SUSY: a lighter 3rd generation?

- SUSY solves Hierarchy Problem only if 3rd generation is “light” i.e. $m(\text{stop}) \sim m(\text{top})$

→ Other gen. can be heavier

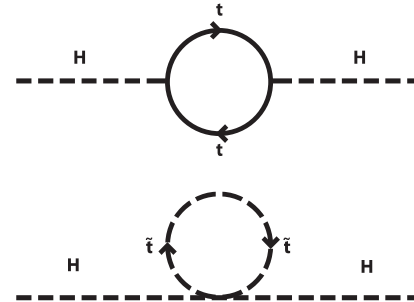
- What if 3rd generation lighter than others?

→ Remove constraint of universality of mSUGRA/CMSSM

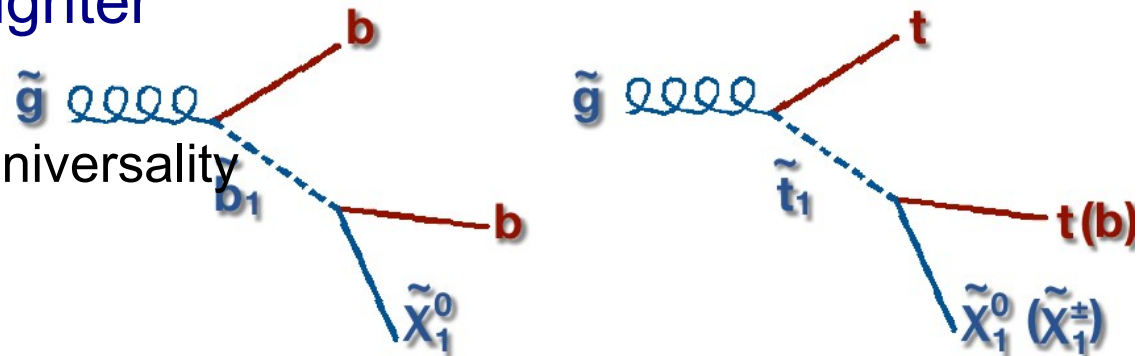
- Look specifically for stop and sbottom:

→ Direct production

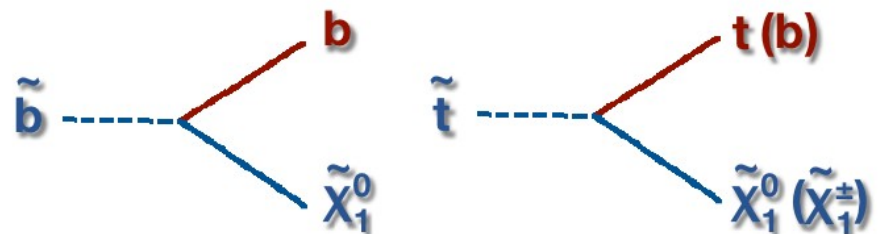
→ Through gluino decays



through gluino decays



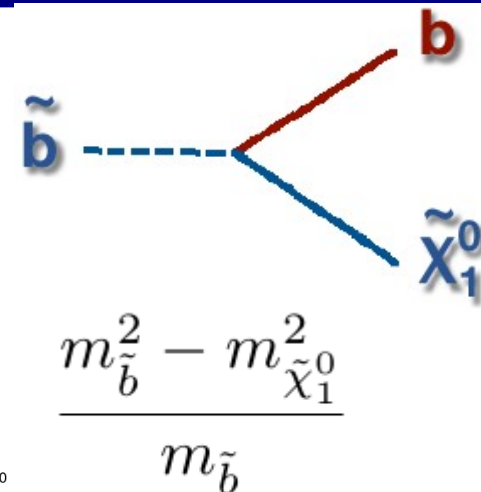
direct production



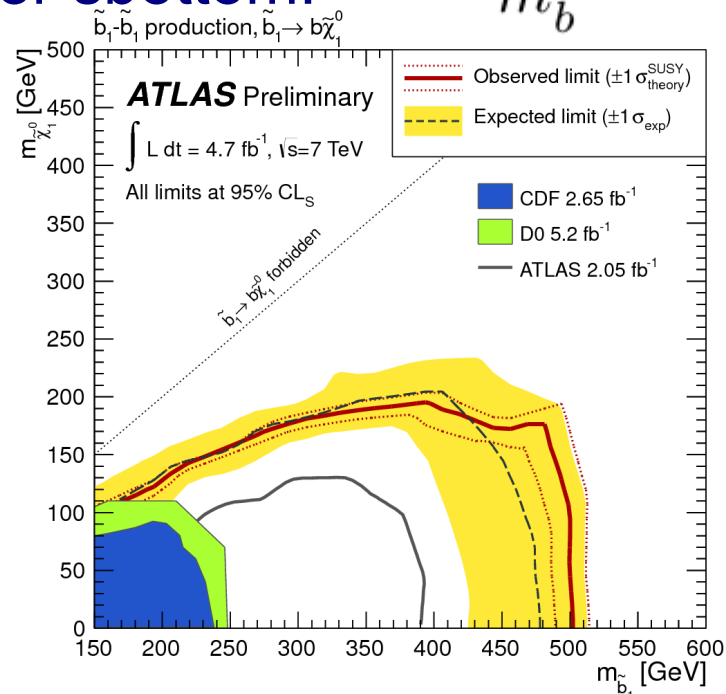
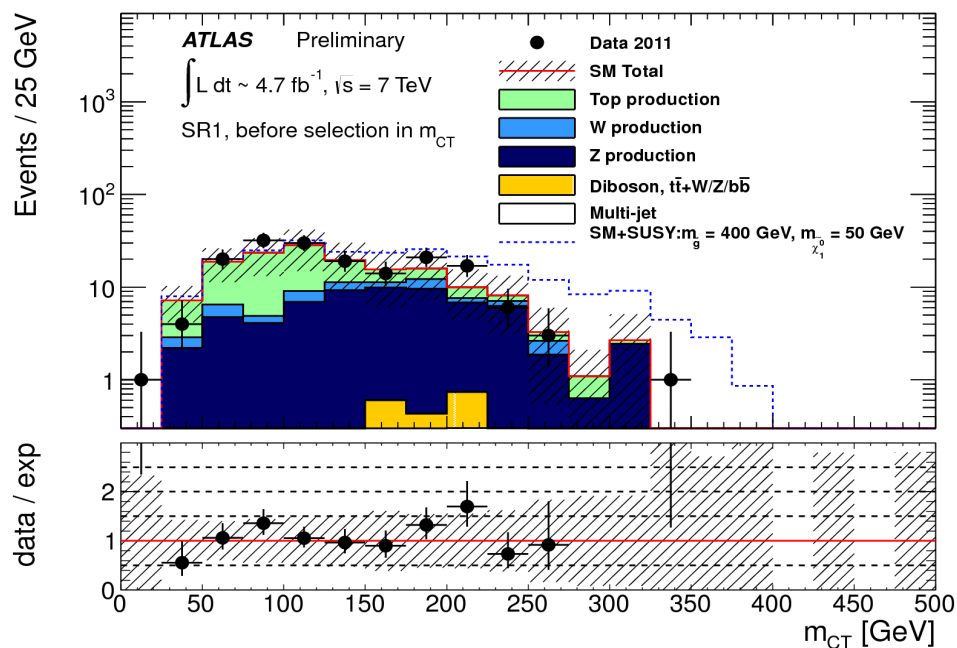
“Natural” SUSY: a lighter 3rd generation? Direct Sbottom Production

