

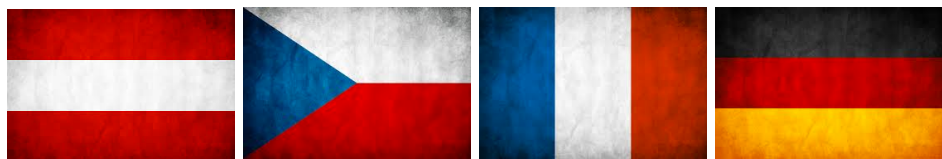


Mesure gravitationnelle de l'antimatière

Patrick Nedelec

IPNL-UCBL

Lyon, France



A E \bar{g} I S collaboration



Stefan Meyer Institute



CERN



Czech Technical University



ETH Zurich



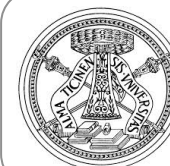
University of Genova



University of Milano



University of Padova



University of Pavia



Institute of Nuclear Research of the Russian Academy of Science



Max-Planck Institute Heidelberg



Politecnico di Milano



University College London



University of Bergen



University of Bern



University of Brescia



Heidelberg University



University of Lyon 1



University of Oslo



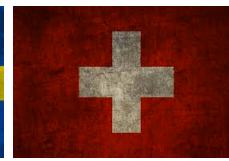
University of Paris Sud



University of Trento



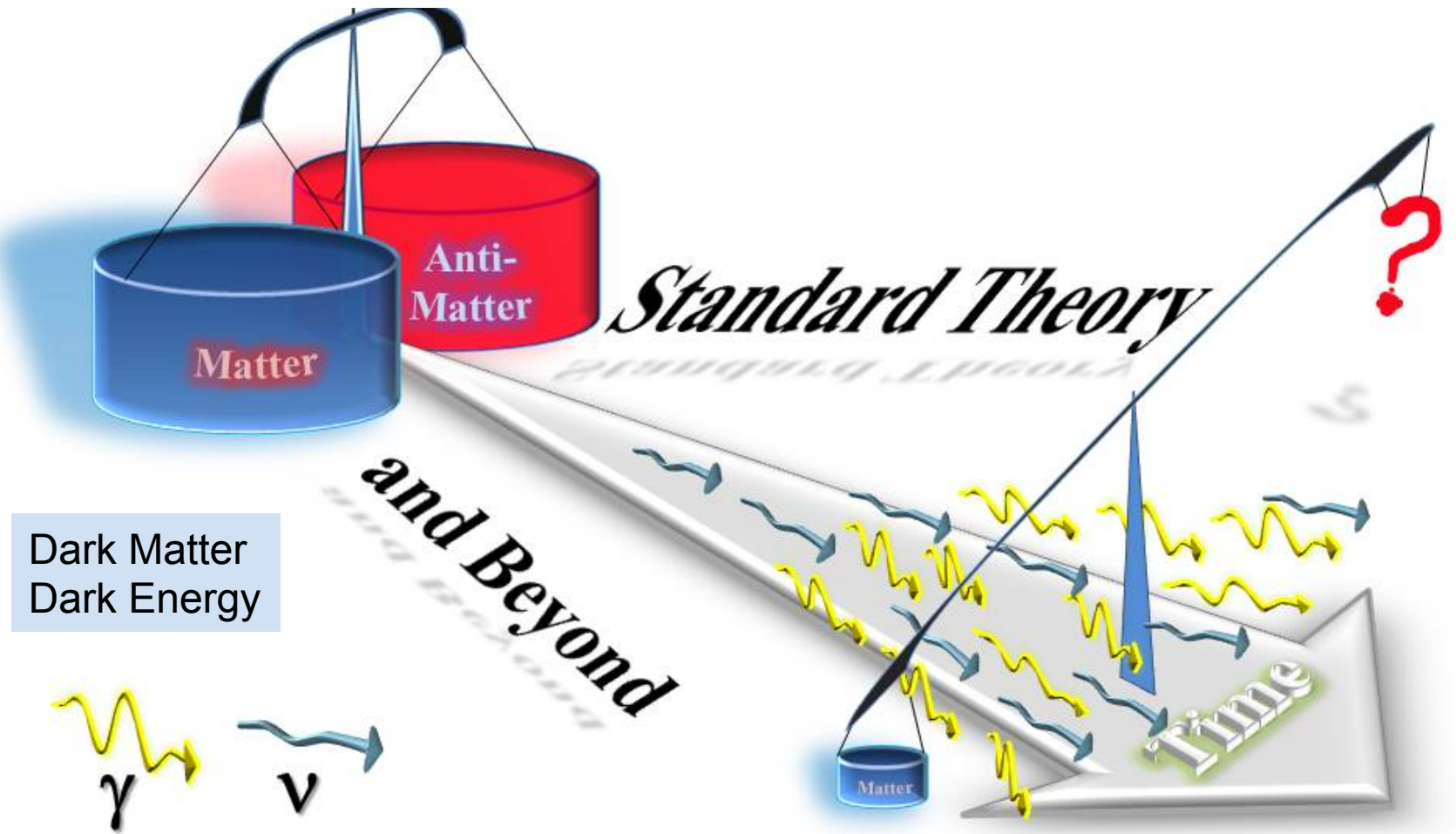
INFN sections of: Genova, Milano, Padova, Pavia, Trento



Plan

- Motivations et buts
- Principe
- État de l'art
- Premiers résultats
- Futur proche - ELENA

Matière-Antimatière et Univers



Antimatière & (anti ?) gravité



$$V = -G \frac{MM'}{r} \left(1 \mp ae^{-\frac{r}{v}} + be^{-\frac{r}{s}} \right)$$



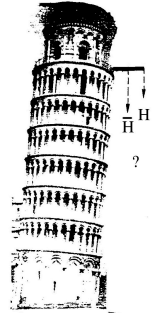
Newton

Supergravité N=2,...,8 : anti-graviton
-> gravité répulsive !

- Contraintes exp. : M. Nieto & al. Phys. Rep. 205 (1991)
- Motivation pour l'antigravité : G. Chardin, Hyp. Int. 109, 83 (1997)
- Violations de Lorentz & CPT : V.A. Kostelecky et al., Phys. Rev. D83 (2011)
- DM & DE : gravitation. pol. & dipole of vacuum : D.S.Hajdukovic, Astro Space Sciences 338, (March 2012)
- Nouvelles expériences :
 - ALPHA (CERN-AD 2013) ; AEGIS (CERN-AD6 -2012) ; Gbar (CERN-AD 2016)

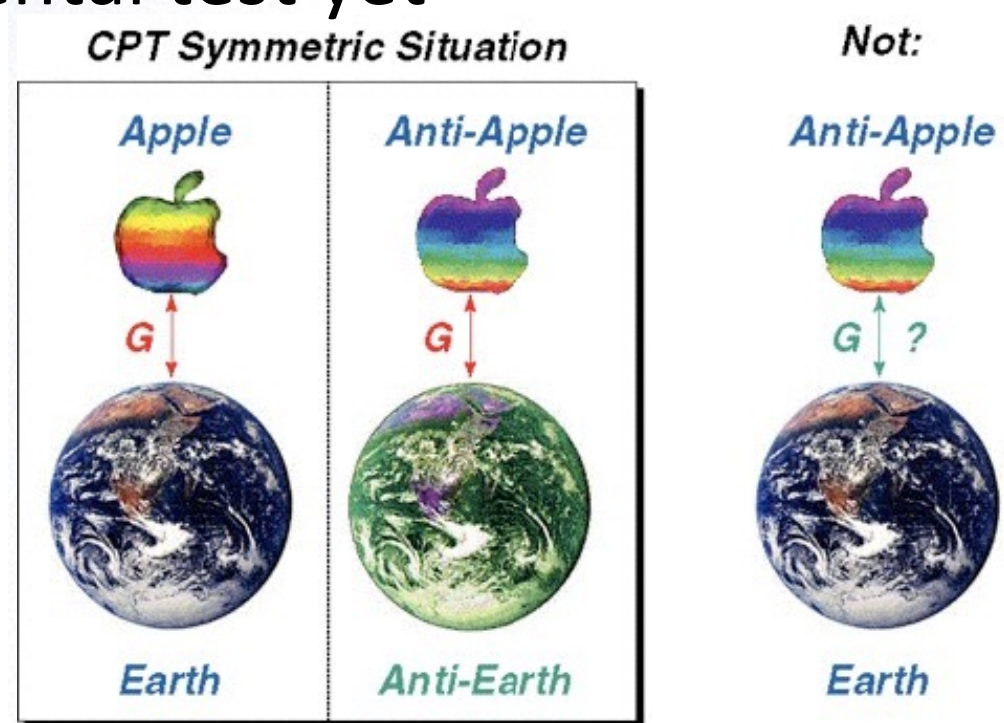
Goals and Motivations

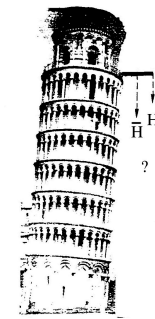
- Measure $\vec{P} = \bar{m} \times \vec{g}$ to 1%
 - Looking at **free fall** of cold \bar{H} atoms
- Test of **WEP** on **antimatter** in the Earth gravitational field
 - First (?) direct measurement
 - Need a large number of atoms (10^5)
- Extra
 - Spectroscopy, Ps physics, H-Hbar comparison,...



Gravitational acceleration of antimatter

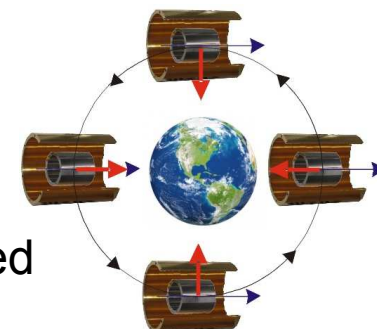
- No direct CPT test
 - Weak Equivalence Principle
- No precise experimental test yet
- Highest precision
 - With neutral antimatter





year	investigator	accuracy	method
500	Philoponus	"small"	drop tower
1585	Stevin	$5 \cdot 10^{-2}$	drop tower
1590	Galileo	$2 \cdot 10^{-2}$	pendulum, drop tower
1686	Newton	$1 \cdot 10^{-3}$	pendulum
1832	Bessel	$2 \cdot 10^{-5}$	pendulum
1910	Southerns	$5 \cdot 10^{-6}$	torsion balance
1918	Zeeman	$3 \cdot 10^{-8}$	torsion balance
1922	Eotvos	$5 \cdot 10^{-9}$	torsion balance
1923	Potter	$3 \cdot 10^{-9}$	pendulum
1935	Renner	$2 \cdot 10^{-9}$	torsion balance
1964	Dicke et al	$3 \cdot 10^{-11}$	torsion balance
1972	Braginski, Panov	$1 \cdot 10^{-12}$	torsion balance
1976	Shapiro	$1 \cdot 10^{-12}$	lunar laser ranging
1987	Niebauer et al.,	$1 \cdot 10^{-10}$	drop tower
1989	Heckel	$1 \cdot 10^{-11}$	torsion balance
1990	Adelberger	$1 \cdot 10^{-12}$	torsion balance
1999	Baebler	$1 \cdot 10^{-13}$	torsion balance
2016	Microscope	$1 \cdot 10^{-15}$	space
20??		$1 \cdot 10^{-18}$	space

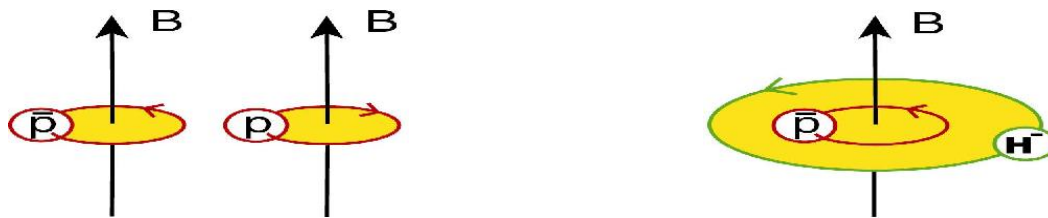
Equivalence
Principle
Tests for
Matter systems



expected

Limites indirectes

- Gravitational Clock Redshift for p/pbar



Gravitational red shift for a clock: $\Delta\omega / \omega = g h / c^2$

→ Antimatter and matter clocks run at different rates
if g is different for antimatter and matter

$$\frac{\Delta\omega_c}{\omega_c} = 3(\kappa - 1) \frac{U}{c^2}$$

for tensor gravity
(would be 1 for scalar gravity)

Hughes and Holzscheiter,
Phys. Rev. Lett. 66, 854 (1991).

grav. pot. nergy difference
between empty flat space time
and inside of hypercluster of galaxies

Experiment: TRAP Collaboration, Phys. Rev. Lett. 82, 3198 (1999).

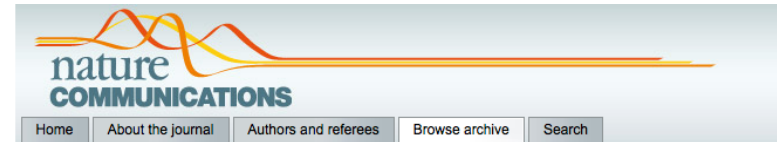
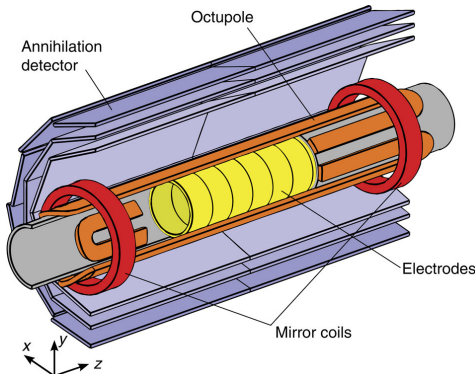
$$\frac{\Delta\omega_c}{\omega_{\partial c}} < 10^{-10} \quad \text{---} > \quad \kappa = 1 \pm (< 10^{-6})$$

Limites directes actuelles

- Antimatière
 - « Limites » ALPHA
 - Expérience de sédimentation

$$-65 < \frac{M_g}{M} < 75$$

Pas de contrainte très forte !

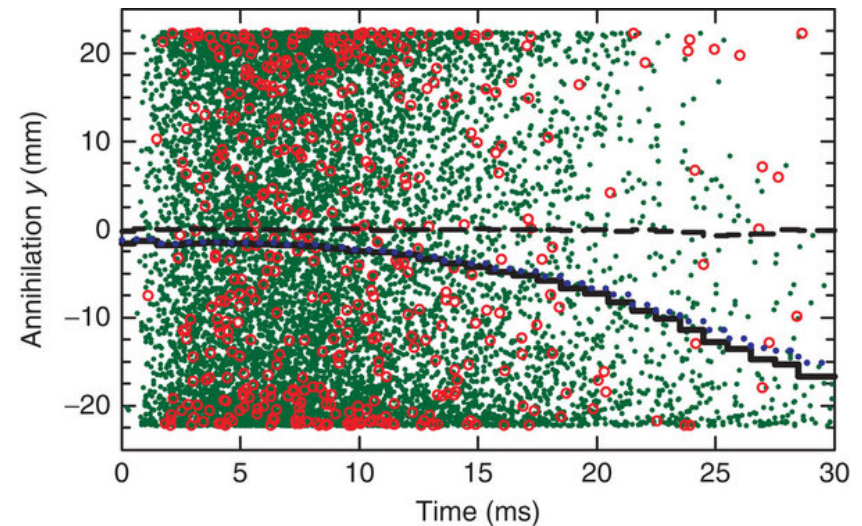


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NATURE COMMUNICATIONS | ARTICLE OPEN

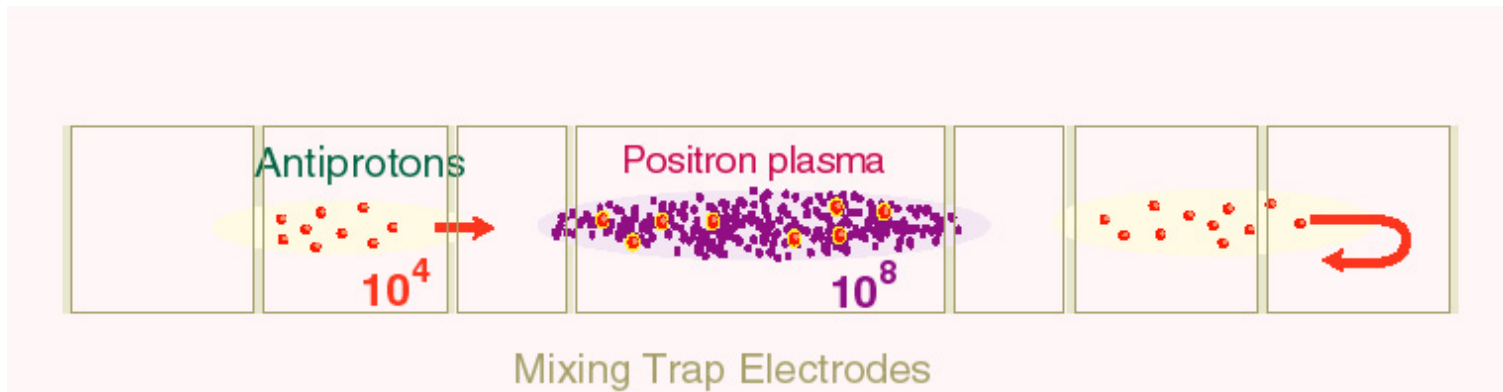
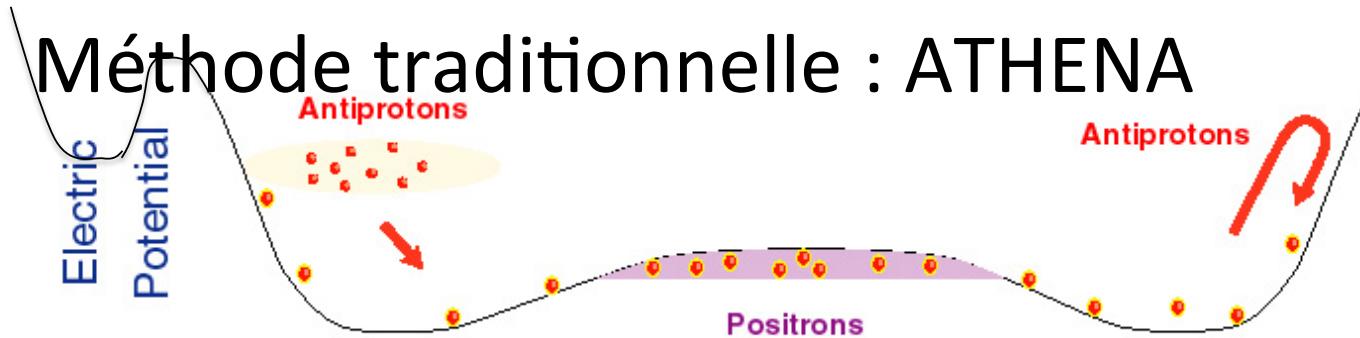
Description and first application of a new technique to measure the gravitational mass of antihydrogen

