

An instrumented absorber detector for hadronic interaction measurements

“The Smart Absorber of Crysbeam project”

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WaveCatcher and SAMPIC International Workshop, Feb. 7-8, LAL Oray, France

The Smart Absorber



64Ch WaveCatcher
readout



Multichannel
Power Supply
System



The Smart Absorber is a Cherenkov calorimeter optimized for study of hadronic shower evolution and for measurement of the hadronic cross sections at a high energy.

It is one of tree key elements of CRYSBEM Project



Outline

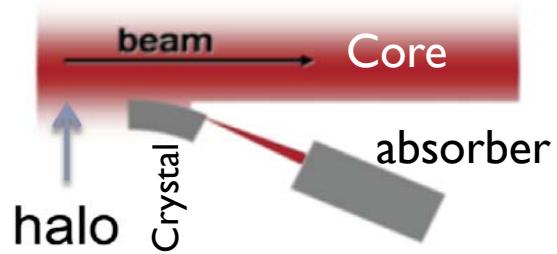


- ▶ The Crysbeam project
- ▶ The Smart Absorber
- ▶ The use of WaveCather : pro, cons and possible improvements

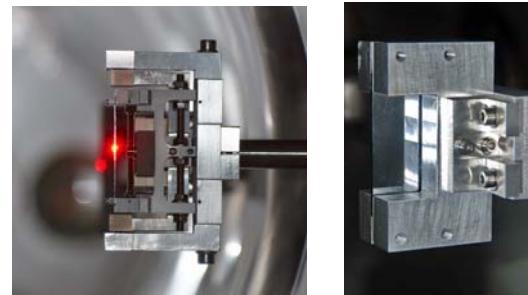


Crysbeam

CRYSB basic idea



Experience of crystals with accelerator is well consolidated at CERN with UA9



UA9 crystals used for Beam Collimation

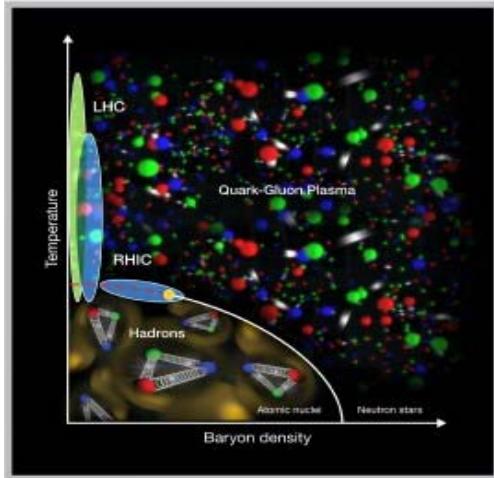
CRYSB (and INFN) propose to demonstrate **multi-TeV crystal extraction** is feasible

Goal → Efficient crystal extraction of a multi-TeV hadron beam, **with a bent crystal in channeling orientation**, for fixed target experiments

Cutting-edge technology in crystal and particle detectors

Physics with a multi-TeV hadron beam

Phase diagram of hadronic matter



- ▶ QCD at **unprecedented** laboratory energies and momentum transfers
- ▶ Proton spin physics
- ▶ **Quark-gluon plasma excitation in the target rest frame**
- ▶ Diffractive physics
- ▶ ... and more with secondary beams

"Physics opportunities of a fixed-target experiment using LHC beams"