- 2 b-jets + Missing ET
- Use “contransverse mass”:

$$m_{CT} = \sqrt{[E_T(b_1) + E_T(b_2)]^2 - [p_T(b_1) - p_T(b_2)]^2}$$



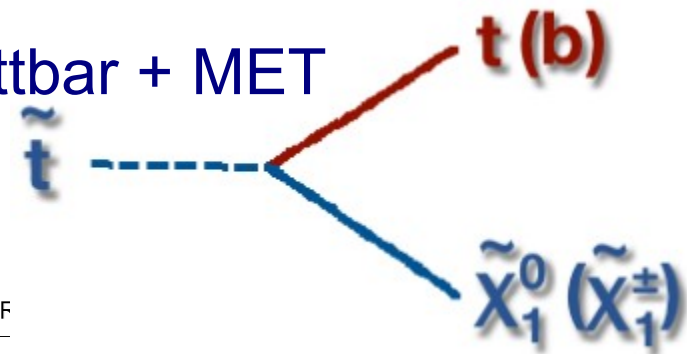
- Endpoint at: for $t\bar{t}$ ~ 135 GeV, for sbottom:



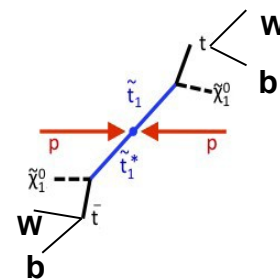
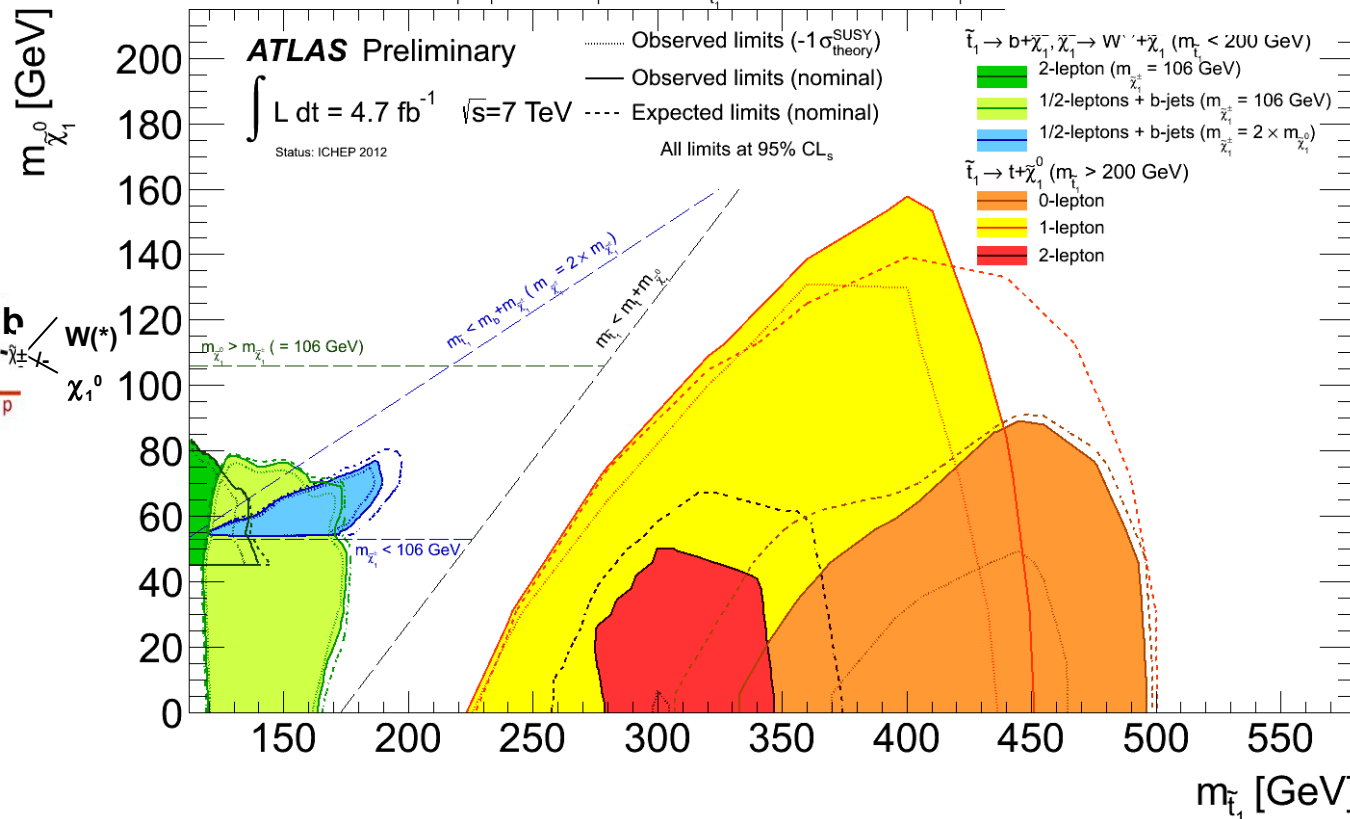
“Natural” SUSY: a lighter 3rd generation? Direct Stop Production

- If $m(\text{stop}) > m(\text{top}) + m(\text{LSP})$: look for $t\bar{t} + \text{MET}$
- If stop is very light, look for decays

involving charginos \rightarrow b's and leptons

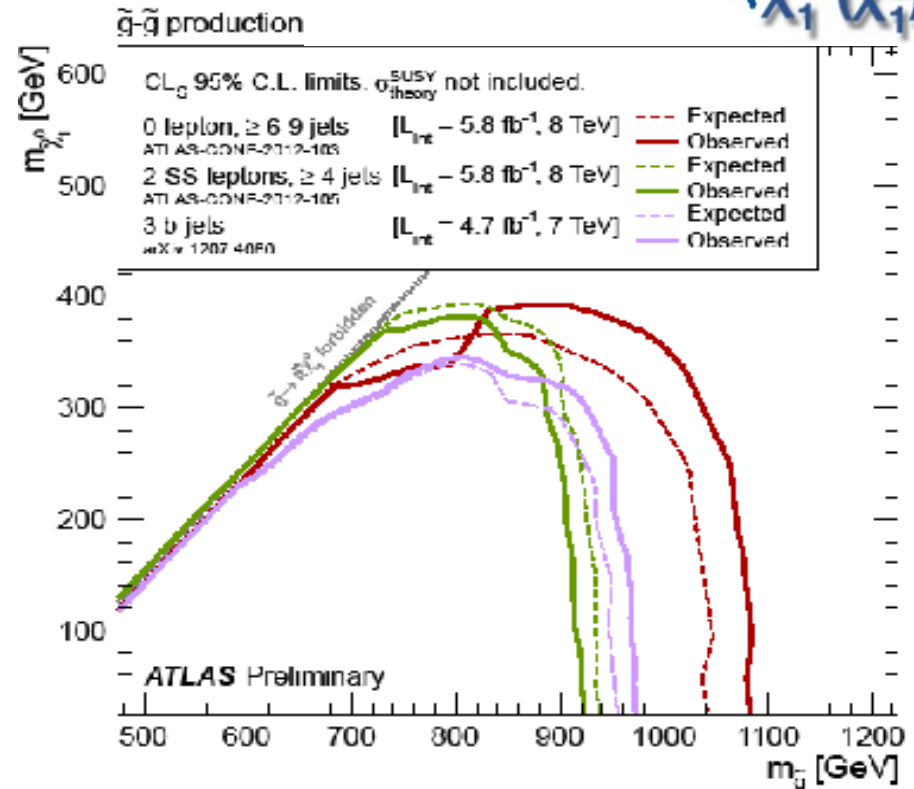
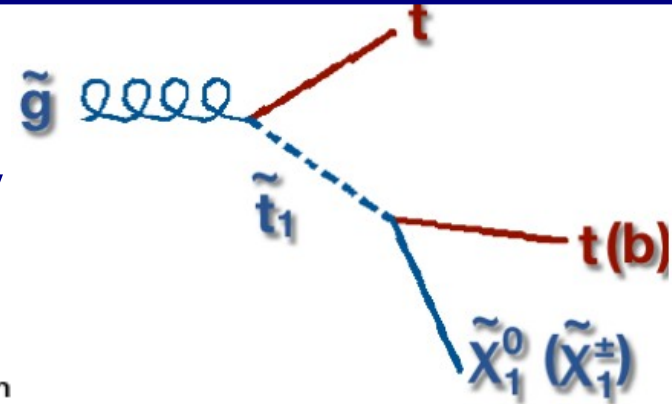