The ALPHA Collaboration & A. E. Charman

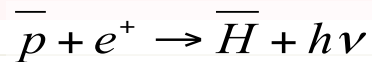


Antihydrogène : production

- Méthode traditionnelle : ATHENA



a) Radiative recombination



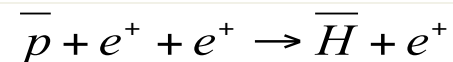
Populate low n ($n < 10$)

$$R \propto n_{e^+} T_{eff}^{-1}$$

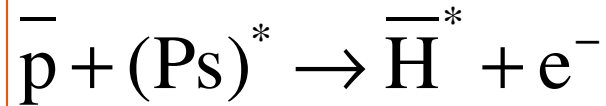
b) 3 body recombination

$$R \propto n_{e^+}^2 T_{eff}^{-9/2}$$

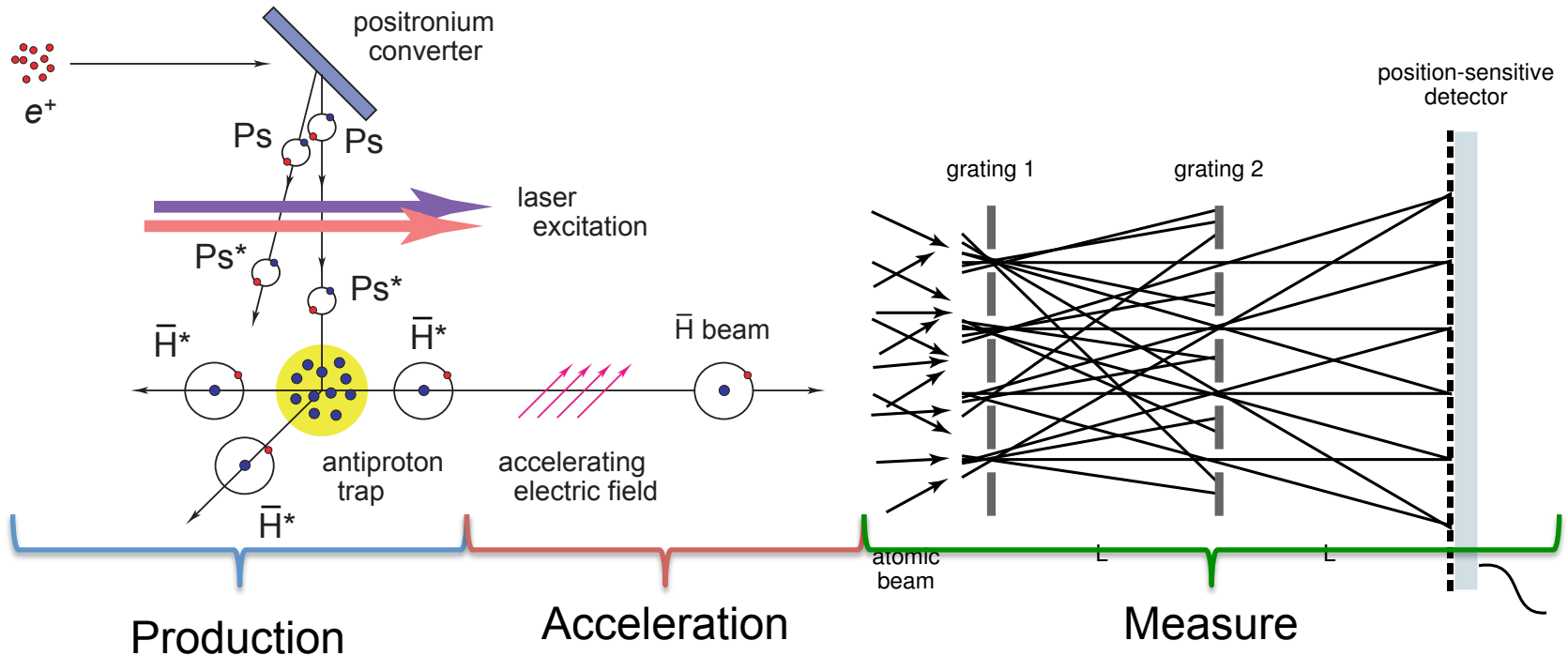
Populate high n, ionization by collisions and field ionization)



Cold pbar : AEgIS recipe



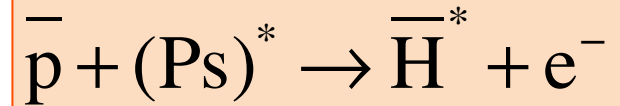
: Charge exchange reaction



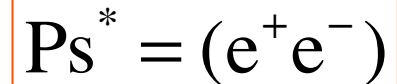
AEGIS $\bar{\text{H}}$ production way

- Antiprotons:

- $\bar{\text{p}}$ (5 MeV) from CERN AD
- Degrade & keep $E < 10$ keV
- Cool down (5T trap)



- Positronium



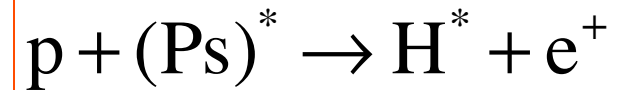
- e^+ on nano-porous SiO_2 , Rydberg states (n)

- Form $\bar{\text{H}}$ by interaction of Ps with $\bar{\text{p}}$ cloud

- Xsection: $\sigma \propto n^4$

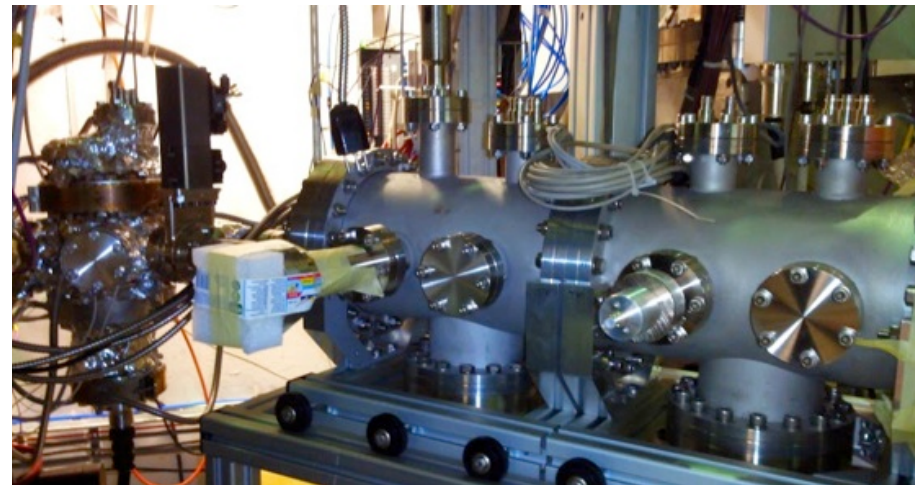
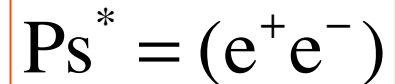
AEgIS H production way

- Adding a p beam gives C-conjugate:
 - p (2 keV)

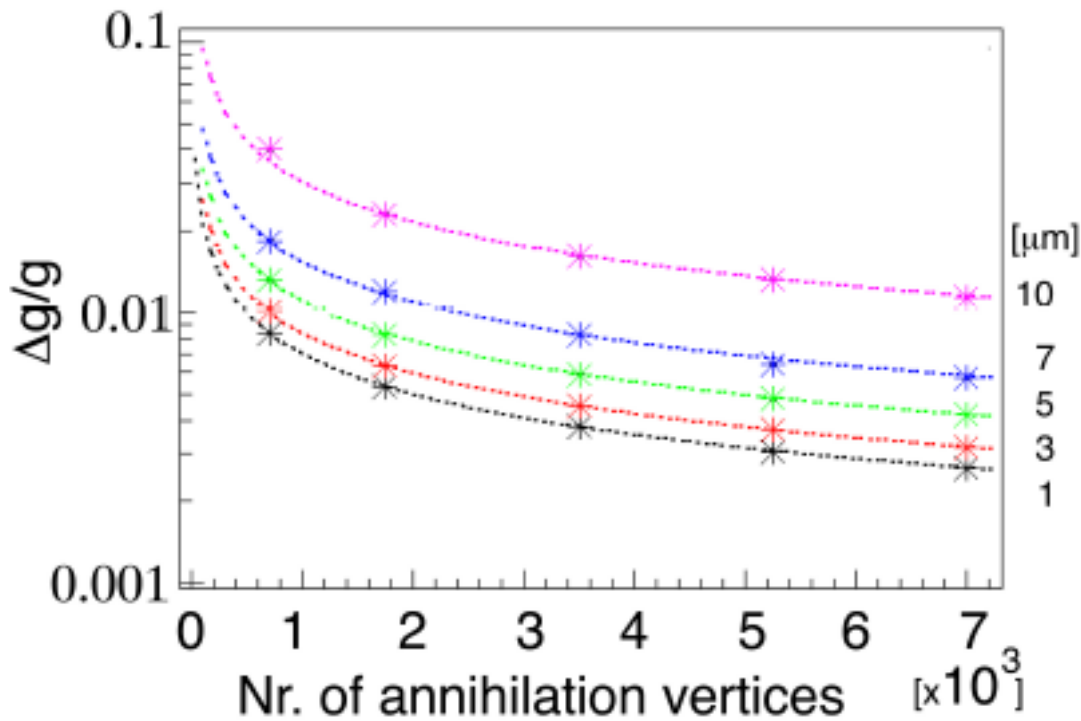


- Form H by interaction of Ps with p cloud

- Comparison H-Hbar
- Extra needs
 - P beam, H-detector
 - Installed in 2014

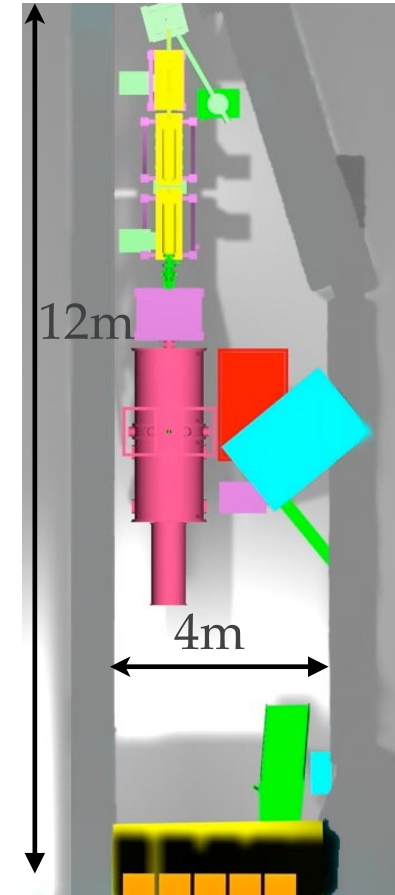
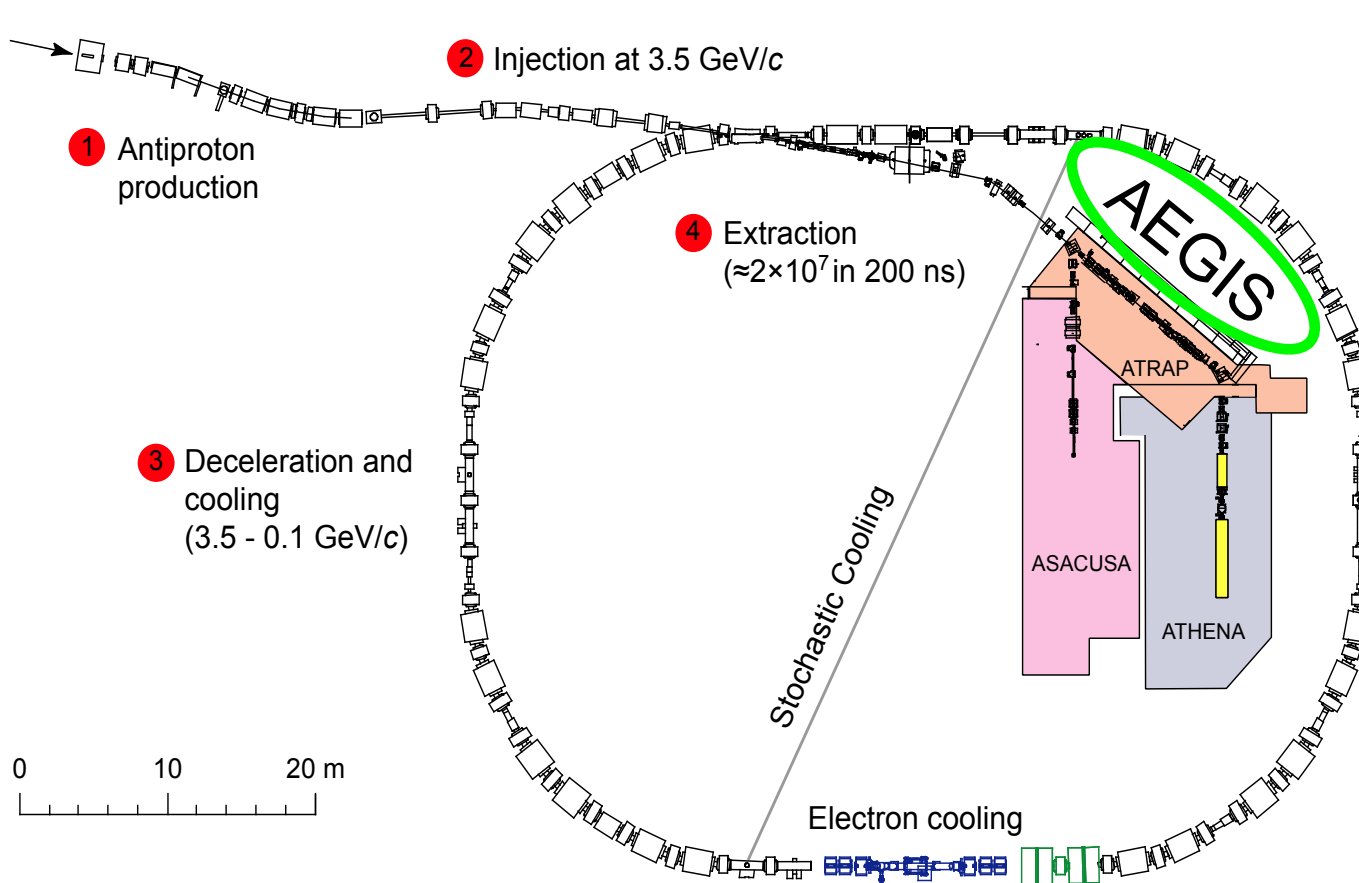


Measurement sensitivity

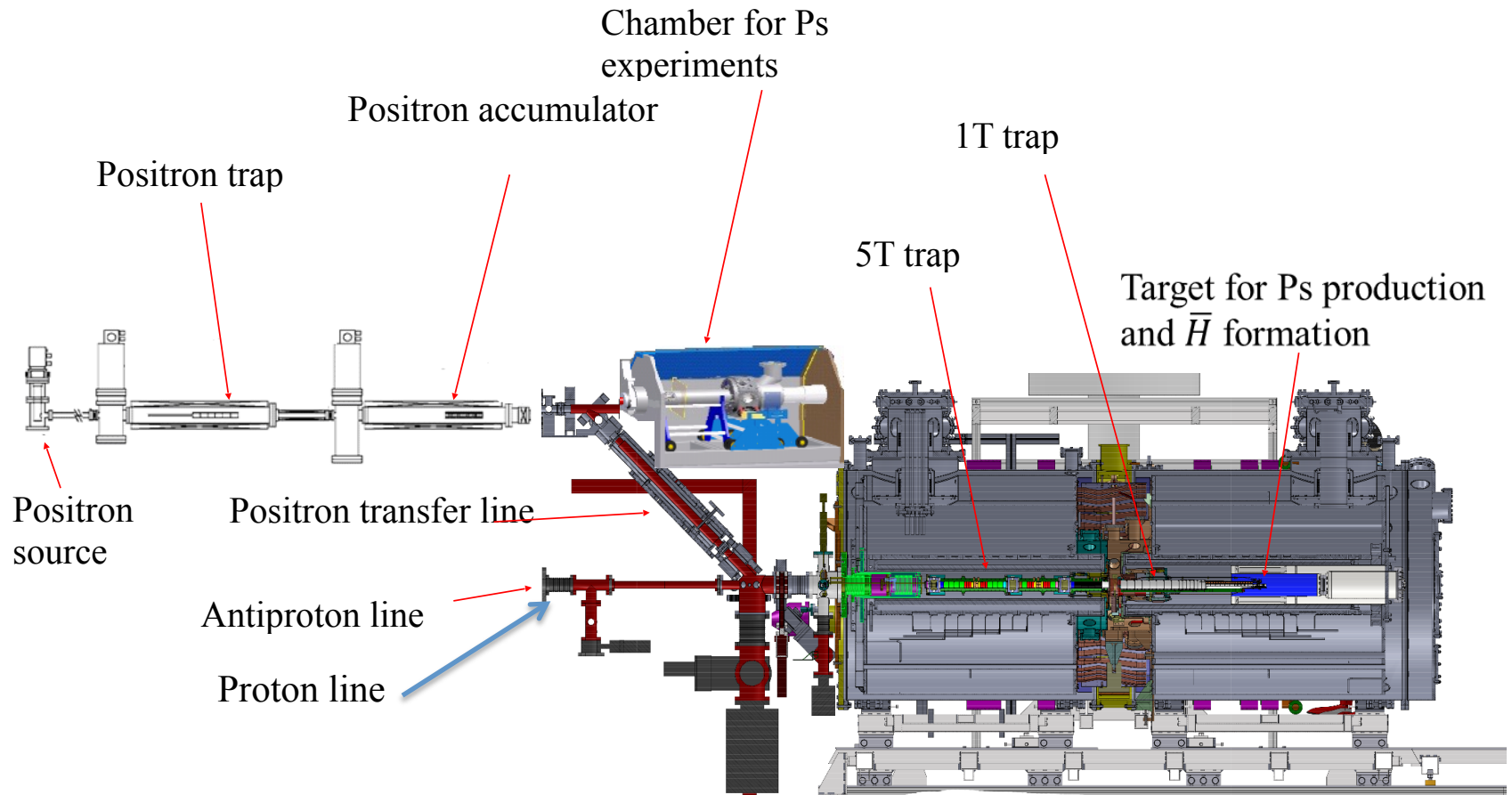


- 1% sensitivity achievable with ~ 600 annihilations
- Measurement resolution $\sim 3 \mu\text{m}$
- Cold antihydrogen necessary for bigger deflection and lower beam divergence.