S. J. Brodsky, F. Fleuret, C. Hadjidakis, and J. P. Lansberg, Phys. Rep. 522 (2013) 239-255.

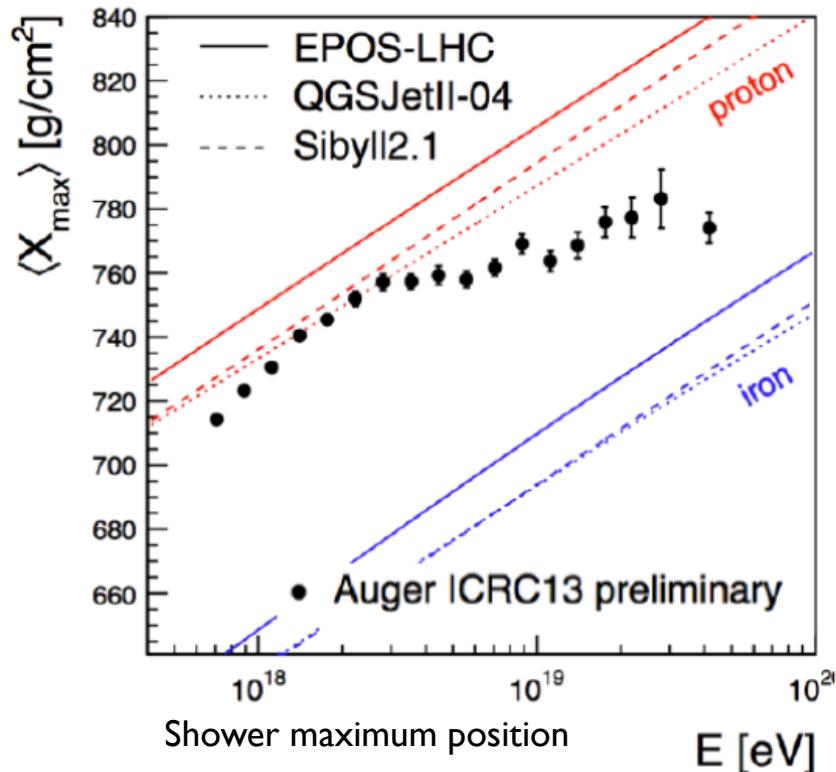
Cosmic ray shower



- ▶ What is the nature of the cosmic rays?
- CRYSEBAM** proposes experiments relevant for Cosmic Rays physics to demonstrate this technique is valid
- Study interaction of hadrons with **different nuclei (targets of C,N,O,...)**

Cosmic Ray Mass composition

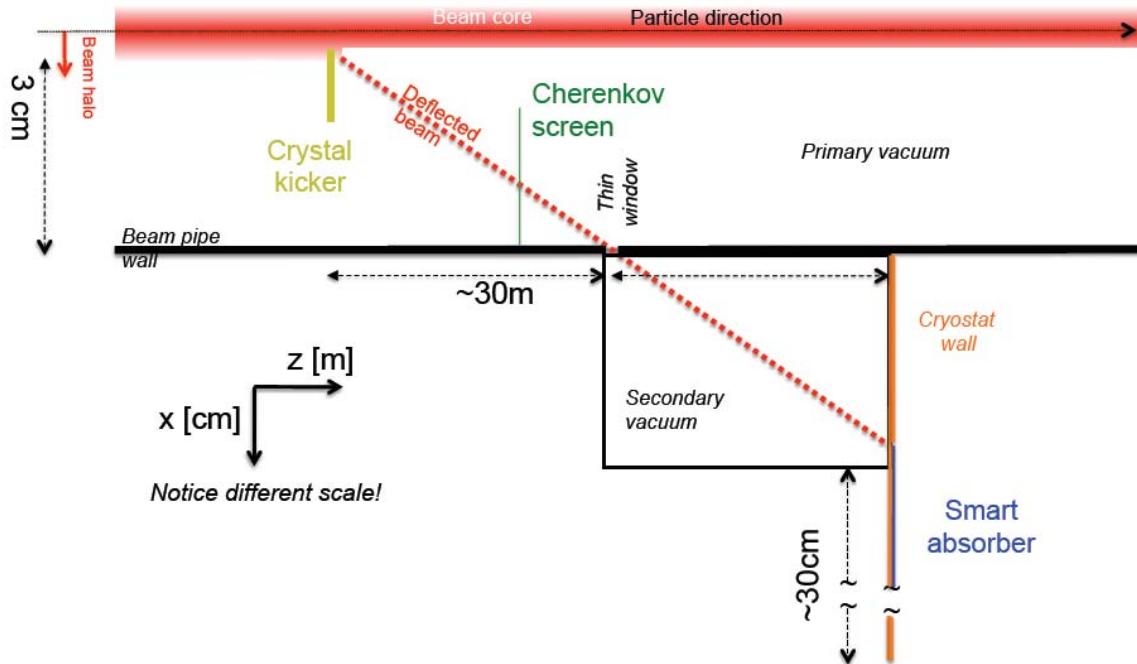
Pierre Auger Observatory



- Cosmic ray experiment observation depends on detailed **MonteCarlo code** to disentangle primary ray
- Data interpretation depends on MC used to described the shower