$\tilde{t}_1\tilde{t}_1$ production: $\tilde{t}_1 \rightarrow b + \tilde{\chi}_1^\pm, \tilde{\chi}_1^\pm \rightarrow W^{(\prime)} + \tilde{\chi}_1^0$ (BR=1, $m_{\tilde{t}_1} < 200$ GeV); $\tilde{t}_1 \rightarrow t + \tilde{\chi}_1^0$ (BF)



“Natural” SUSY: a lighter 3rd generation? Stop in gluino decays

- Look for 4 top quarks and missing ET
- Main background is $t\bar{t}$ → reject it by requiring either of the following:
 - 3 b-jets
 - 2 same-sign leptons
 - Large jet multiplicity
- Sensitive to
 - $m(\text{gluino}) < 1000 \text{ GeV}$
 - for $m(\text{LSP}) < 380 \text{ GeV}$



Supersymmetry: Summary

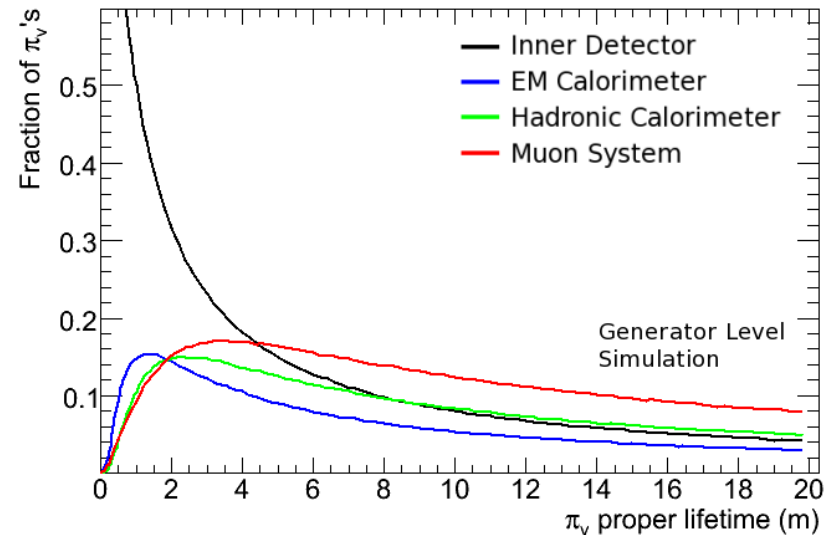
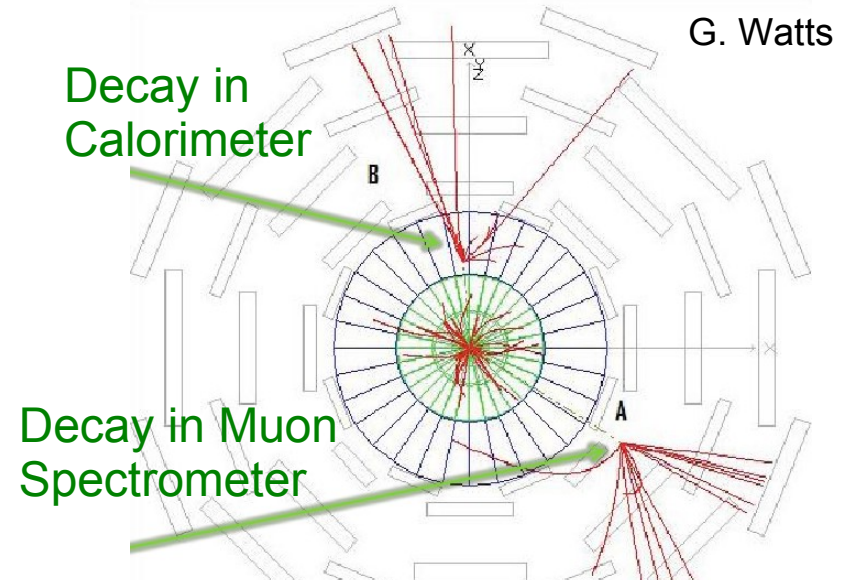
- SUSY MSUGRA/CMSSM is starting to be fine-tuned
 - Might be its last year to live...
- With more than 5 fb^{-1} , SUSY prod. mechanisms open up → exclusive 3rd generation and gaugino production
Look for “Natural SUSY”
- Focusing more and more on non-CMSSM scenarii that make SUSY harder to find:
“Split”, “compressed”, R-parity violation
More exotic signatures:
 - SUSY with low Missing ET
 - Multi-jet resonances
 - Long-Lived Particles (R-hadrons, staus) → see next slides

Long-Lived Particles



Long-Lived Particles

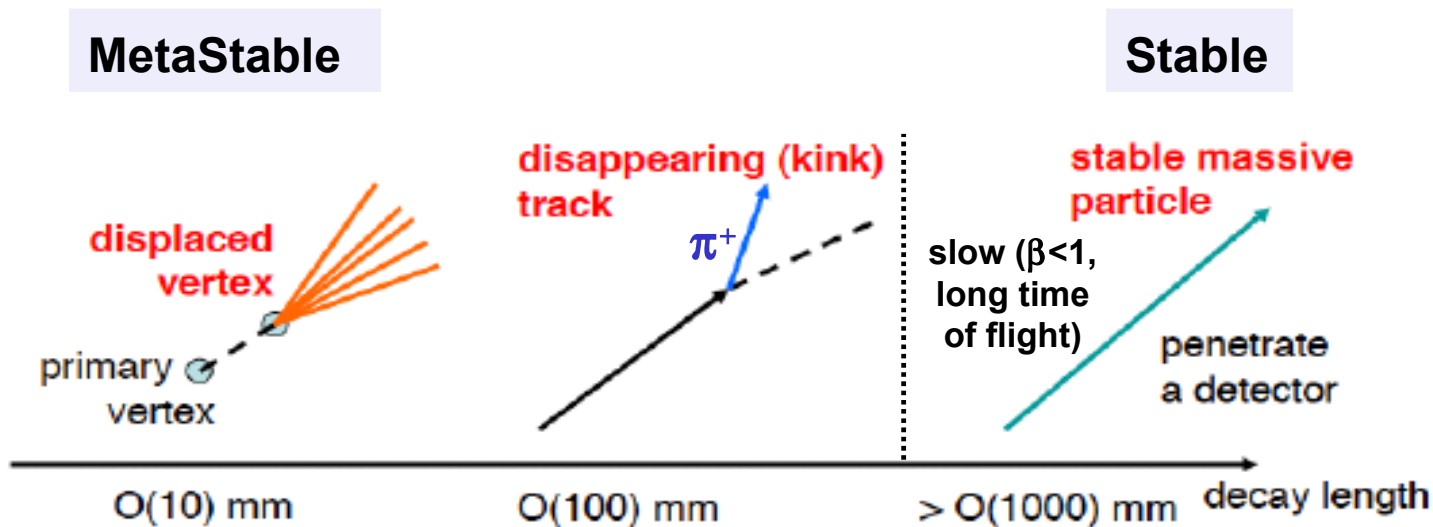
- Predicted by:
 - SUSY (R-parity violating or split/compressed mass spectra): stau, or gluino/stop hadronized into R-hadrons
 - Hidden Valley
- Experimentally very diverse:
 - Depends widely on particle's properties: life-time, charge, decay
 - highly displaced vertices
 - highly ionizing (dE/dx)
 - slow (time-of-flight)
 - kinked tracks
 - disappearing tracks
 - out-of-time (wrt collision) decay