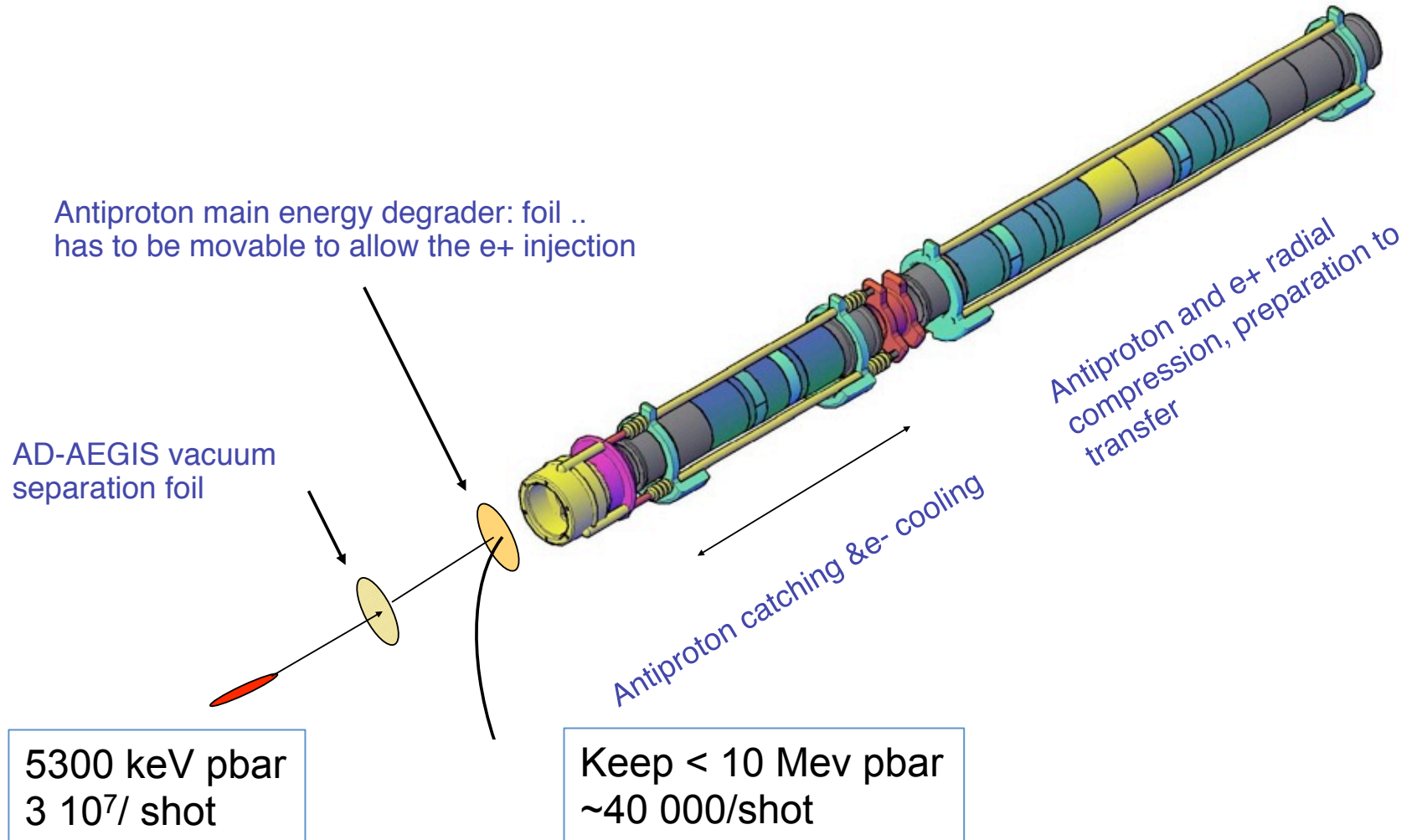
Experimental site CERN Anti Decelerator(AD) area



Experimental setup



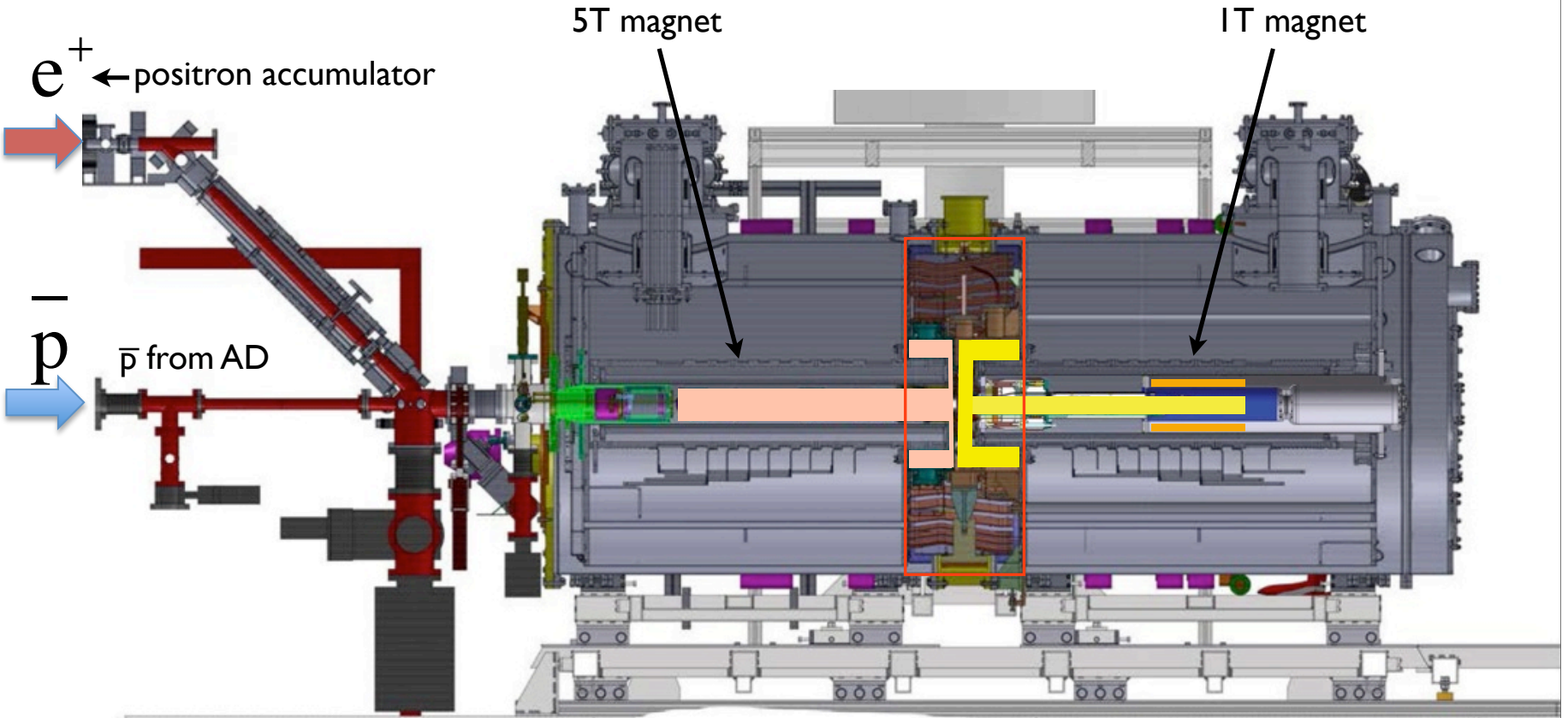
pbar degrader



Apparatus

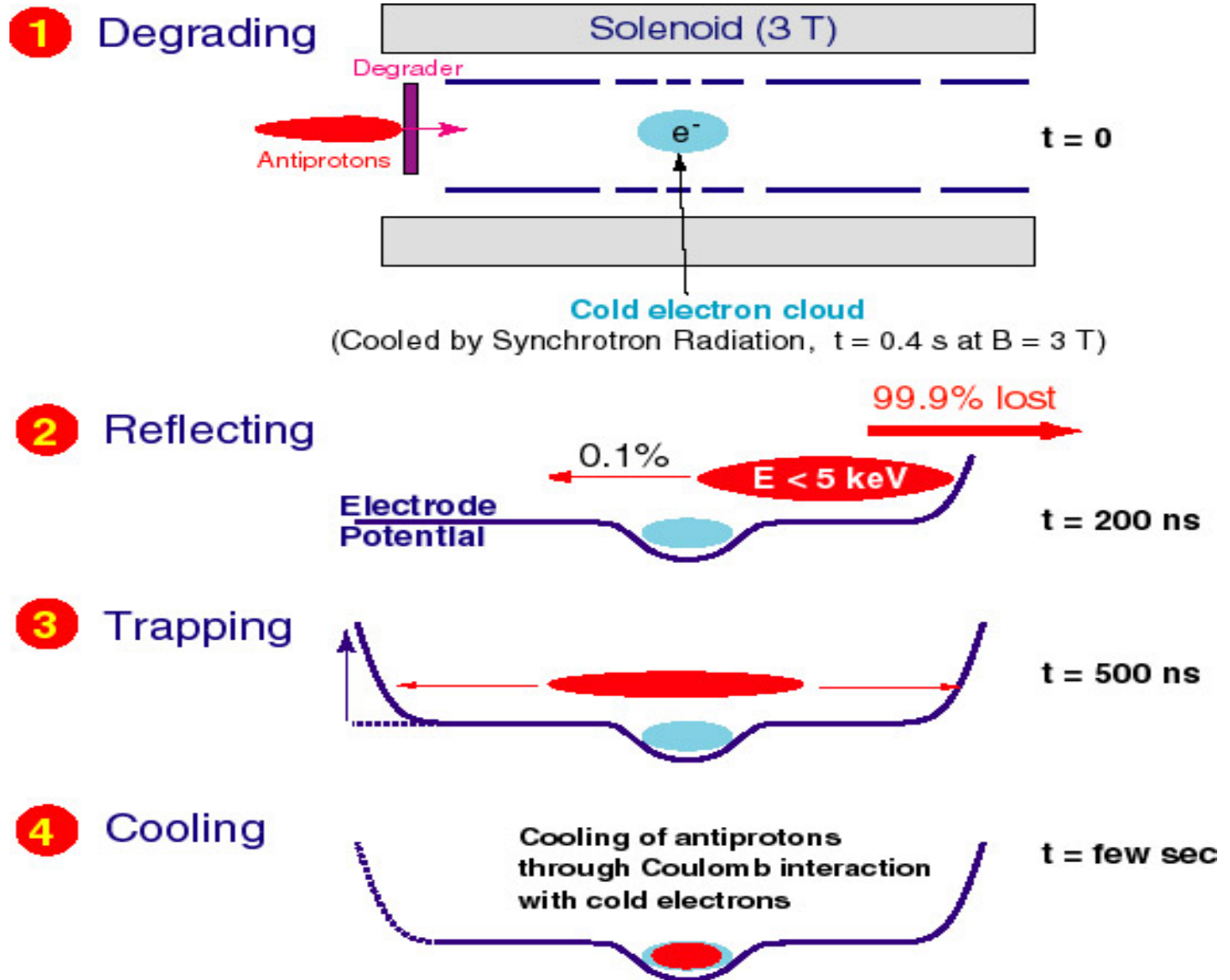
Pbar catching & Cooling
E+ storage (T~10K)