SUSY and Long-Lived Particles

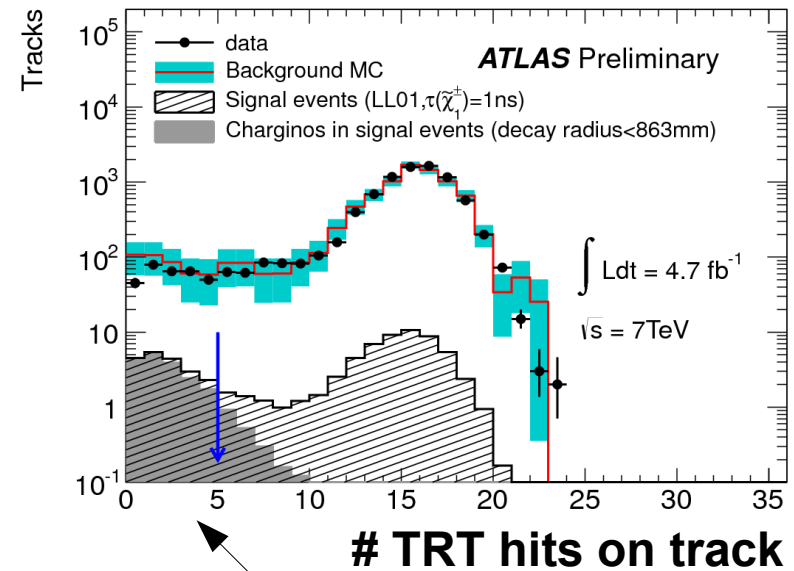
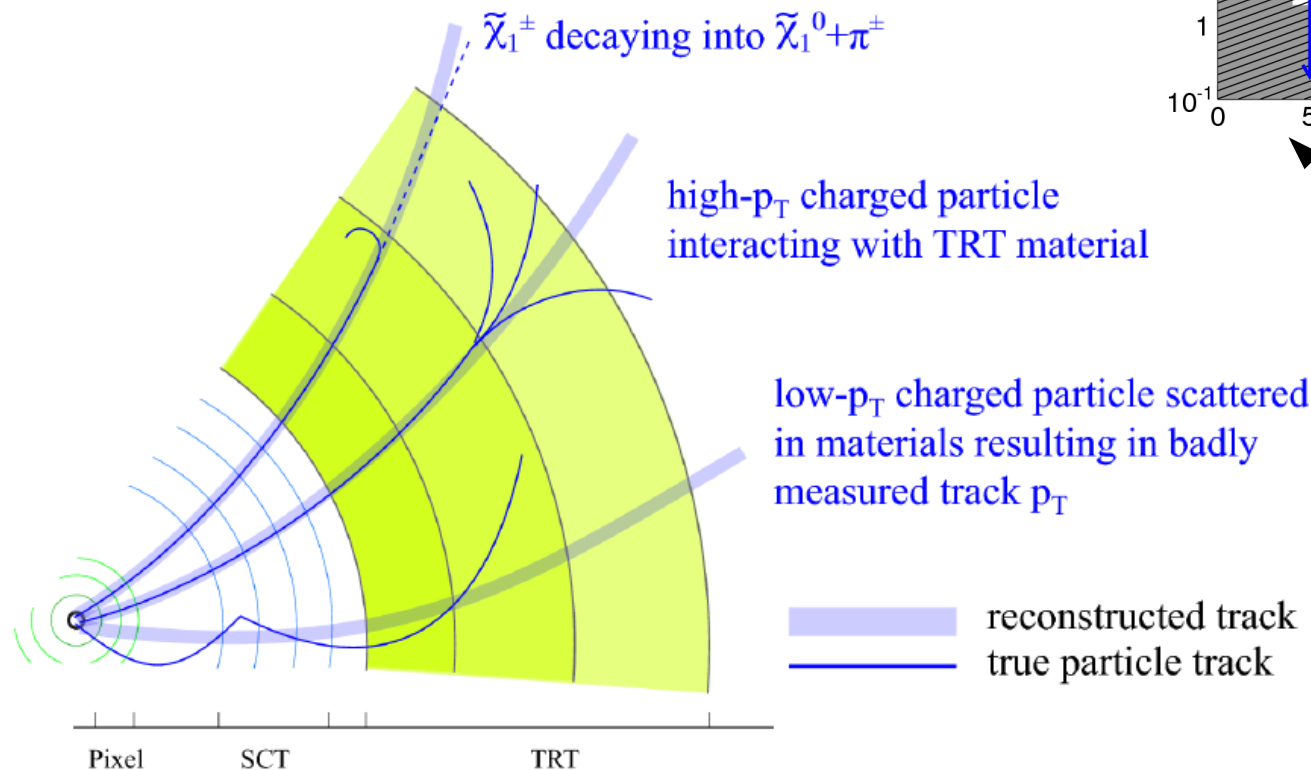
Three main mechanisms

- R-Parity violation: Lifetime proportional to $\lambda^2, \lambda'^2, \lambda''^2$ → Displaced vertex if $\lambda, \lambda', \lambda'' < 10^{-7}$
- Low $\Delta m(\tilde{\chi}_1^+ - \tilde{\chi}_1^0) \sim 100$ MeV in AMSB → Low π emitted, kinked track
- Low $\Delta m(\tilde{g}/\tilde{q} - \tilde{\chi}_1^0)$ for coloured particles → R-hadron (\tilde{g} or \tilde{q})
- Weak coupling to \tilde{G} in GMSB → Stable sleptons



Long-Lived Particles: Disappearing Track

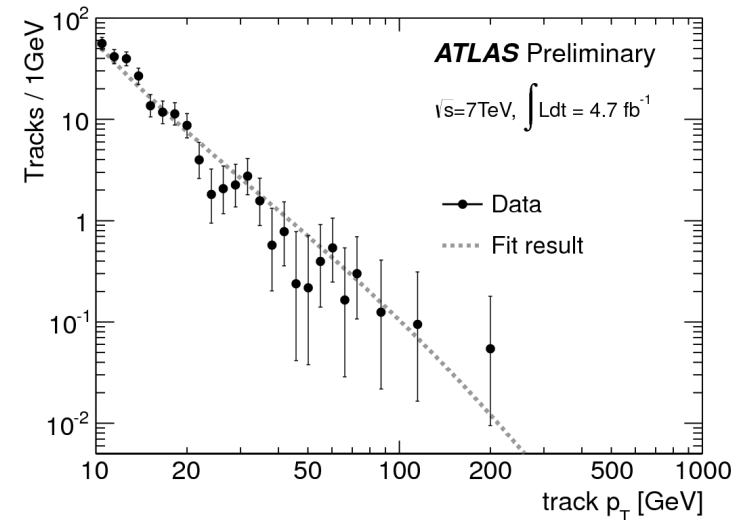
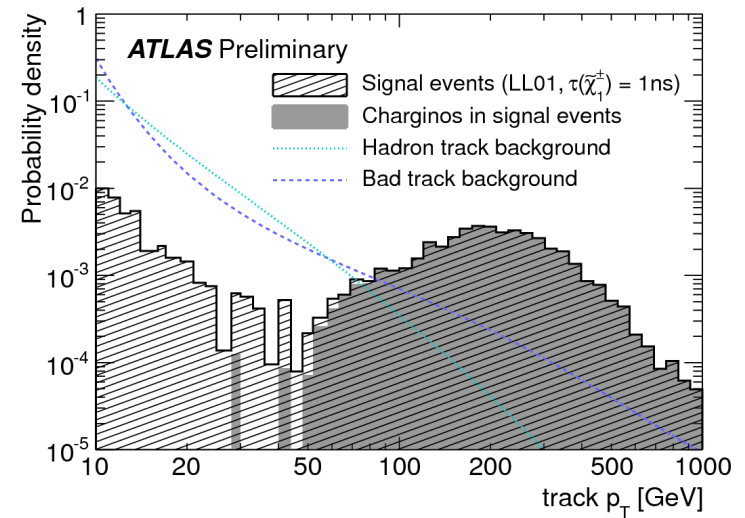
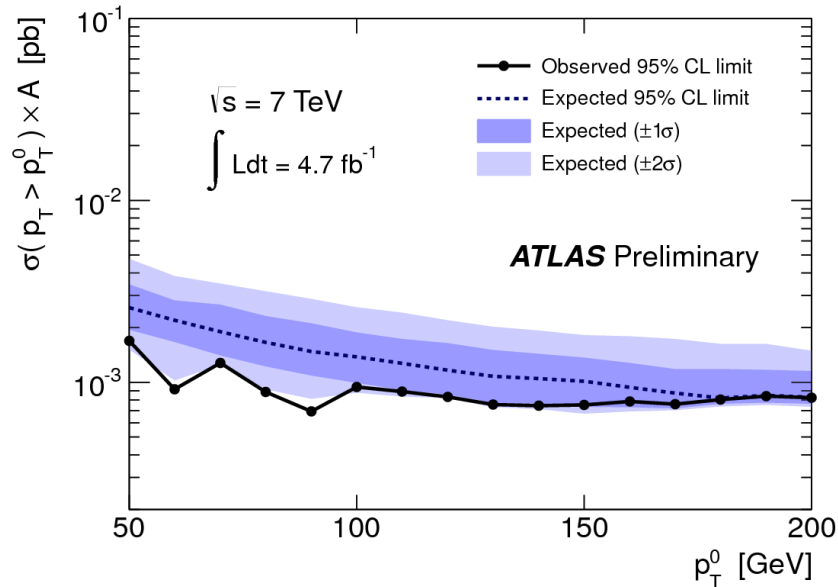
- $m(\text{chargino}) \sim m(\text{neutralino})$
- Chargino long-lived
- Chargino \rightarrow neutralino + soft pion



Select tracks
With less than
5 TRT hits

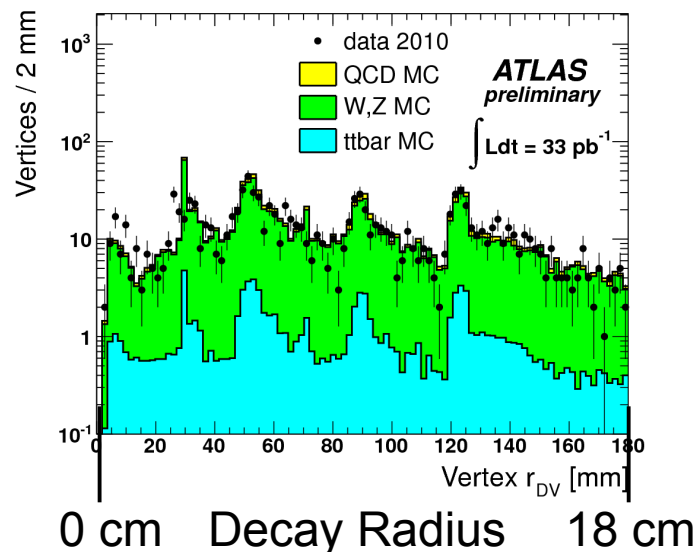
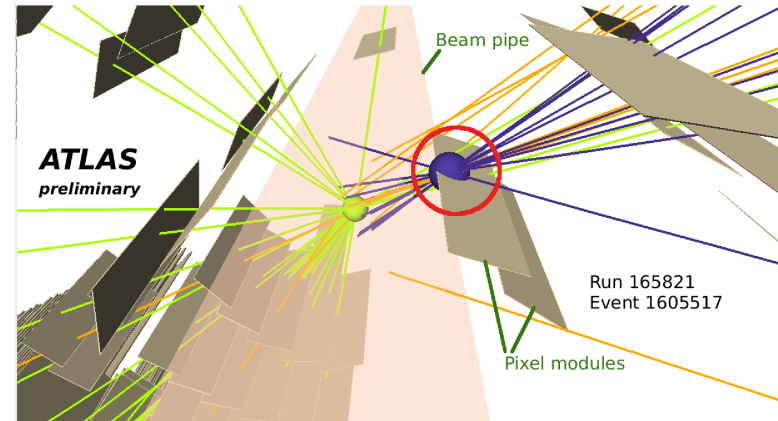
Long-Lived Particles: Disappearing Track

- Look at events with at least 3 jets and large missing ET
- Discr. Variable: p_T of tracks with less than 5 TRT hits

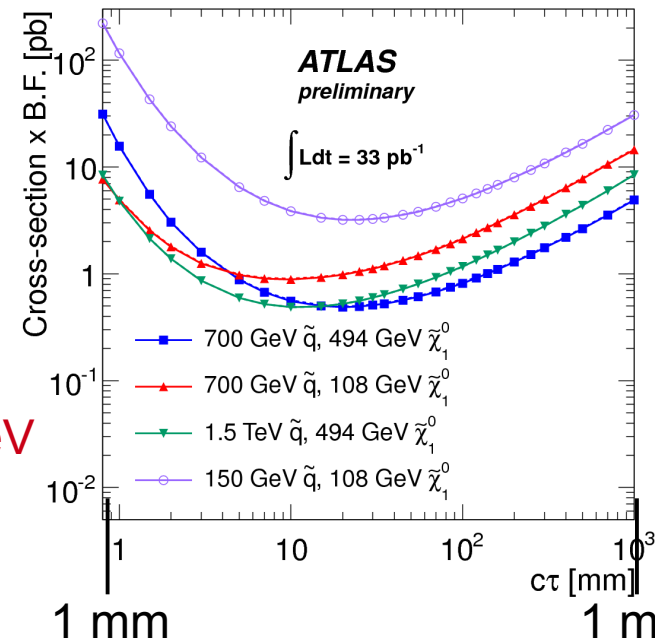


Long-Lived Particles: Decay in the Inner Detector

- R- hadrons (hadronized squarks or gluinos)
- Vertex outside the beampipe, in association with a high- p_T muon
- Requires good understanding of tracking, detector passive material