p-pbar cooling
Ps production
H-Hbar production

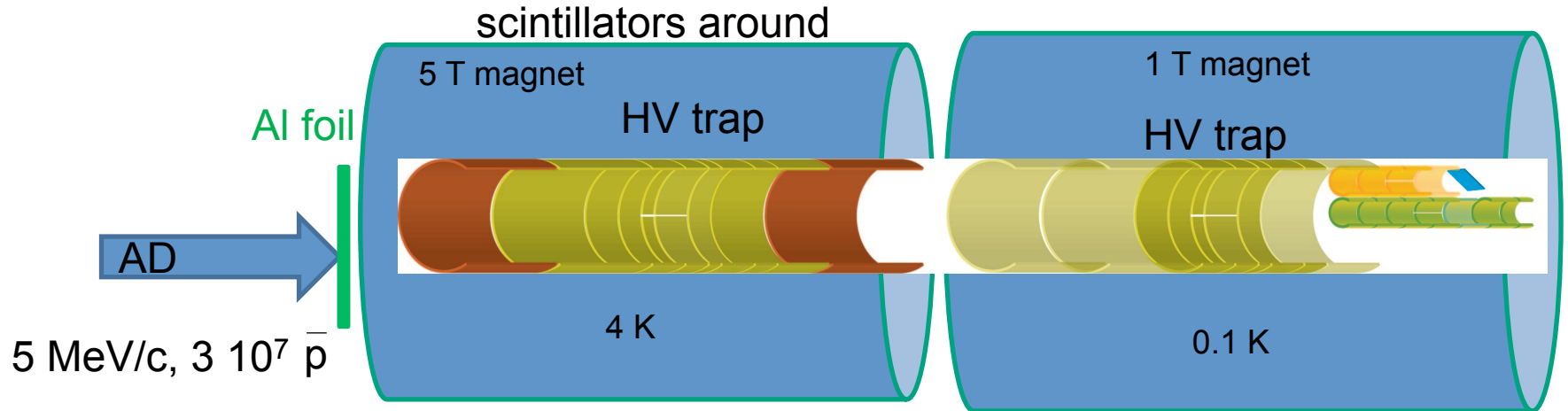


CERN AD area

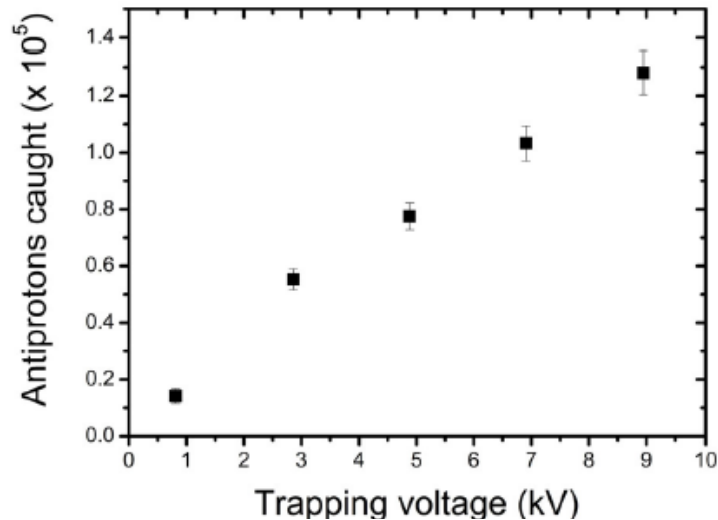
Trapping antiprotons



Antiproton trapping

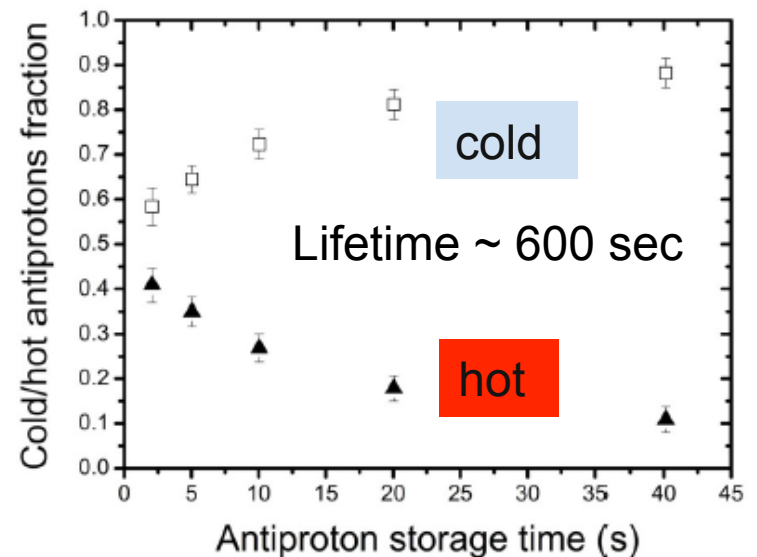


Anti-p catching after Al degrader in the 5T trap



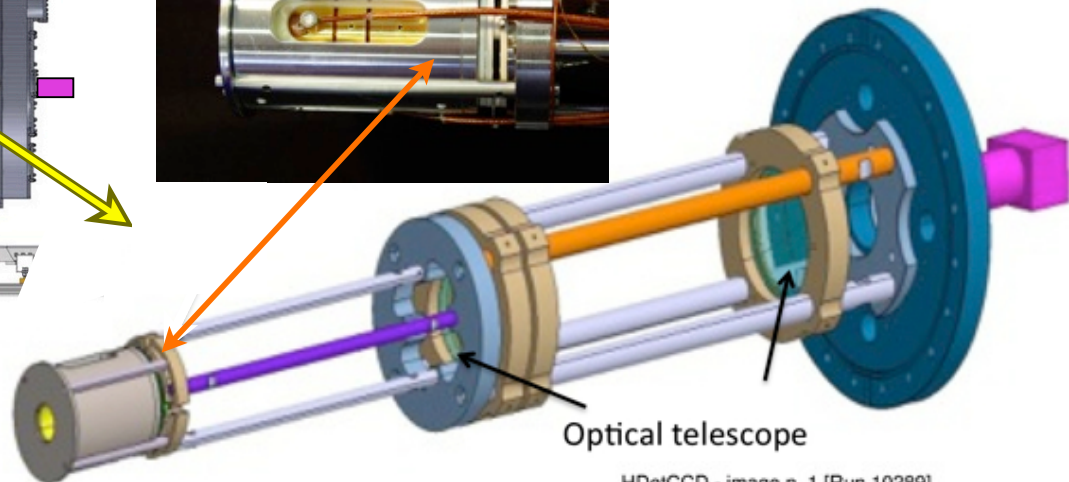
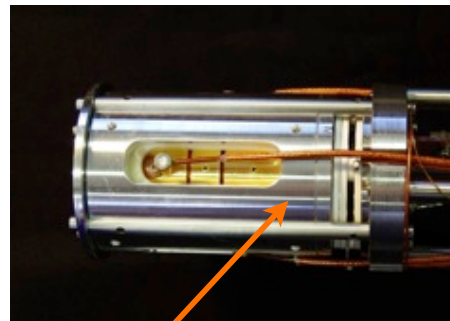
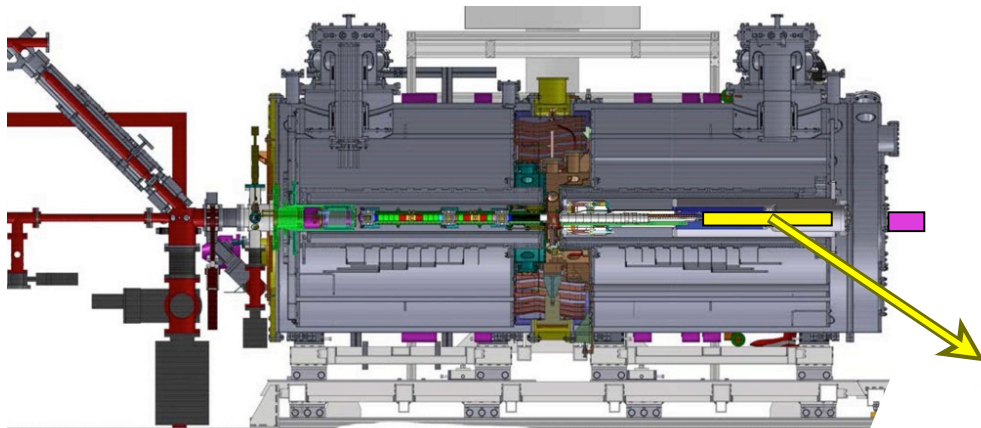
D. Krasnický *et al.*, AIP Conf. Proc. 1521, 144 (2013)

Cooling by electrons



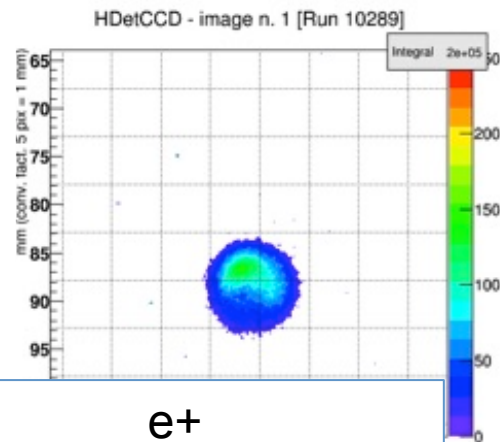
(anti)hydrogen detector

- Tune/Analyse traps



field-ionizing electrodes + MCP / Faraday cup + optics + CCD camera

Optical telescope



Commissioned (2014) with e-

pbar,

e+

mm (conv. fact. 5 pix = 1 mm)

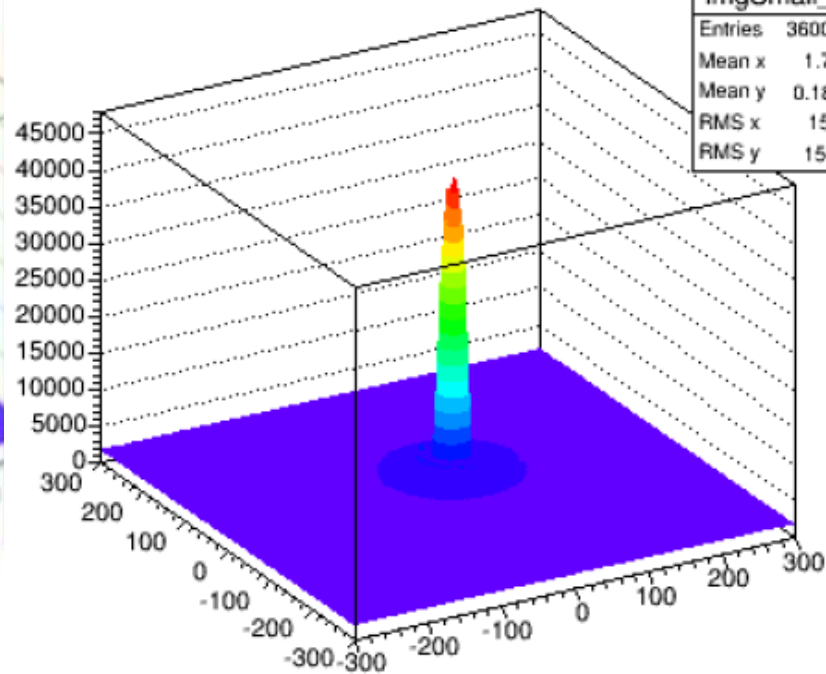
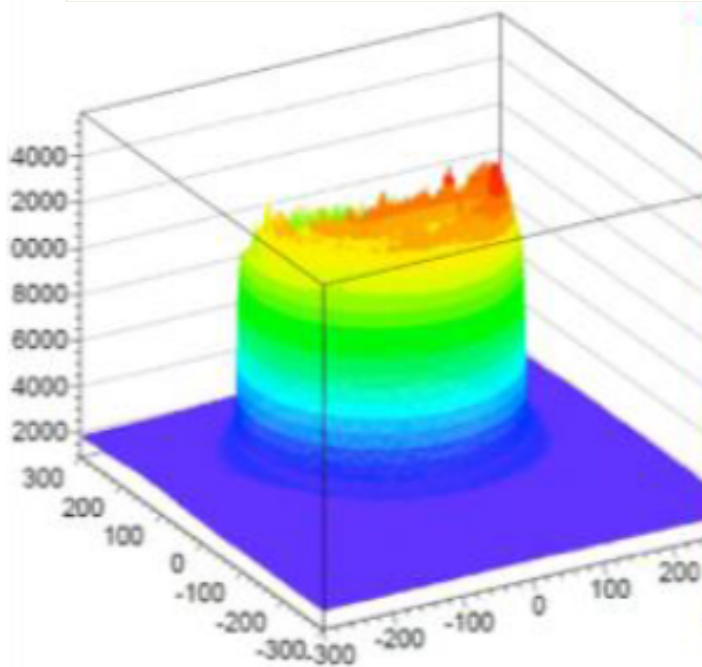
Radial compression of the Trapped plasma with RF field: Rotating Wall

Cold a

1

0.9

imgSmall_0	
Entries	360000
Mean x	1.715
Mean y	0.1813
RMS x	158.1
RMS y	157.9

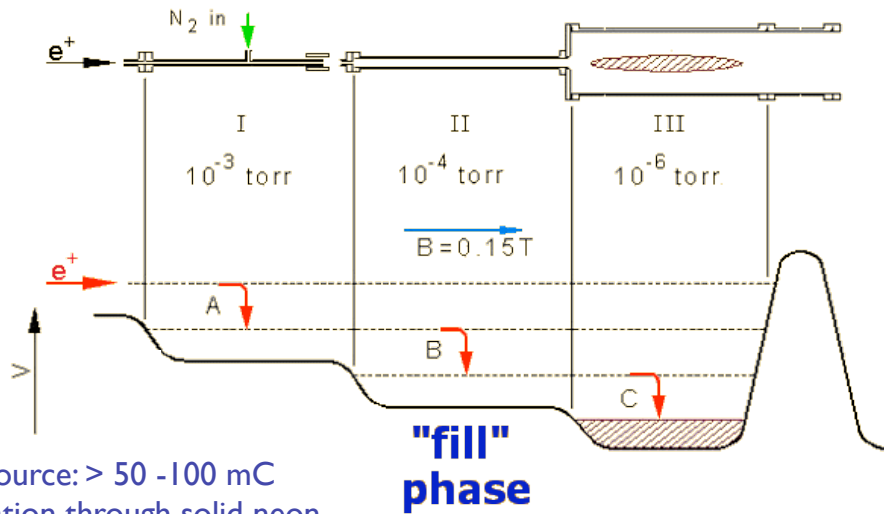
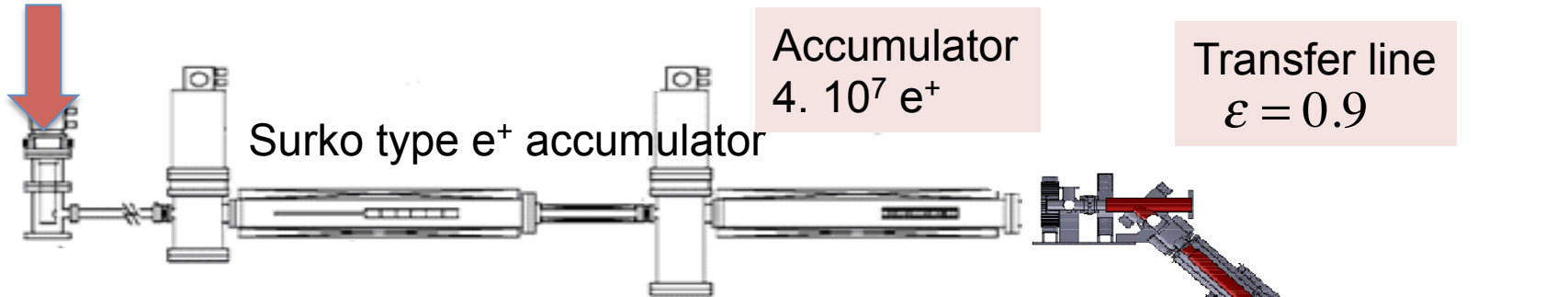


Positrons line

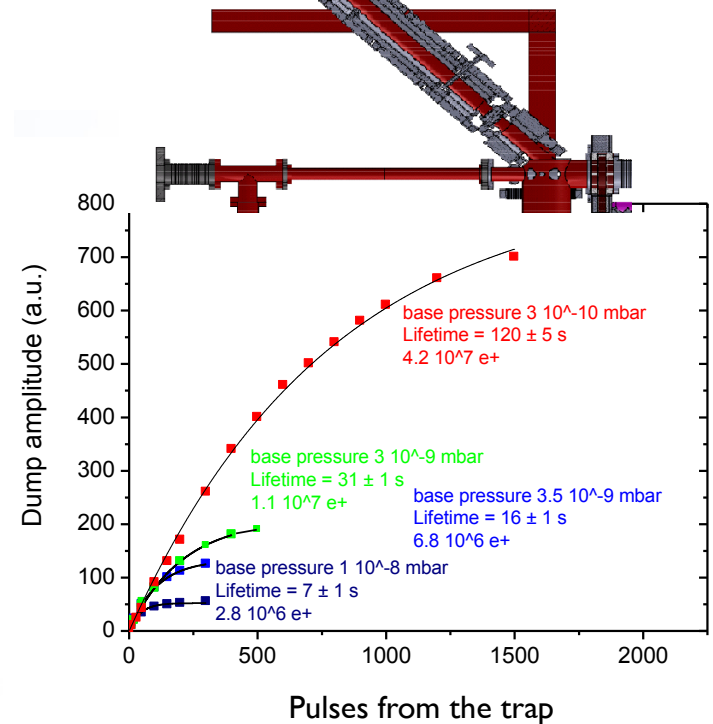
^{22}Na (14 mCi)

Solid Ne moderator: $\varepsilon = 2 \cdot 10^{-2}$

Trapping: $\varepsilon = 0.14$



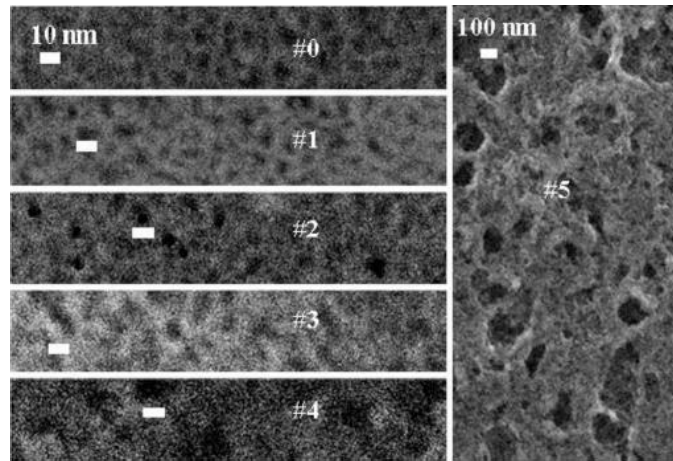
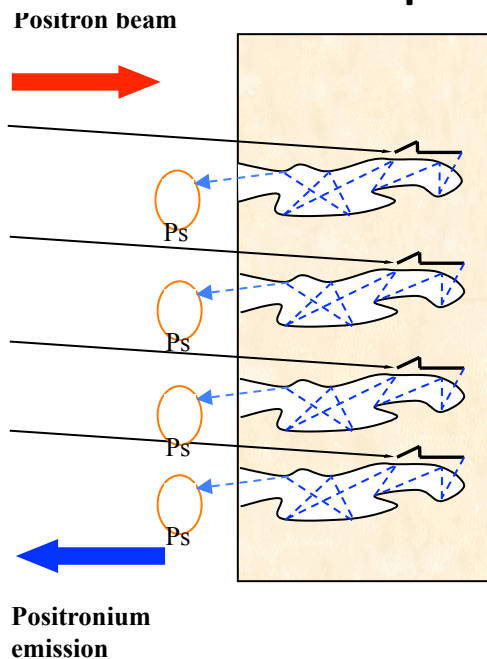
- ^{22}Na source: > 50 -100 mCi
- Moderation through solid neon
- Accumulation in trap
- Buffer gas cooling





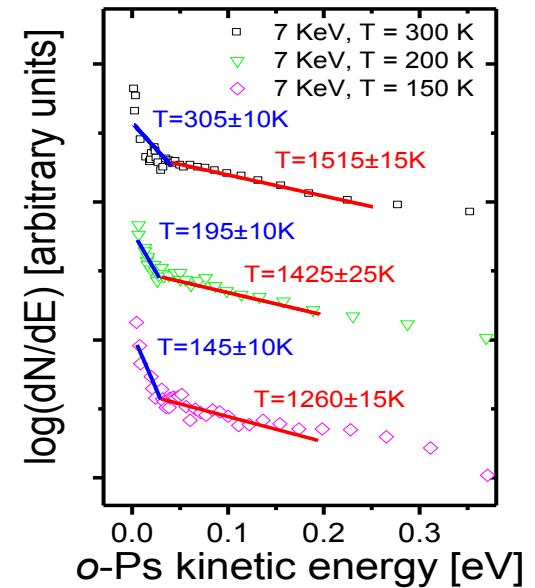
Positronium formation

- Implantation of e^+ in nano-porous target
 - SiO_2 - 8-14 nm pores; $T \sim 75$ K oPs
 - Tune pore size to tune Ps temp.



Nanochannels size

0 : 4 – 7 nm # 3 : 10 – 16 nm
1 : 8 – 12 nm # 4 : 14 – 20 nm
2 : 8 – 14 nm # 5 : 80 – 120 nm

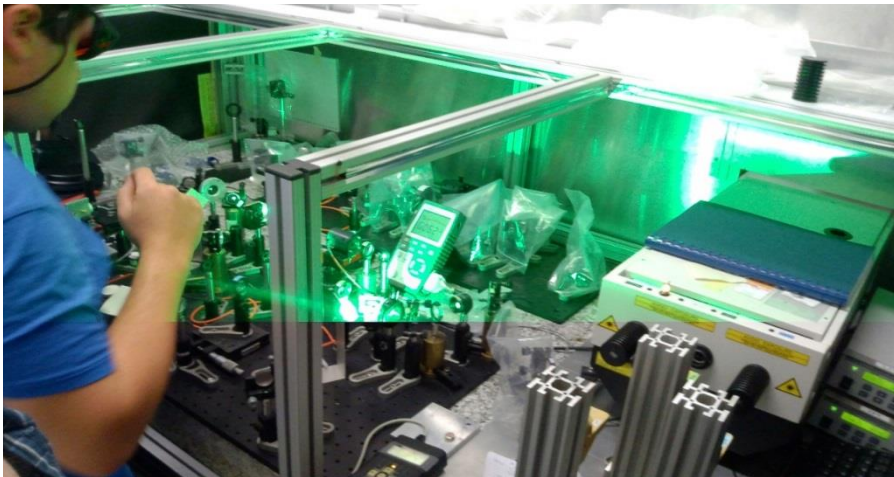
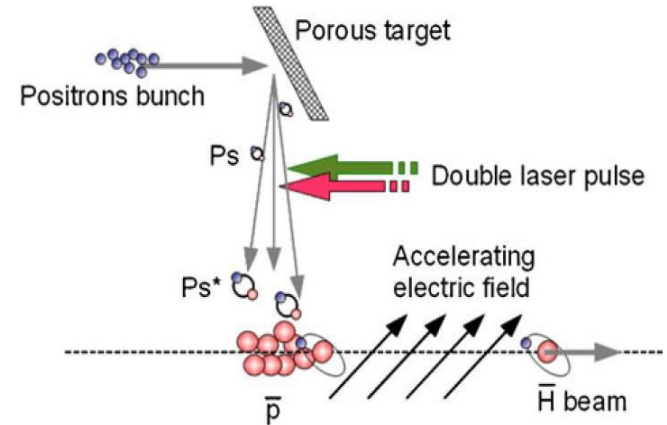


S. Mariuzzi et al., Phys. Rev. B 78, 085428 (2008)

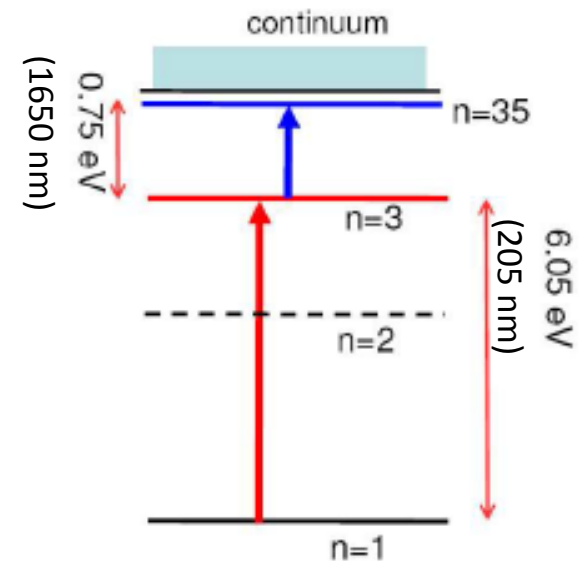
S. Mariuzzi et al., Phys. Rev. B 81, 235418 (2010)

Ps excitation

- Two stages excitation
 - UV (205 nm): $n=1 \rightarrow 3$
 - IR (1650-1700 nm): $n=3 \rightarrow 25-3$

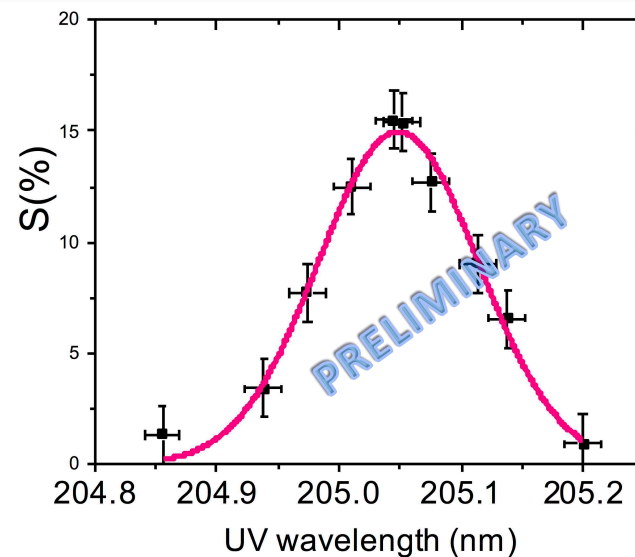
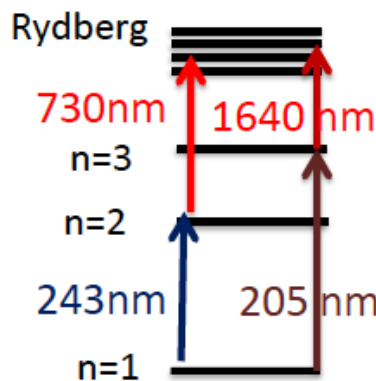
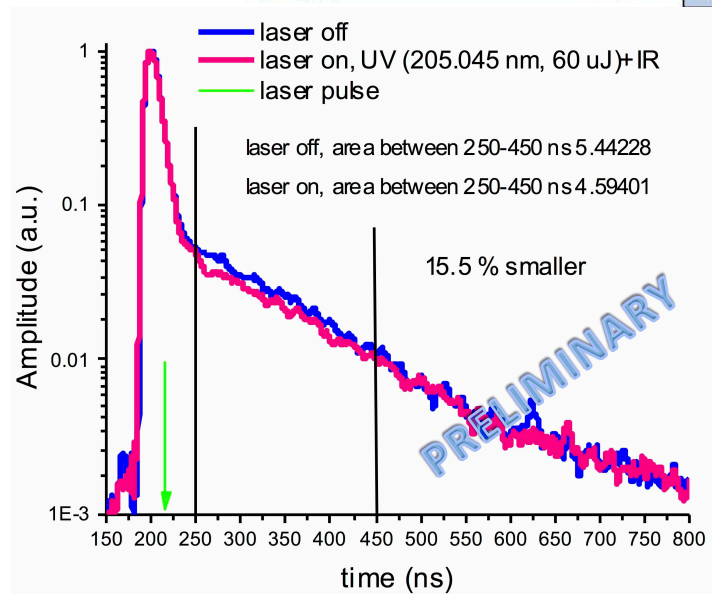
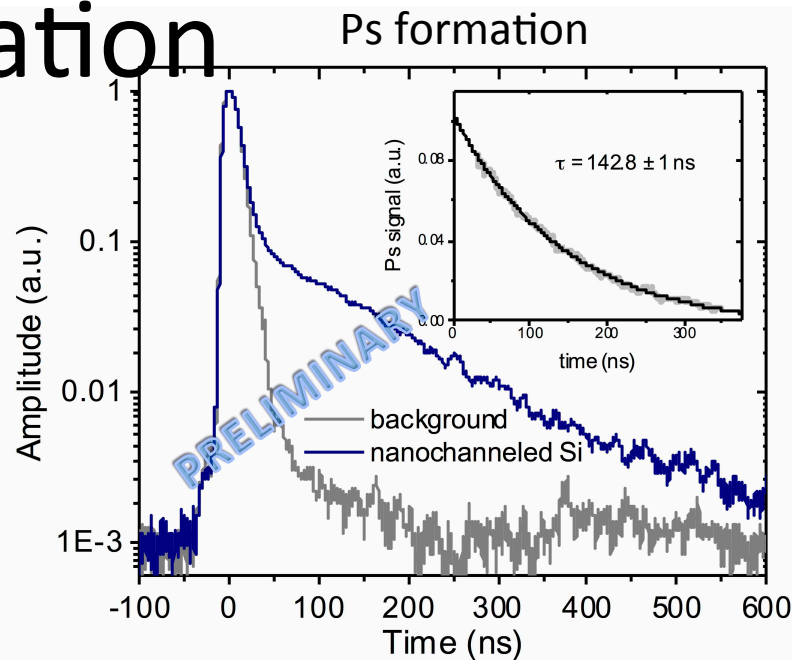
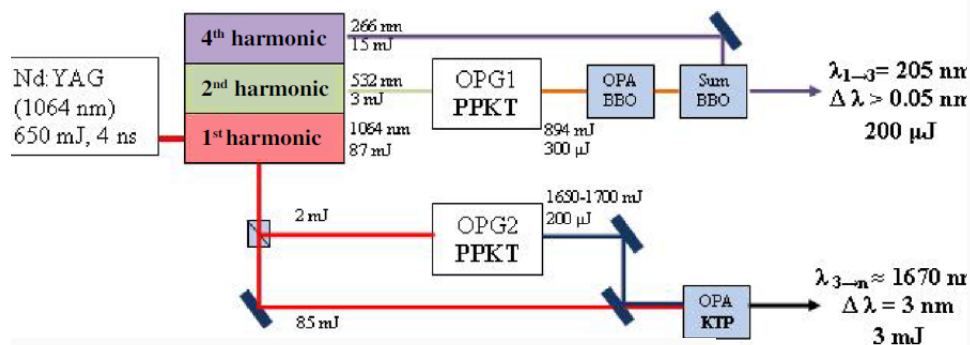


Phys. Rev. A 78 (2008) 052512; NIM B 269 (2011) 1527

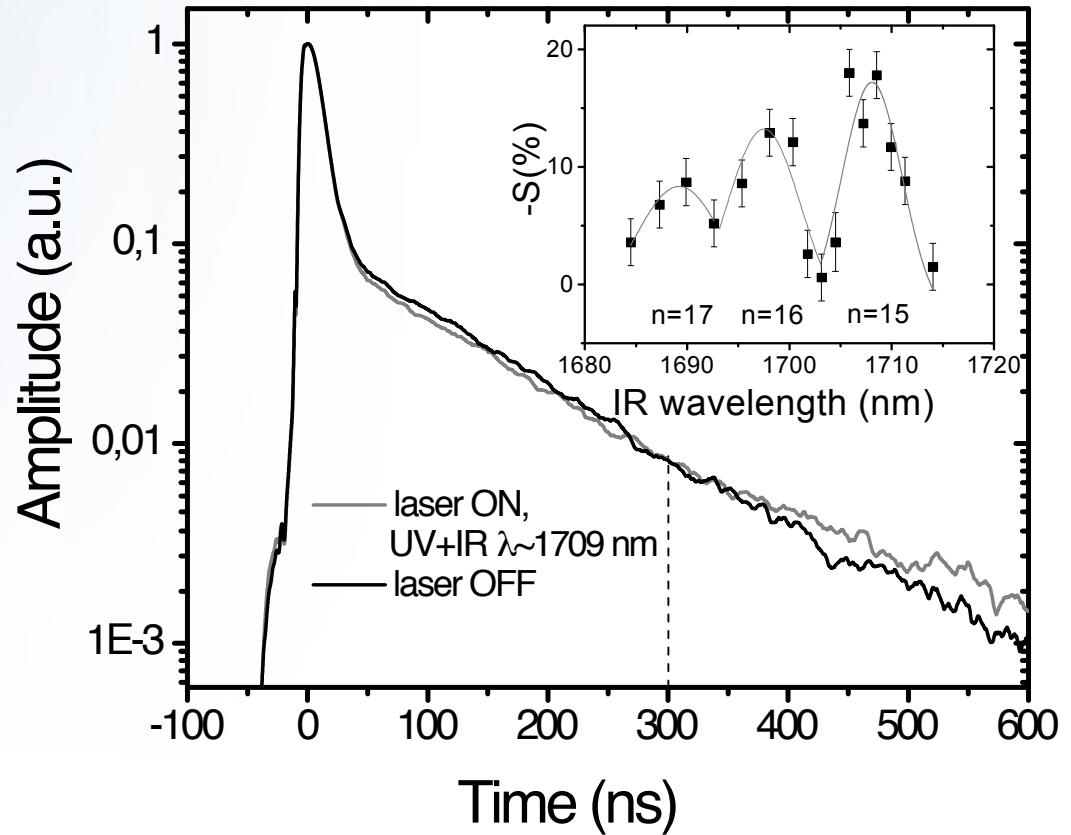
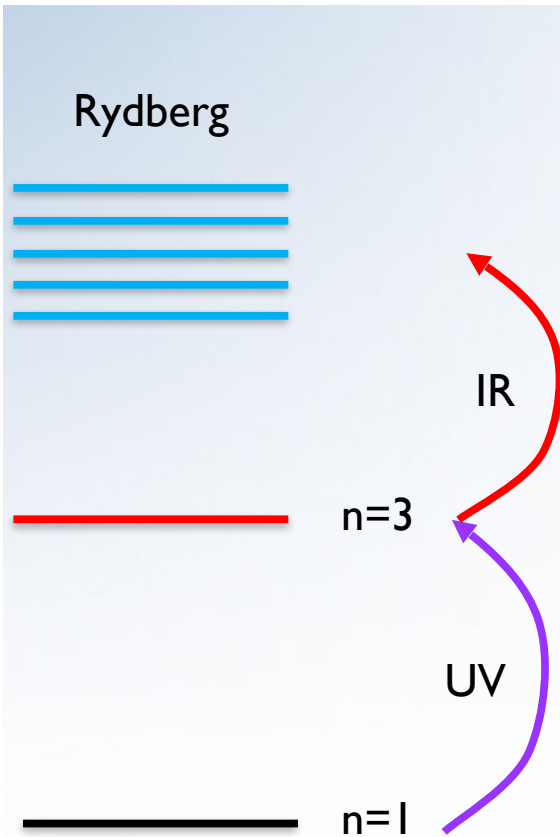


Ps excitation

- Test on Ps formation & excitation in a separate breadbox
- Laser excitation to $n=3$ demonstrated by photoionization measurements



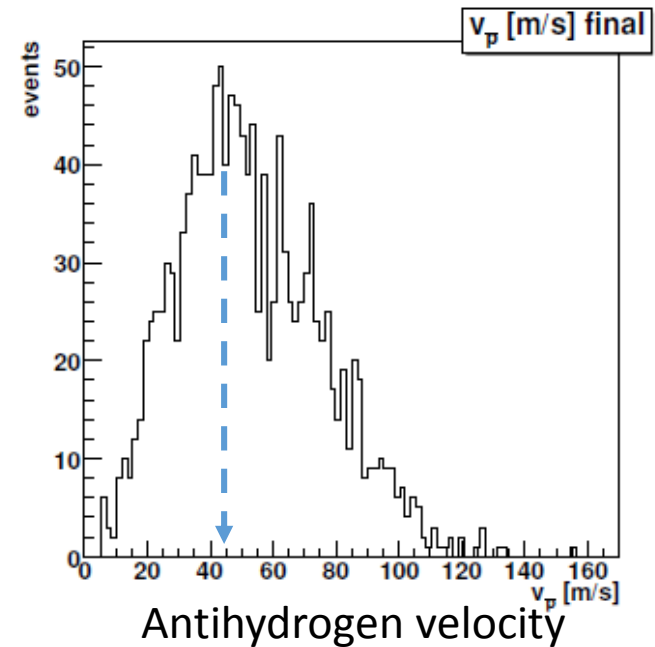
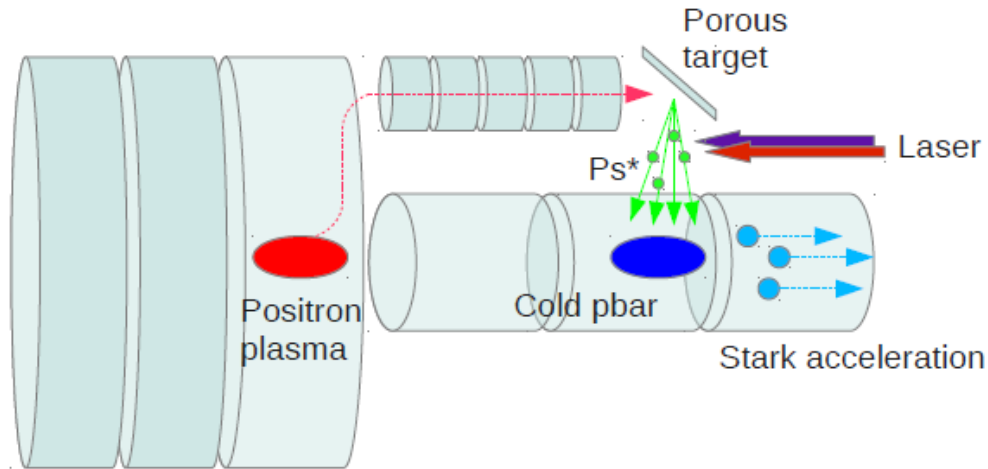
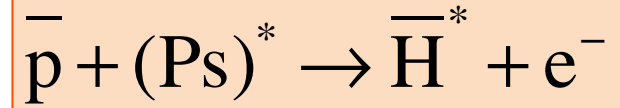
Rydberg excitation



$$S(\%) = (\text{Area laser OFF} - \text{Area laser ON}) / \text{Area laser OFF}$$