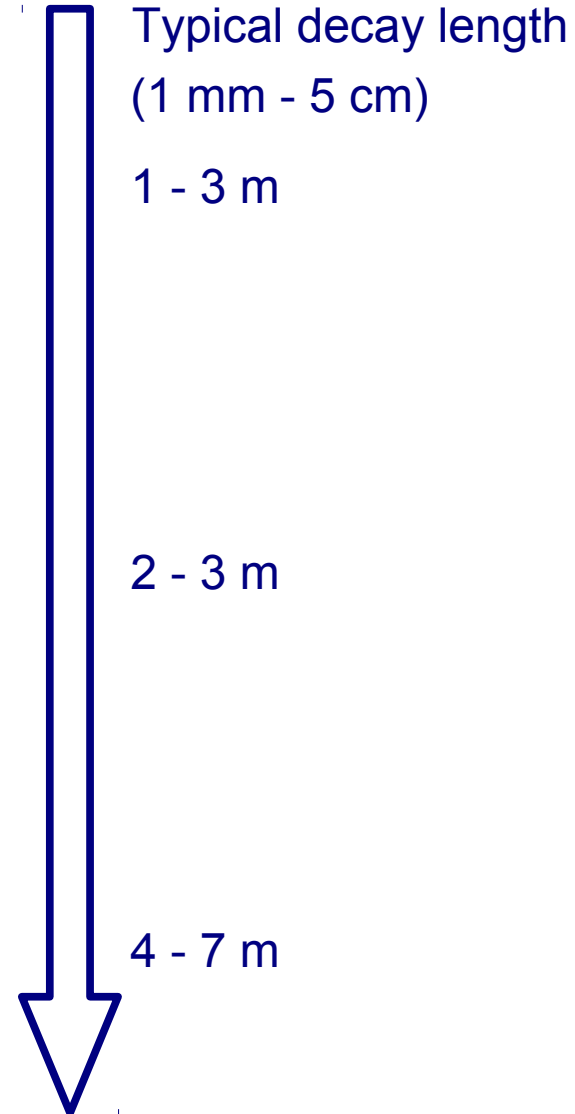


Signal Region:
 * Ntracks > 4
 * Vertex Mass > 10 GeV

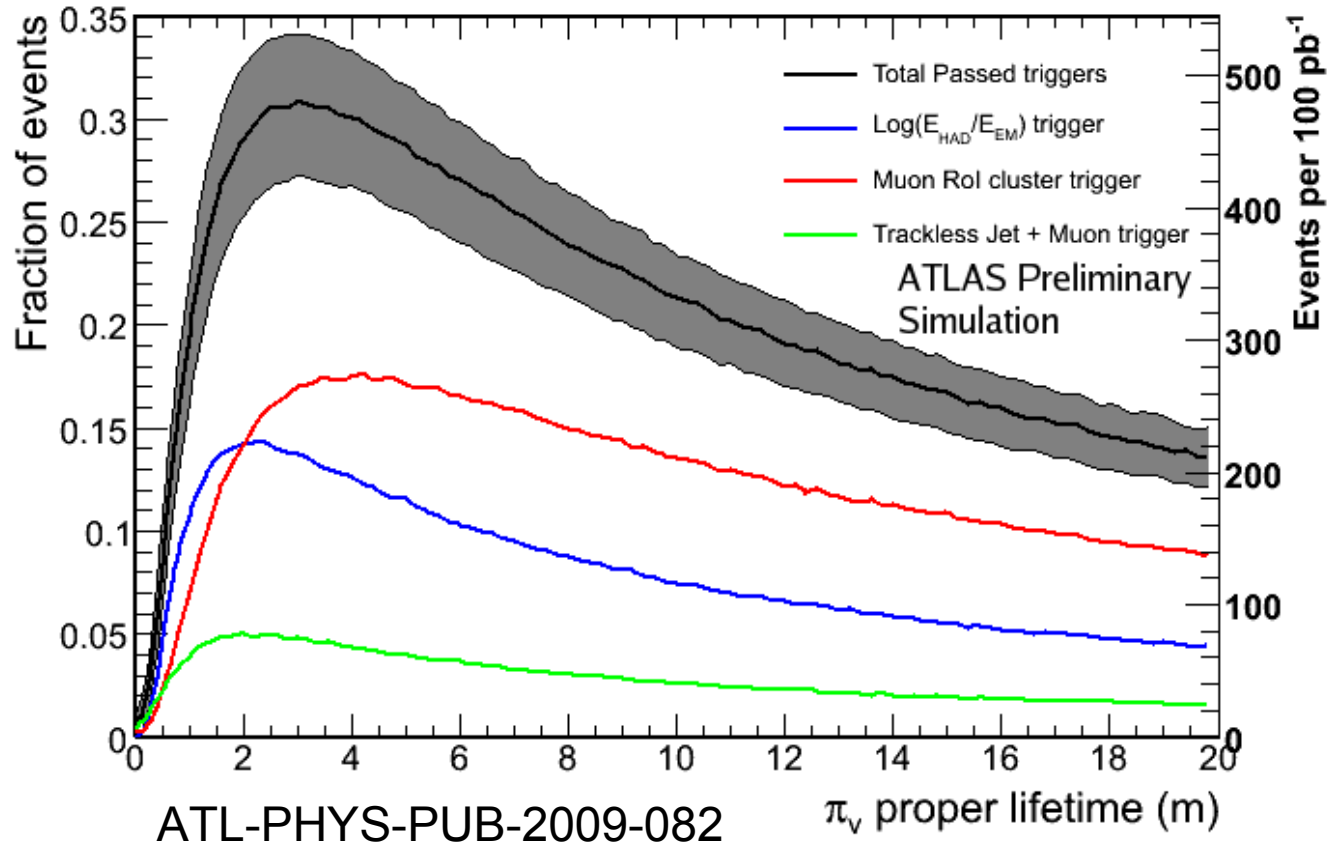


Long-Lived Particle Vertex Triggers

- (b-tagging triggers)
- Trackless jet trigger:
 - decays late in inner detector
 - jet $E_T > 35$ GeV
 - no tracks with $p_T > 1$ GeV near jet
 - muon spectrometer activity
- Hadronic / EM (decays beyond the EM calorimeter)
 - jet $E_T > 35$ GeV
 - no tracks with $p_T > 1$ GeV near jet
 - $E_{\text{had}} / E_{\text{EM}} > 10$
- Muon spectrometer cluster trigger
 - 3 muon triggers close from each other
 - no jets, no tracks