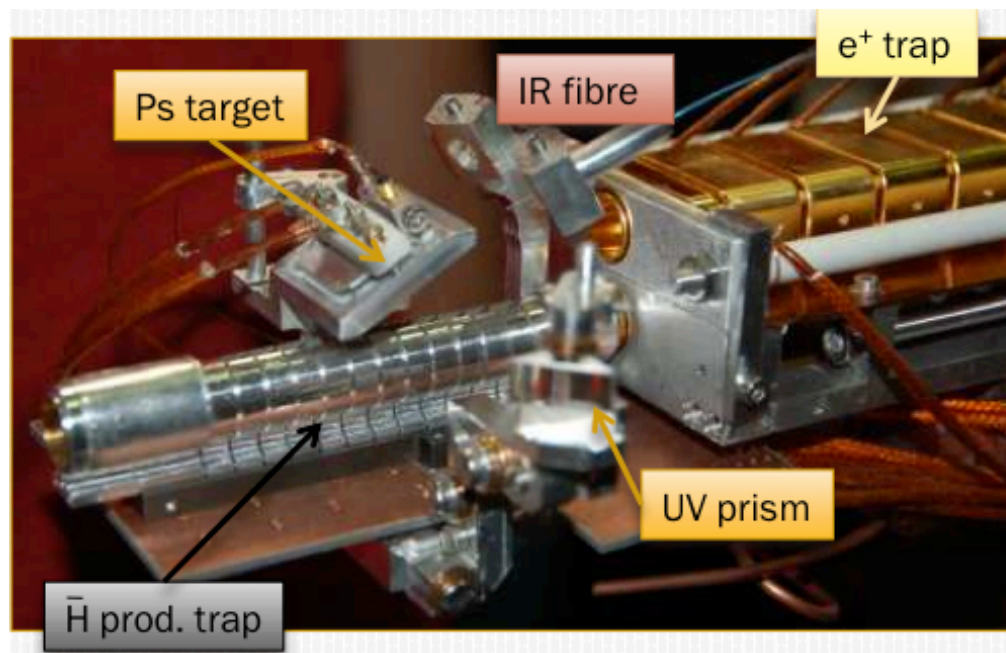
Hbar atoms Production

- Charge exchange reaction:

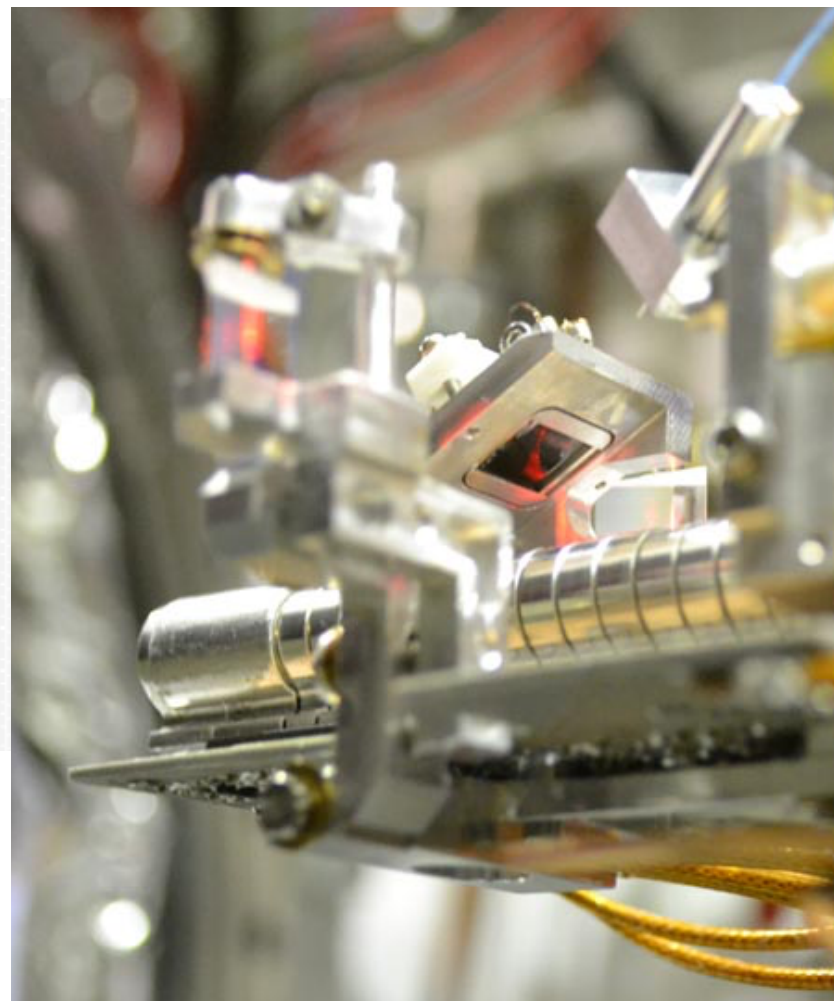


- large xsection: $a_0 n^4$
- $V(\text{Hbar})=45\text{m/s}$ @ $T=100\text{nK}$, $n=30$

Production zone

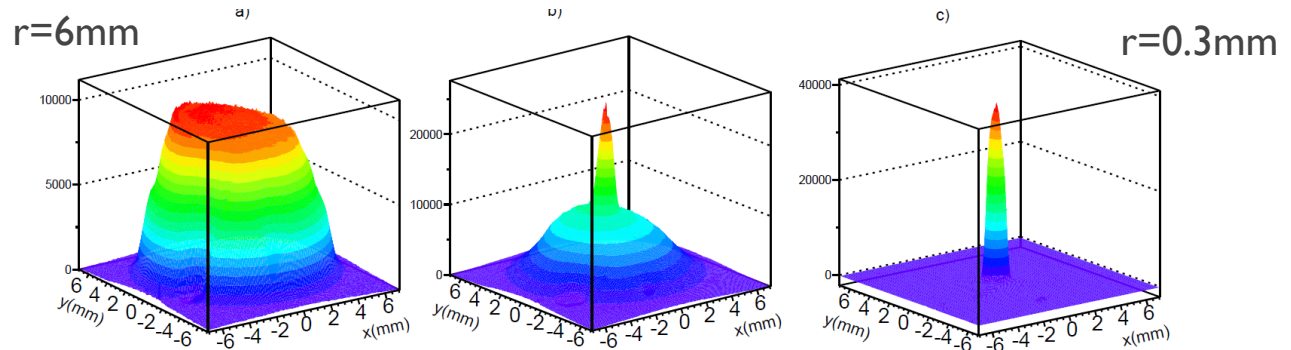


In the 1T region

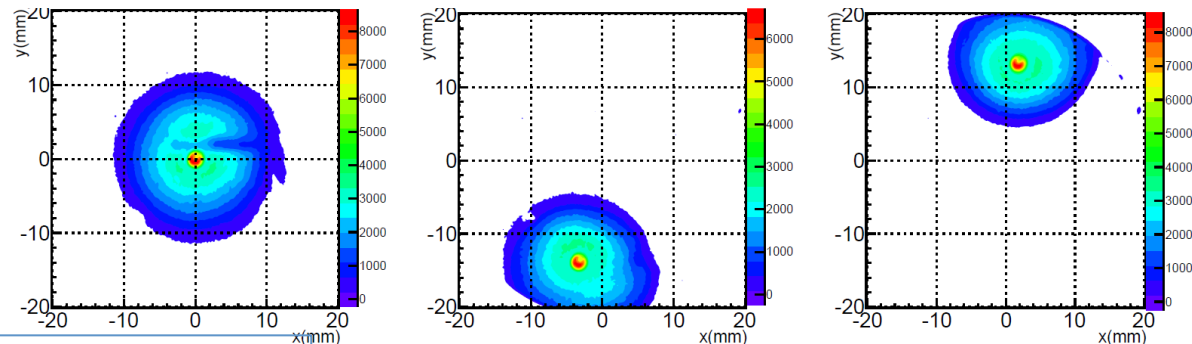


Plasma manipulations in traps

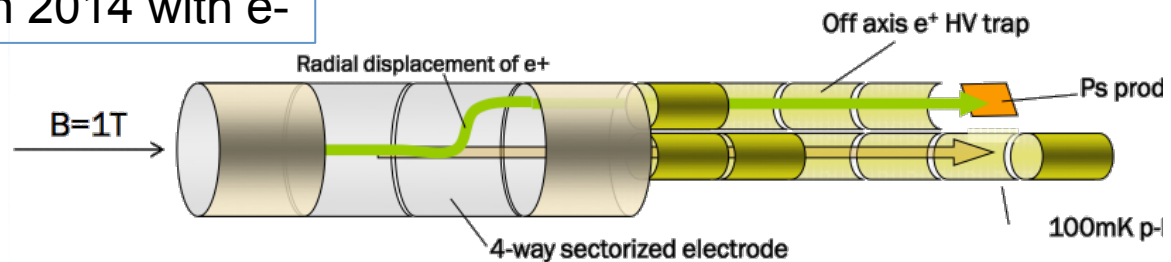
rotating wall



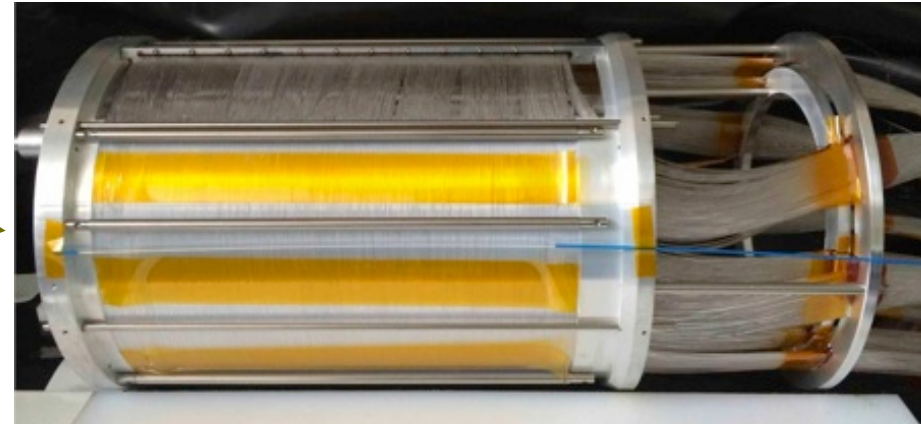
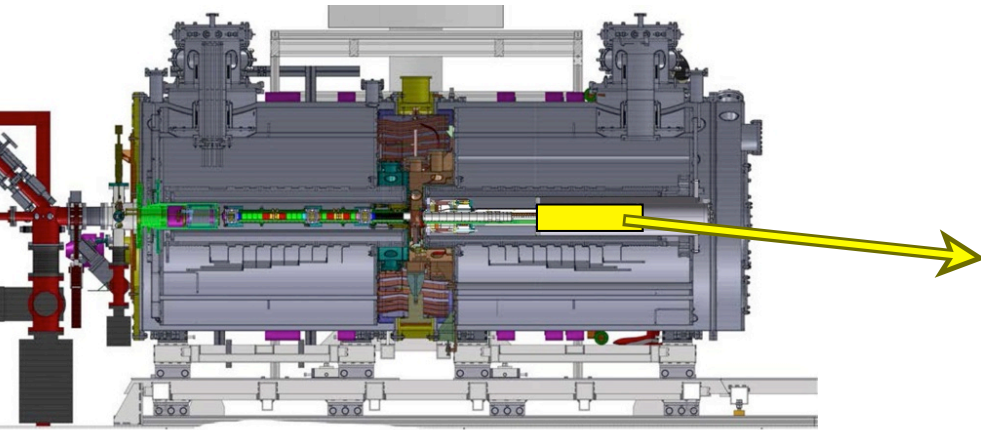
diocotron excitation



Tests performed in 2014 with e-



Anti-H monitoring



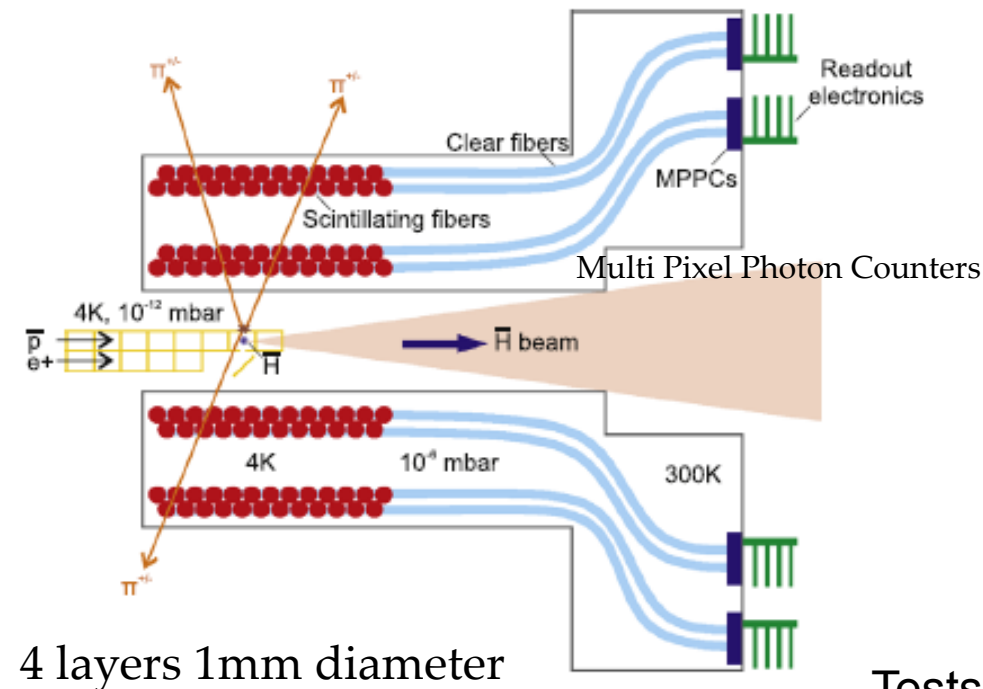
400 scintillating fibers

Constraints:

- operate at 4 K, in high vacuum
- inside 1 T
- power dissipation < 10 W.

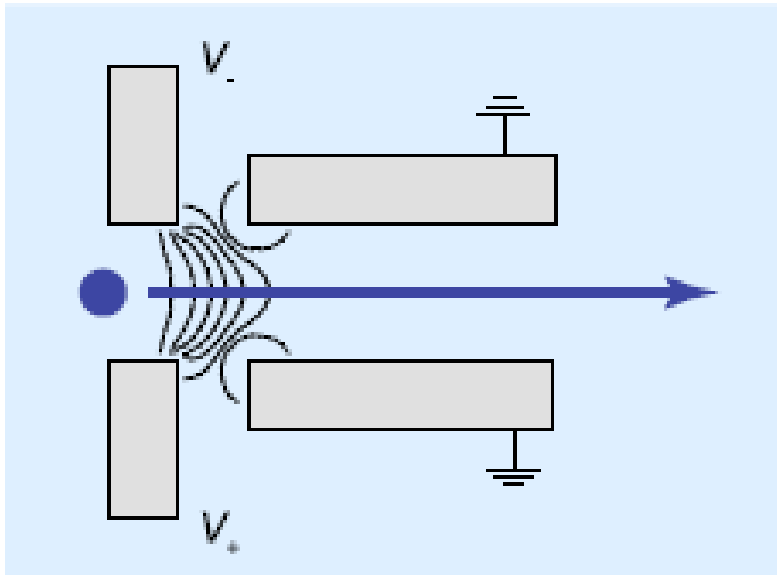
Fast Annihilation Particle Tracking
resolution ~ 2 mm

Tests with cosmic rays \rightarrow almost no degradation



Hbar beam

- Accelerate an electric dipole (H^*) by an electric field gradient
 - Hbar bea along z axis @ 100 m/s
- Demonstrated with hydrogen

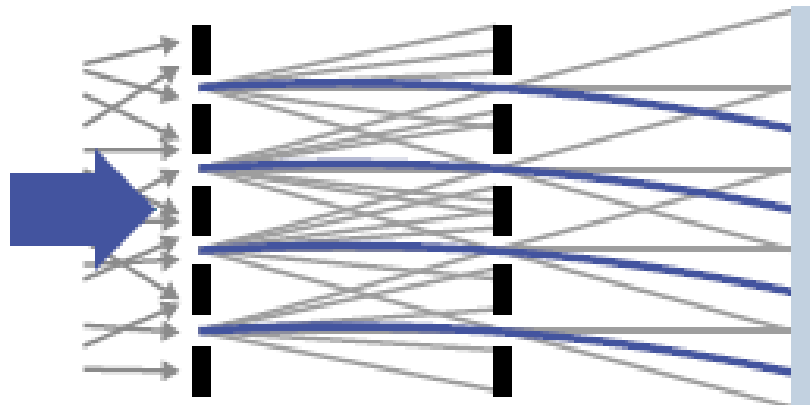
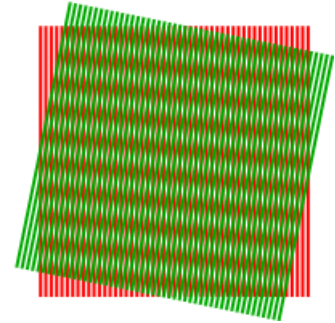


$$\vec{F} = -\frac{3}{2} e a_0 n (n-1) \nabla \vec{E}$$

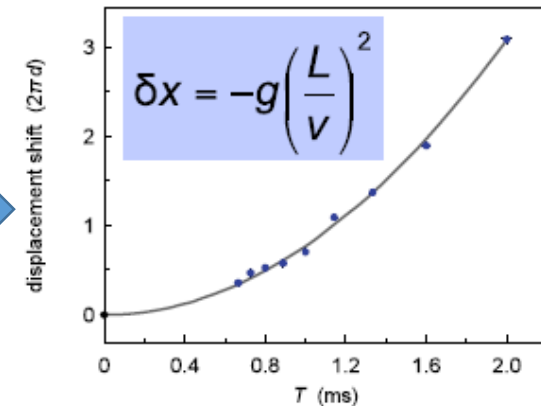
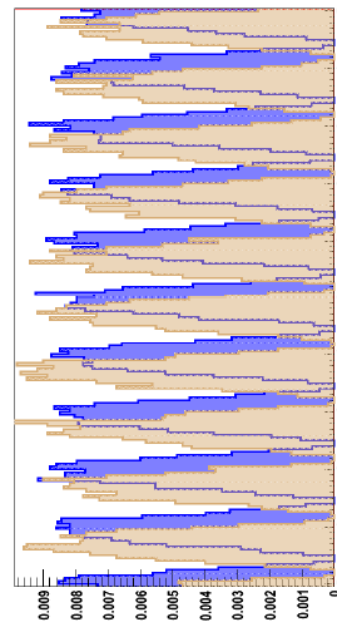
- $n = 22, 23, 24$
- Acceleration up to $2 \times 10^8 \text{ m/s}^2$ achieved

Gravity measurement

- Using a moiré deflectometer + position sensitive detector
- Detection a periodic shift (due to gravity)

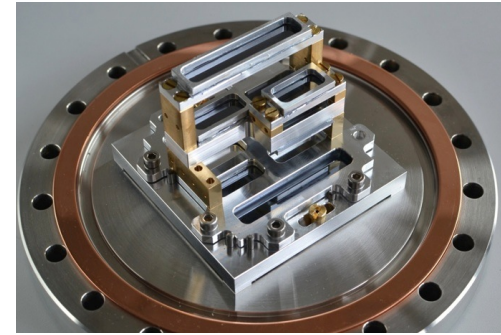


Moiré deflectometer + position detector

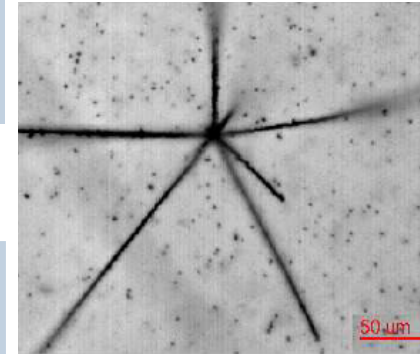
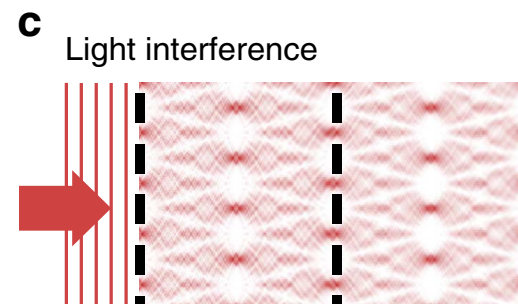
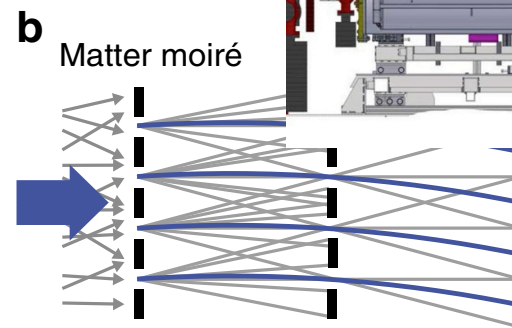
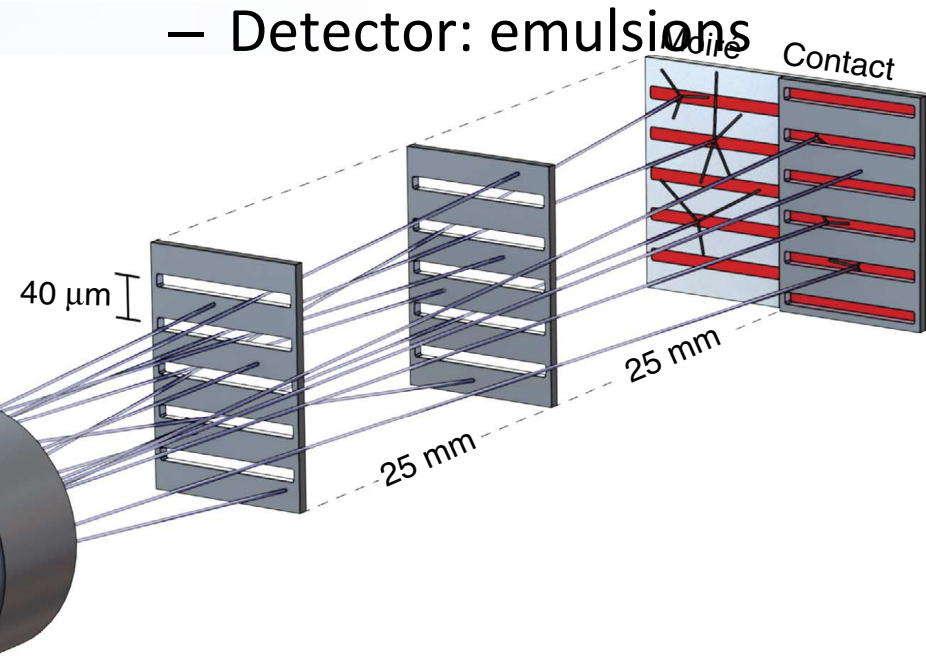
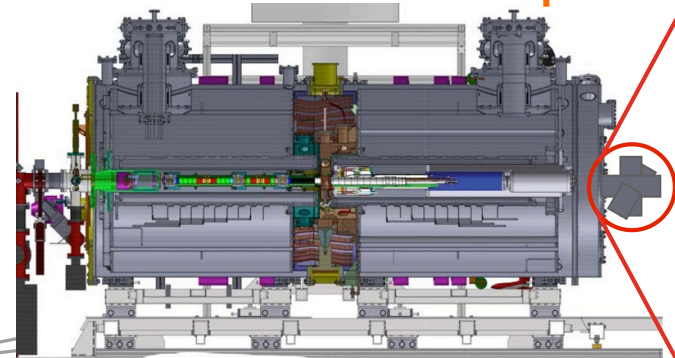


Moiré deflectometer

- Proof of principle tested with pbar
 - D=25 mm, pitch=40 μ m
 - Beam (pbar) 100-150 keV
 - Laser beam – reference
 - Detector: emulsions

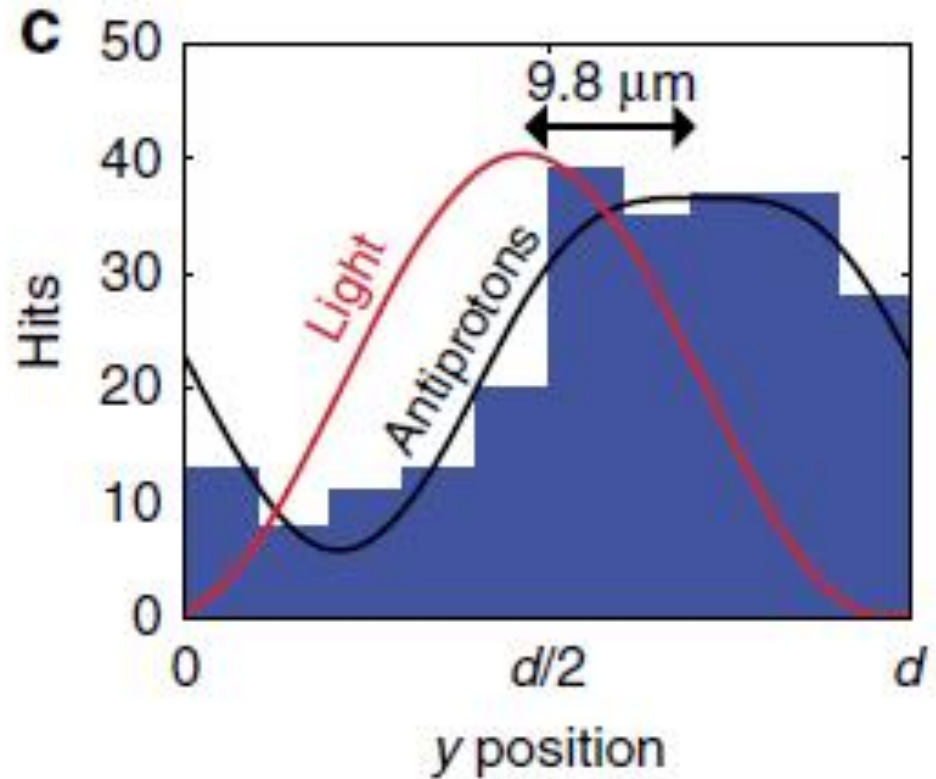
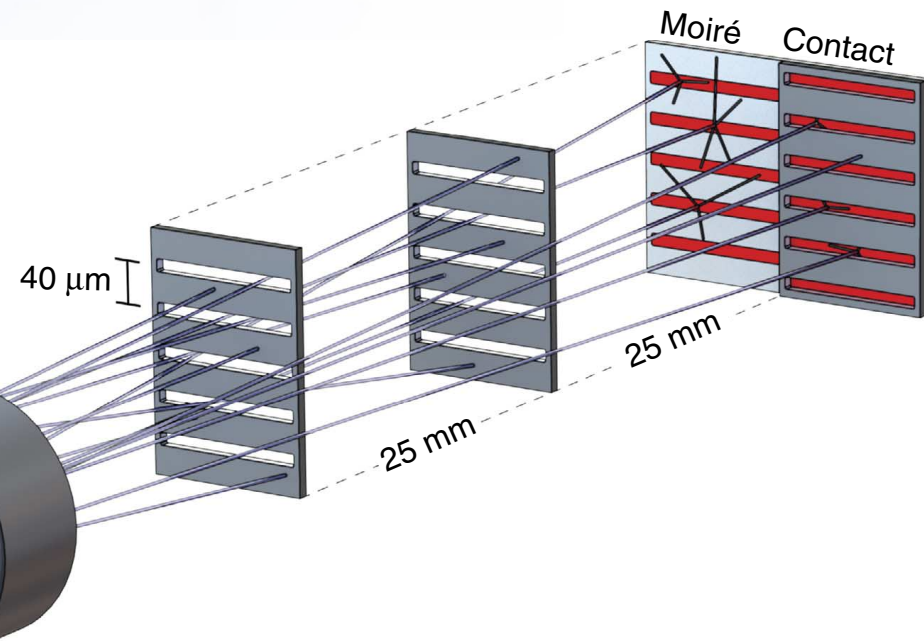


2012 setup



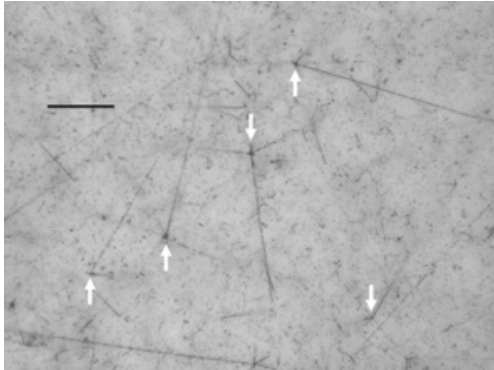
Moiré deflectometer

- First demonstration of a moiré deflectometer technique with pbar
 - Displacement $\sim 9.8 \mu\text{m}$
 - $F \sim 530 \text{ aN}$ ($B_{\text{eq}} \sim 10 \text{ G}$)

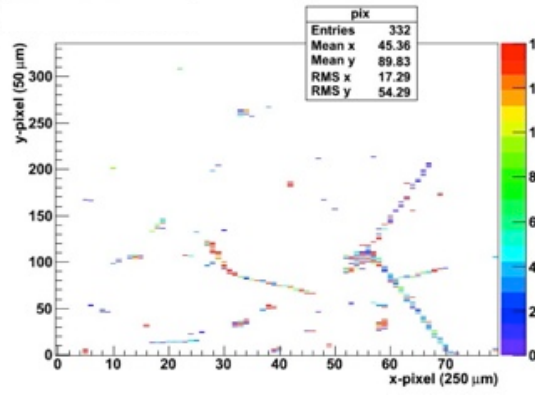


Detection

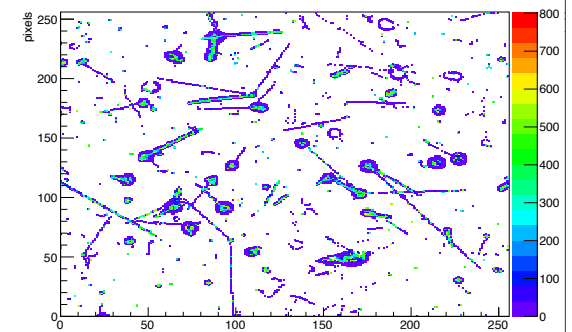
- We have tested (2012/2014) emulsions



Silicon detector

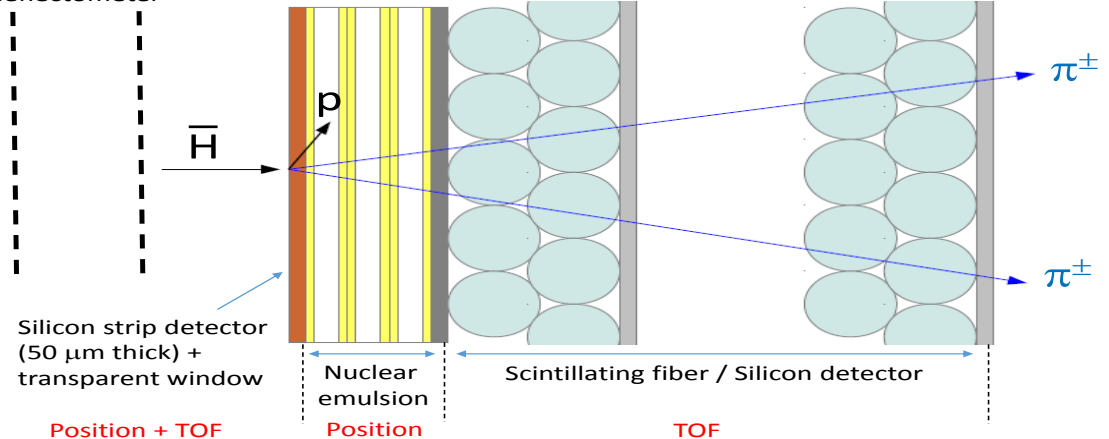


Timepix3



- Prepare hybrid (emulsion+Si) detector

Moiré deflectometer



AEGIS - current status

Achieved so far:

- **Antiprotons** (compression and fast transfer in 1 Tesla) + electron manipulation
- Positron transfer
- oPs, $n=1$; $n=3$; $n>15$ (Rydberg)
- H acceleration (Stark)
- Silicon sensor and emulsion detector tests
- Moiré deflectometer and time resolution tests

To be done (in situ): 2015/16

- The first goal is to produce pulsed antihydrogen formation at 4 K
- Ps* spectroscopy

Further improvements (in situ & test labs): 2016/2017

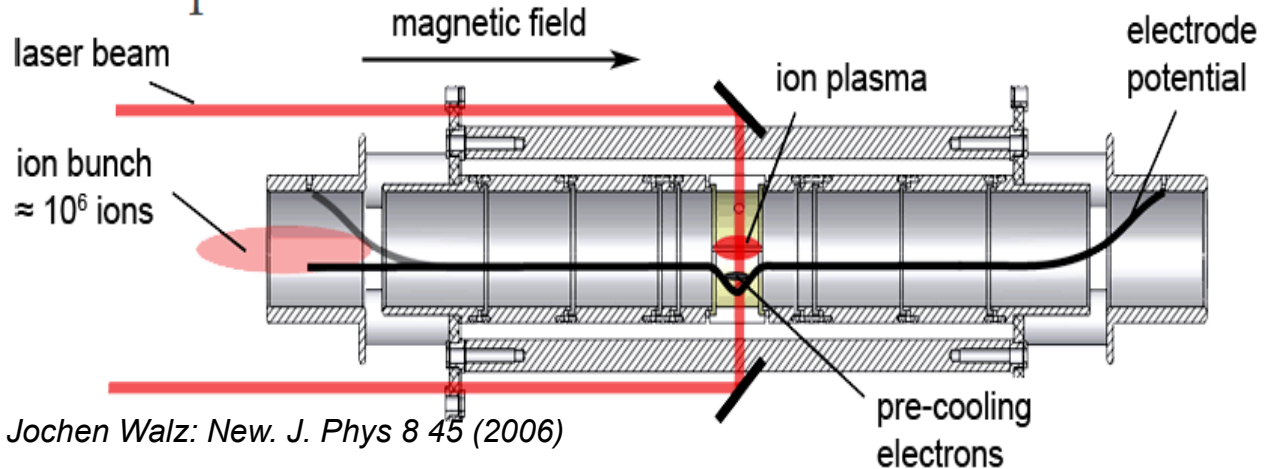
- Ps formation, anti-H beam formation,
- Gravity setup – **First measurement**
- Increase e^+ (new source)
- Improve pbar cooling (C_2^- cooling)
- New pbar beam (ELENA)

AEGIS

Colder pbar

- Cold pbar to maximize flux
 - $T \sim 100$ mK (7K now)
- Sympathetic laser cooling with negative ions

Experimental procedure:



Alban Kellerbauer & Jochen Walz: New. J. Phys 8 45 (2006)

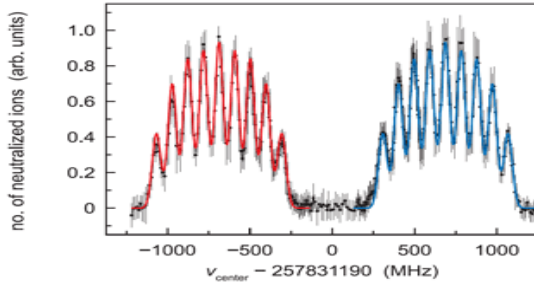
1. Capture of anions in the Penning trap
2. Pre-cooling of anions with electrons
3. Laser cooling of anions
4. Simultaneous confinement of antiprotons in the trap
5. Sympathetic cooling of antiprotons by anions

Atomic anions ?