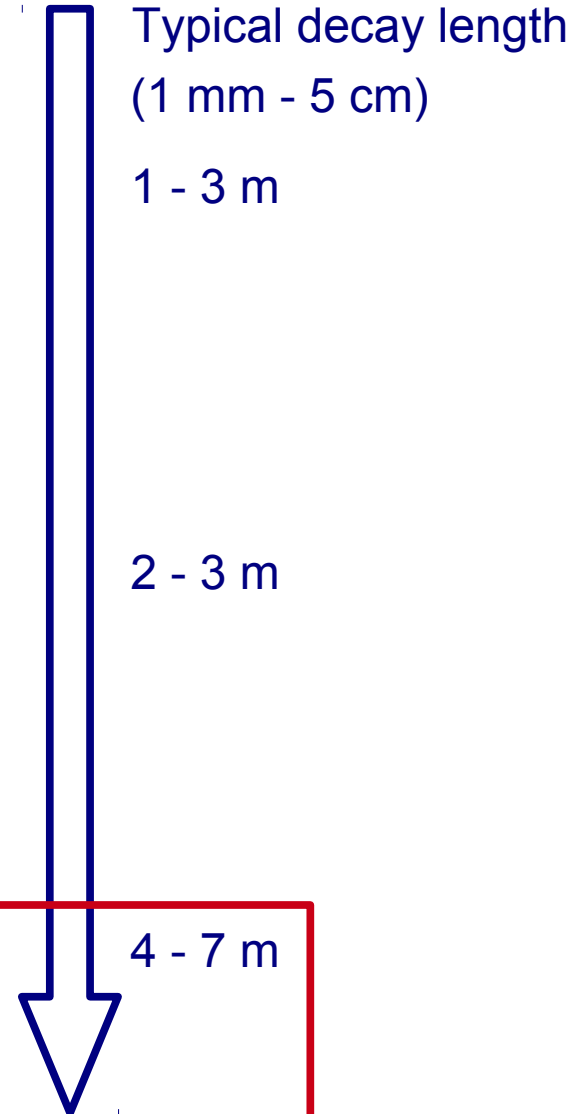


Long-Lived Particles Vertex Triggers



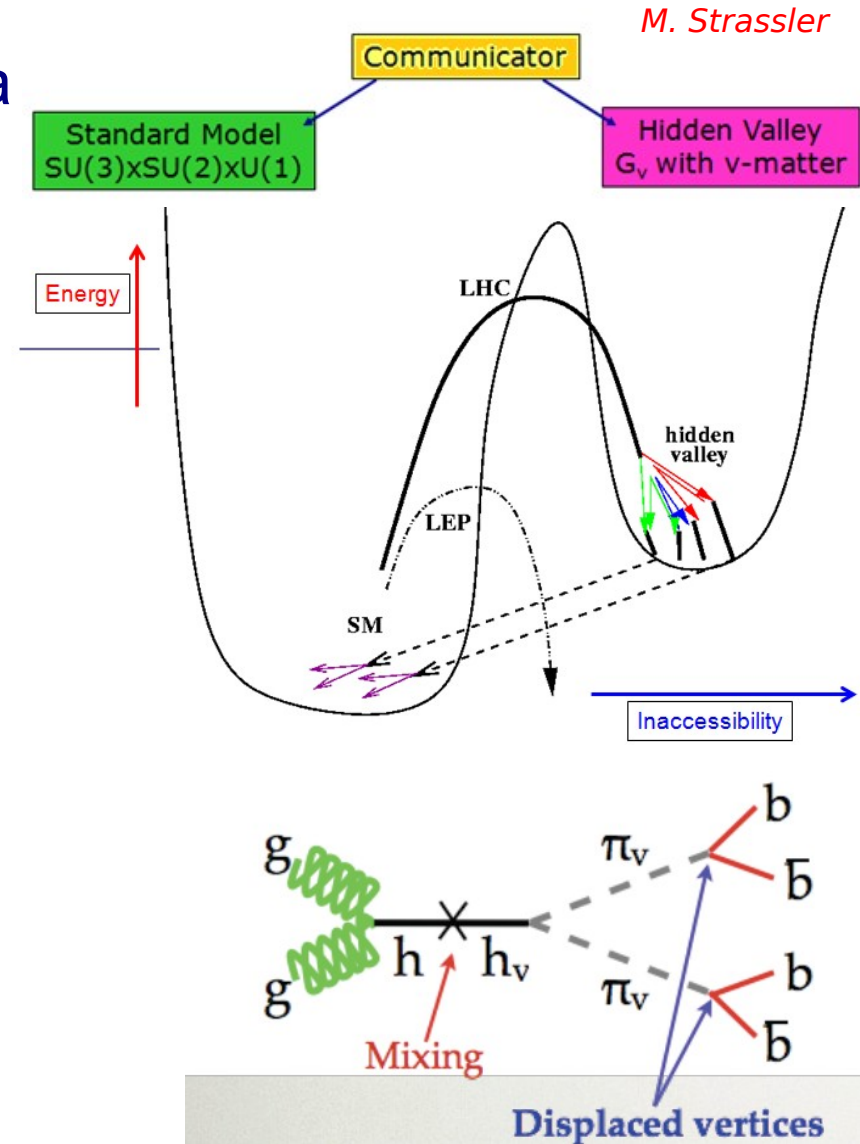
Long-Lived Particle Vertex Triggers

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- Hadronic / EM (decays beyond the EM calorimeter)
 - jet $E_T > 35$ GeV
 - no tracks with $p_T > 1$ GeV near jet
 - $\log(E_{had} / E_{EM}) > 1.0$
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 - 3 muon triggers close from each other
 - no jets, no tracks



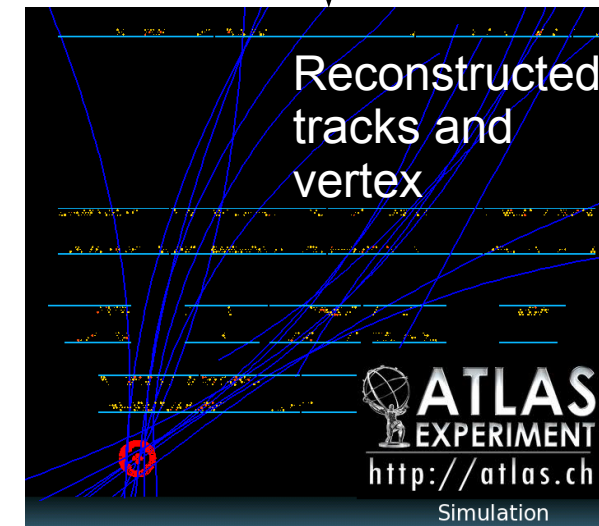
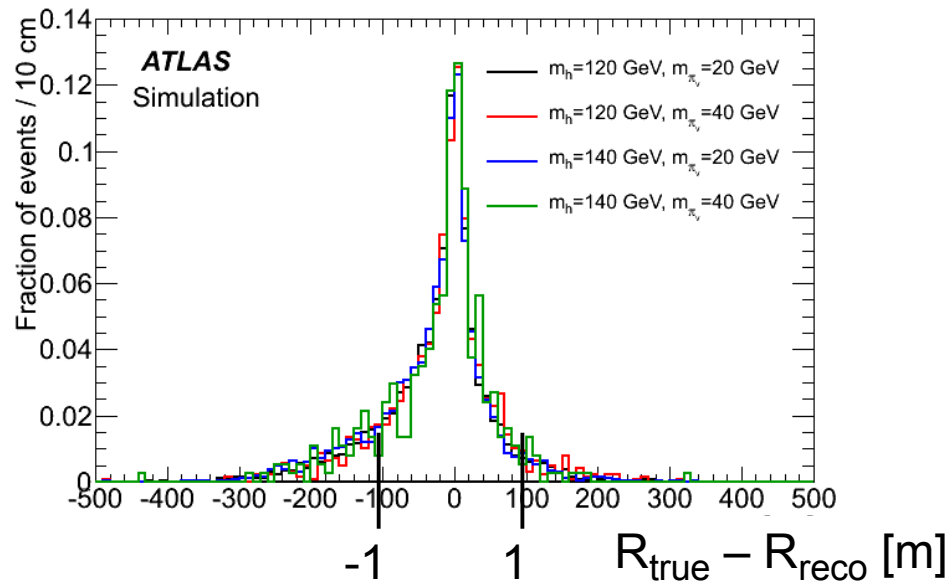
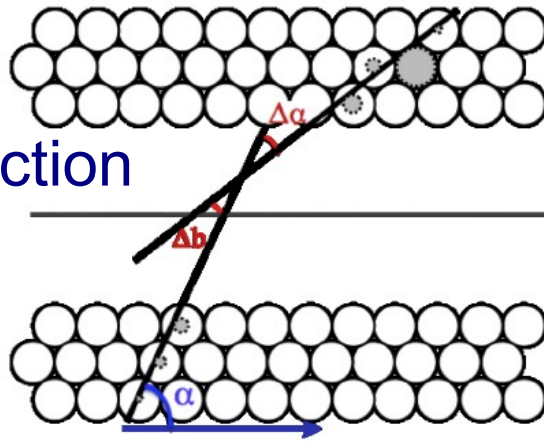
Long-Lived Particles: Decay in the Muon Spectrometer