One bound state lying below the neutral

IA	IIA	IIIB	IVB	VB	VIB	VII B	VIII B	IB	IIB	IIIA	IVA	VA	VIA	VIIA	VIIIA			
H															He			
Li	Be									B	C	N	O	F	Ne			
Na	Mg									Al	Si	P	S	Cl	Ar			
K	Ca		Sc	Ti	V	Cr	Mn	Fe	Co	Ni	Cu	Zn	Ga	Ge	As	Se	Br	Kr
Rb	Sr		Y	Zr	Nb	Mo	Tc	Ru	Rh	Pd	Ag	Cd	In	Sn	Sb	Te	I	Xe
Cs	Ba	La-	Lu ^c	Hf	Ta	W	Re	Os	Ir	Pt	Au	Hg	Tl	Pb	Bi	Po	At	Rn
		Yb																
Fr	Ra	Ac-	Lr ^c	Rf	Db	Sg	Bh	Hs	Mt	Ds	Rg	Cn	Uut	Fl	Uup	Lv	Uus	Uuo
		No																

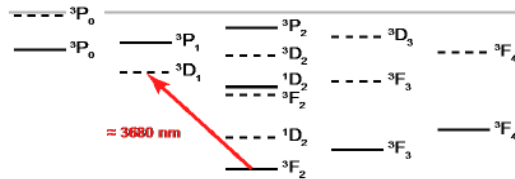
Os⁻



A. Kellerbauer et al., PRA 89 (2014) 043430

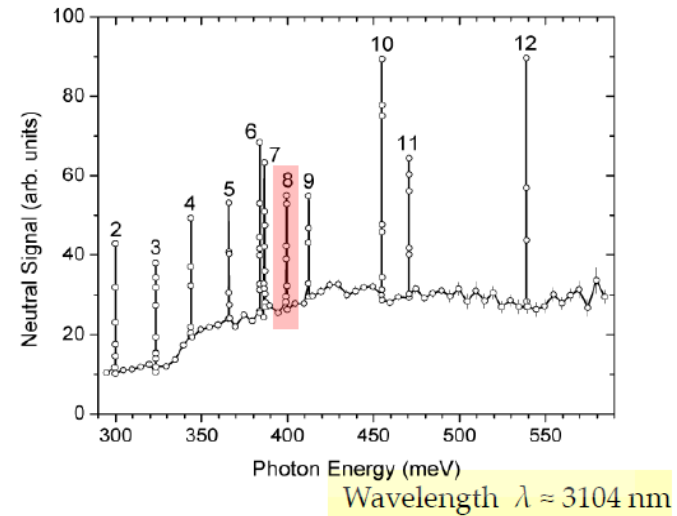
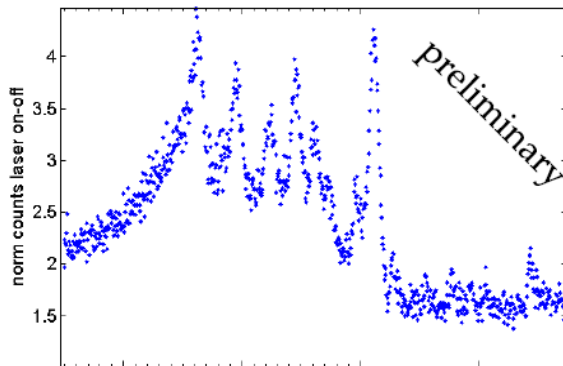


La⁻



La⁻ $5d^1 6s^2$ Transition rate $\Gamma \approx 4.5$ kHz
 $5d 6s^2 6p$ 99.98% closed transition

- First La⁻ resonance recorded at MPIK:



[S. M. O'Malley & D. R. Beck,
Phys. Rev. A 81 (2010) 032503]

[C. W. Walter et al.,
Phys. Rev. Lett. 113 (2014) 063001]

Difficult to produce + Low cooling rate + spectroscopy unknown + photodetachment

Why not molecules!



2010: DeMille's group demonstrates laser collimation with a radiative force on **SrF**
 2014: Show 3D cooling and trapping: *Nature* 512, 286–289 (21 August 2014)

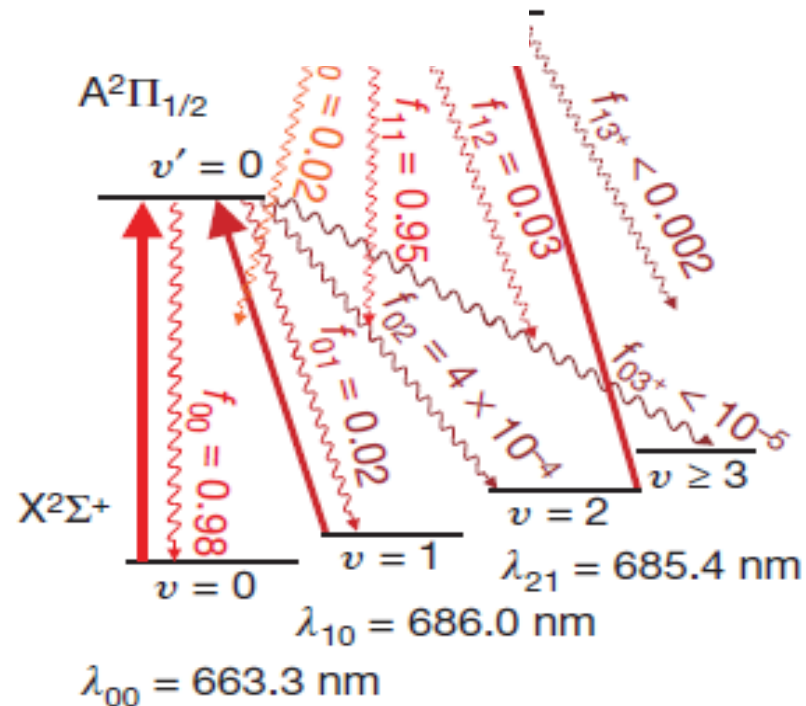
LETTER

doi:10.1038/nature09443

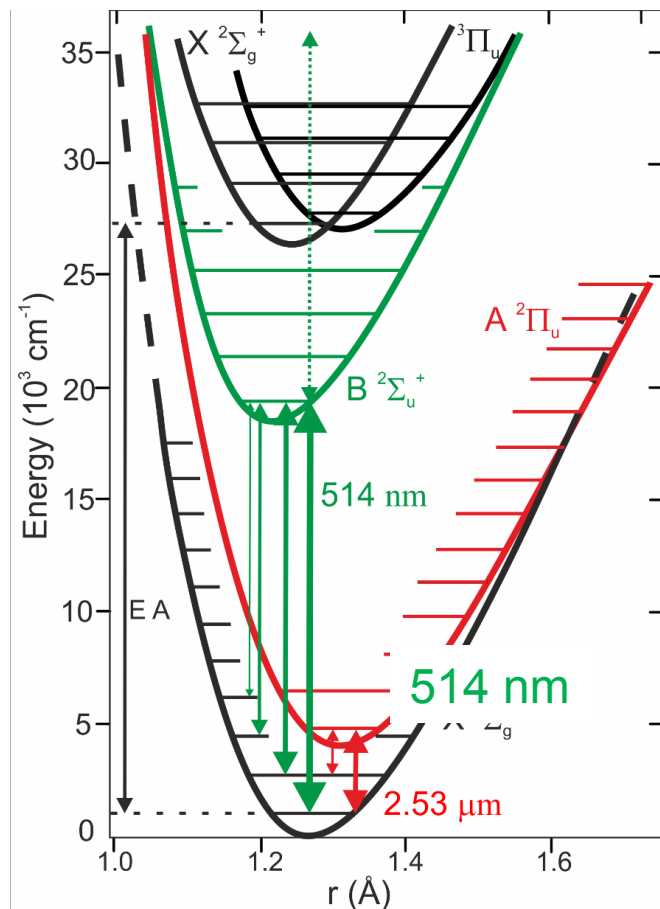
Laser cooling of a diatomic molecule

E. S. Shuman¹, J. F. Barry¹ & D. DeMille¹

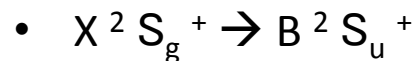
Like a 2 level atom: 98% of branching ratio



Example (best known molecule) C_2^-



2 possible cooling schemes



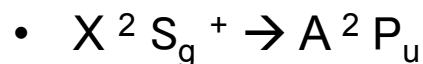
541, 598, 667, 753 nm

probability (Franck-Condon) 72, 23, 3, 0.8

Ultra fast cooling : lifetime 70ns

high photodetachment $s \sim 10^{-17} \text{ cm}^2$

\rightarrow rate $I s / h n = 1/s$ for $I = 30 \text{ mW/cm}^2$



2.53 μm and 4.50 μm

10 lasers

10 K \rightarrow 100mK in 100 ms

30% losses (photodetachment)

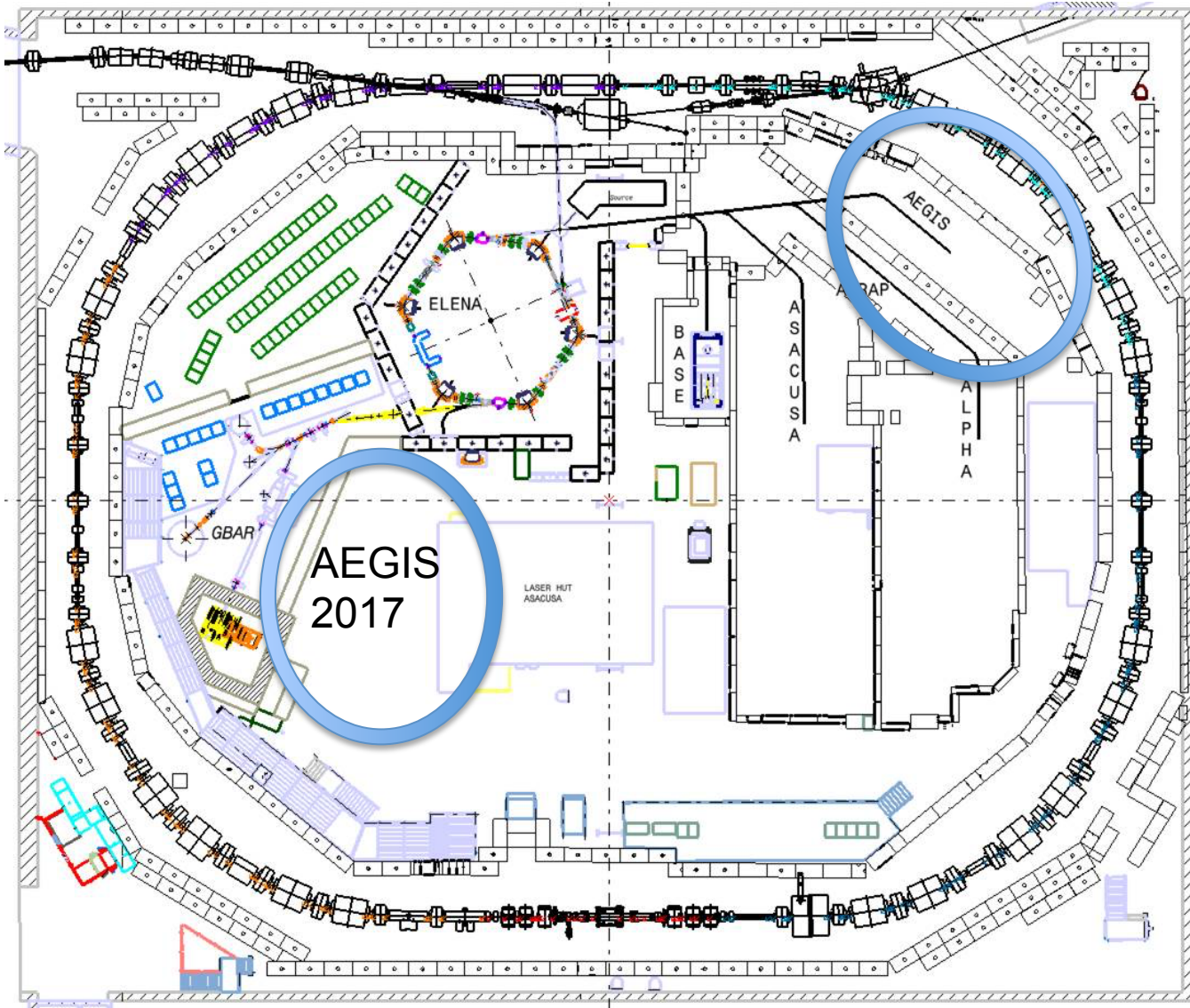
Less laser needed

slower cooling (100 μs)

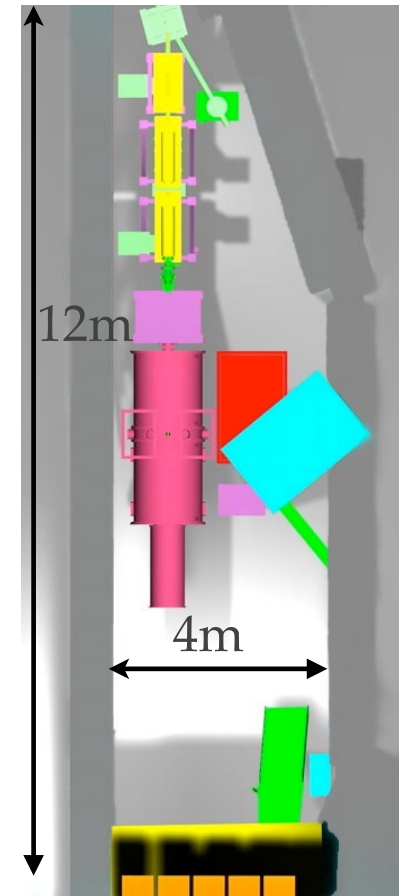
No photodetachment

Easy to produce + Fast cooling rate + spectroscopy known + no-photodetachment

Extremely Low Energy Antiproton area



Pbar 100 keV
En //



Matter wave interferometer and deflectometer

$$\delta\varphi_{\text{beam}} < \varphi_{\text{diffr}} = \lambda_{\text{dB}} / d$$

$$v = 1000 \text{ m/s}$$

$$d = 0.1 \text{ } \mu\text{m}$$

$$\partial\varphi_{\text{beam}} < 4 \text{ mrad}$$

$$T < 1 \text{ mK}$$

where d – grating period

v – beam velocity

← Currently not feasible

- diffraction requirement on angular beam divergence:

Mach–Zehnder interferometer:

$$\lambda_{\text{dB}} > d^2 / L \quad (\text{far-field diffraction})$$

Talbot–Laue interferometer:

$$\lambda_{\text{dB}} = d^2 / L \quad (\text{near-field diffraction})$$

Moiré deflectometer:

$$\lambda_{\text{dB}} \ll d^2 / L \quad (\text{no diffraction})$$

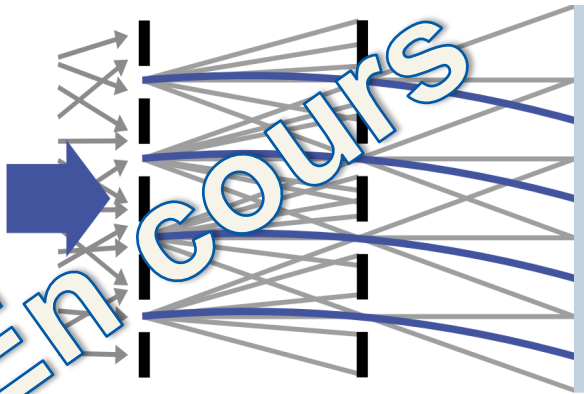
AEGIS



Interferences quantiques

Proton interferometry – outlook

classic regime

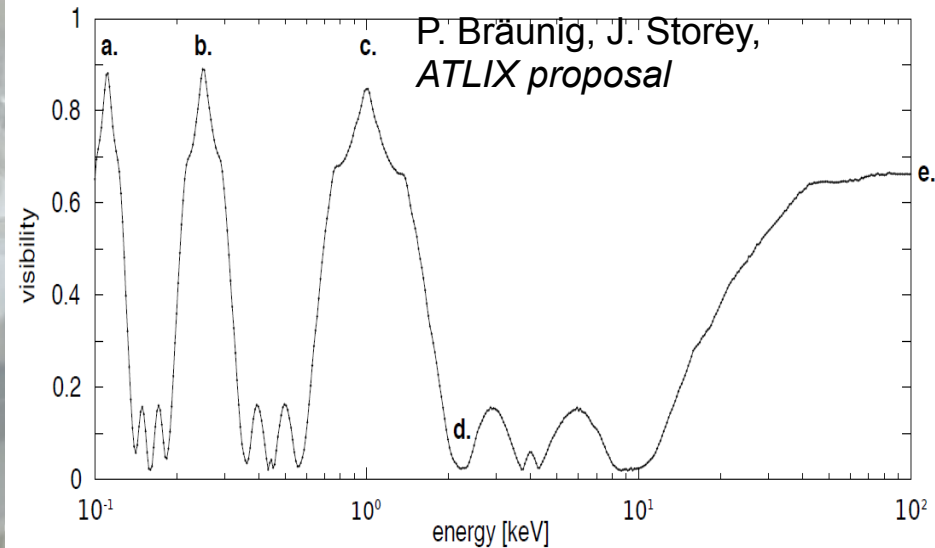
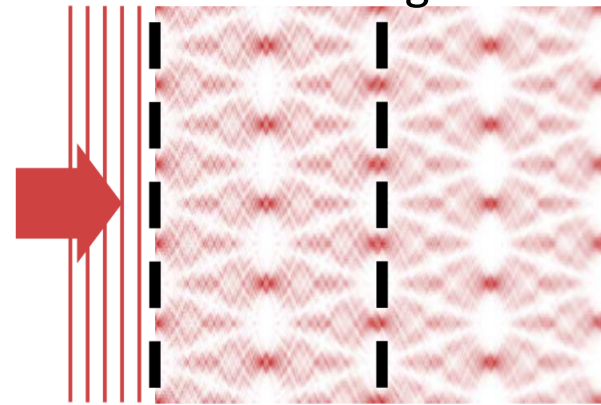


Talbot-length:

$$L_T = \frac{d^2}{\lambda}$$

visibility is the signal!

wave regime



Merci