- Hidden-Valley theories predict a hidden sector coupled to the SM only through some **heavy communicator** → **weakly coupled** → **long-lived particles**
- Ex: $h \rightarrow h_\nu \rightarrow \pi_\nu \pi_\nu \rightarrow 4b$'s
- Life-time of π_ν is unknown
- Look for 2 pairs of b-jets appearing outside the calorimeter.
- Sort of b-tagging with the Muon Spectrometer!



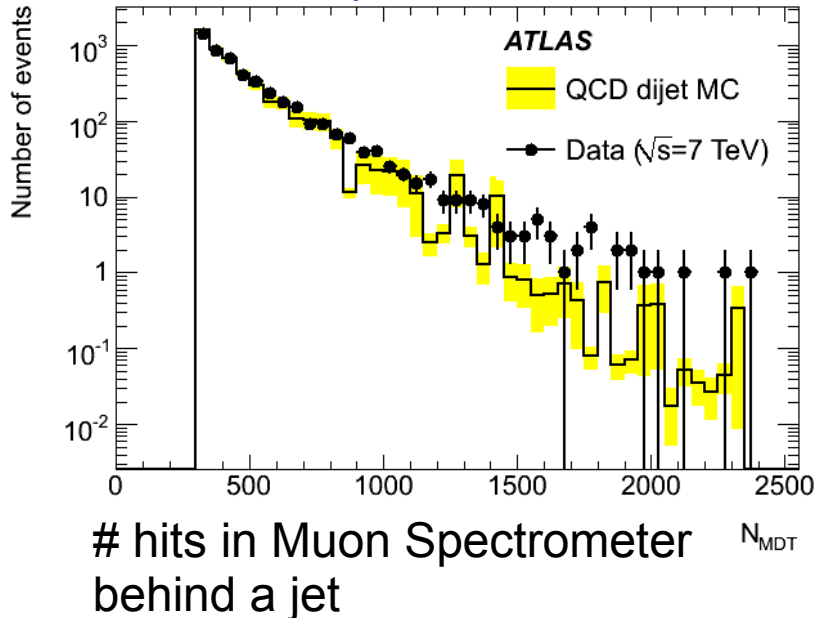
Long-Lived Particles: Vertex Reconstruction in the Muon Spectrometer

- Very high occupancy
- Partial track reconstruction
- Efficiency $\sim 50\%$
- Very coarse spatial resolution:

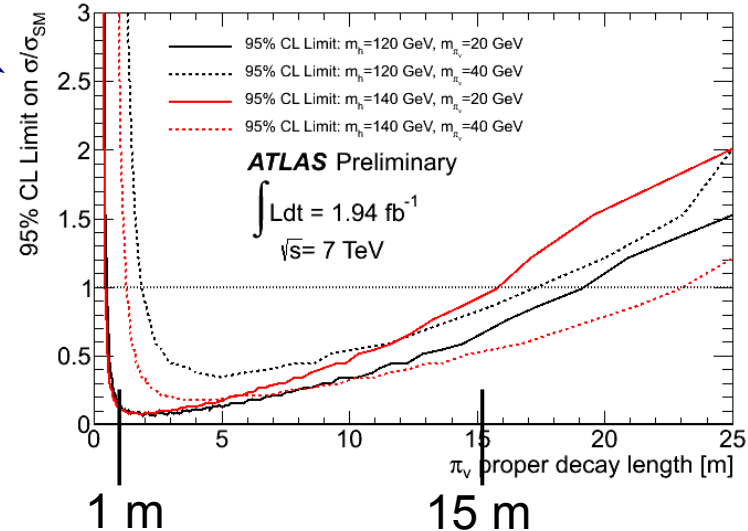
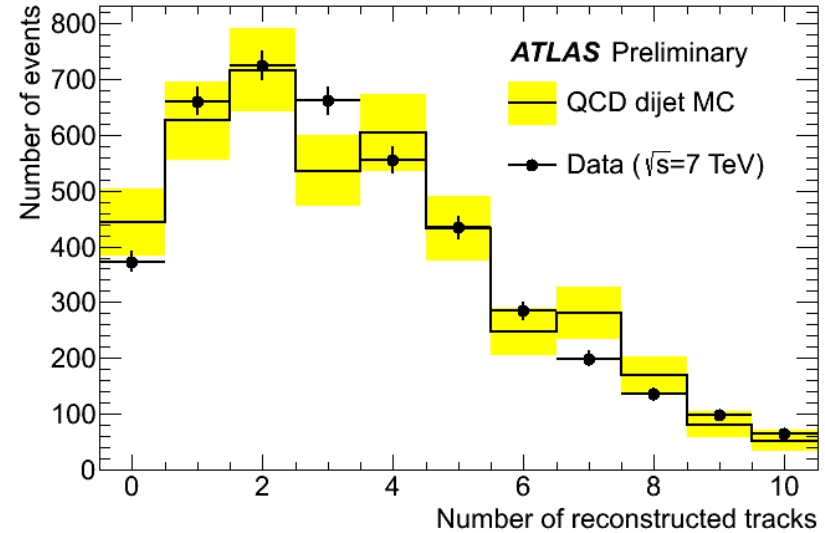


Long-Lived Particles: Decay in the Muon Spectrometer

- Validation of simulation with punch-through events
- **Note: punch-through's very well described by the simulation!**
- After final selection: no event observed (exp: 0.03 ± 0.02 ev.)



Tracks caused by jets in Muon Spectrometer



Backup

Sidenote: How to avoid bias

- “The easiest person to fool is yourself” - R. Feynman
- You have two methods to estimate your background
 - Method 1 predicts $N(\text{background events}) = 2.2 \pm 0.2$ (sys.)
 - Method 2 predicts $N(\text{background events}) = 5 \pm 3$ (sys.)
- Case 1: choice made a priori (before looking at the data)
- Methods are compatible and method 1 gives best sensitivity
=> choose Method 1
- In the data, you observe $N(\text{events}) = 7$ events
- 3-sigma excess: you will not graduate this year

Sidenote: How to avoid bias

- “The easiest person to fool is yourself” - R. Feynman
- You have two methods to estimate your background
 - Method 1 predicts $N(\text{background events}) = 2.2 \pm 0.2$ (sys.)
 - Method 2 predicts $N(\text{background events}) = 5 \pm 3$ (sys.)
- Case 2: choice made a posteriori
- You observe 7 events (and you must graduate before the end of the year)
- You decide to choose the most conservative method, just to be more conservative
- No excess: you just missed the Nobel price.