Measuring the cross section at various energies (100 GeV-10 TeV) on atmosphere-like target can help in those MC calibration

CRYSBEM conceptual layout



- ▶ **Crystal kick about 1 mrad** technically feasible.
- ▶ **Detection and timing** of deflected/extracted beam at the vacuum/air interface with a detector based on **Cherenkov** light emission.
- ▶ **Instrumented beam dump** in air to measure the hadronic shower structure (**the Smart Absorber**)

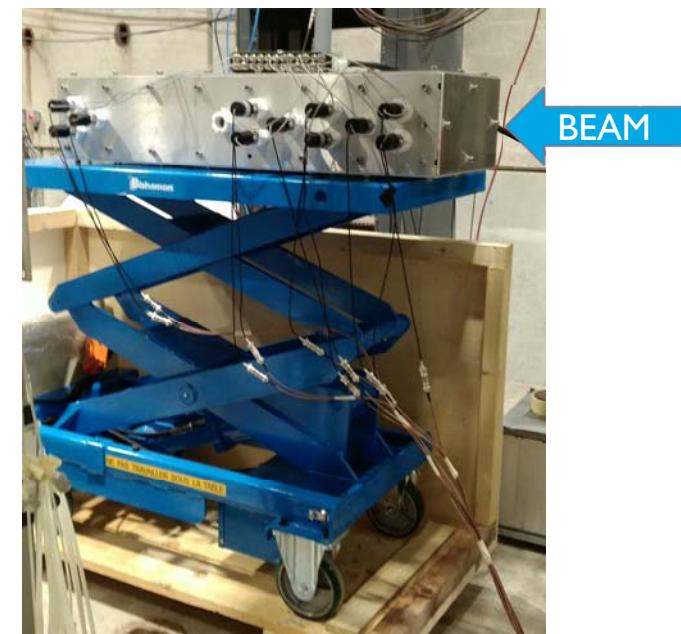
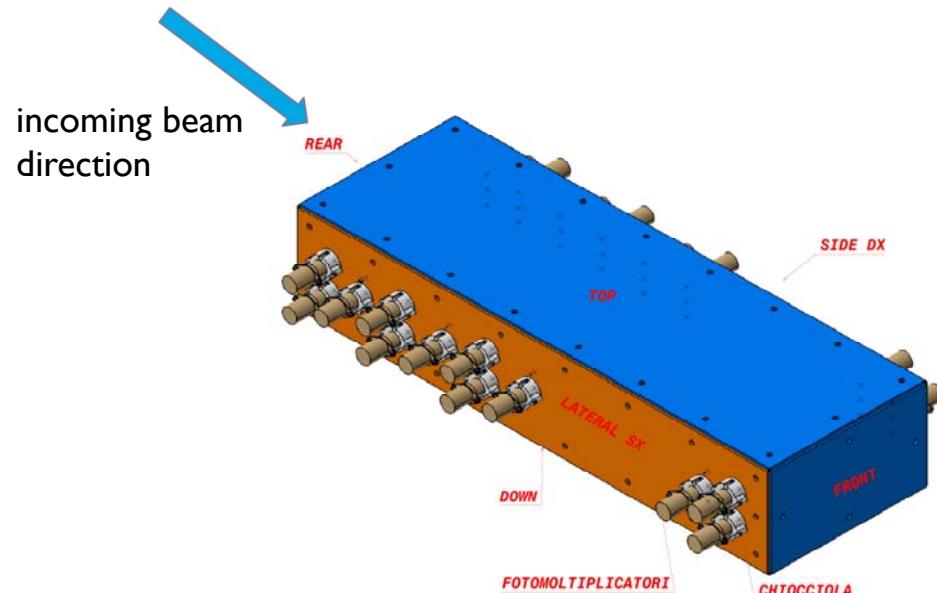


European Research Council
Established by the European Commission

Smart Absorber

The Smart Absorber

The smart absorber has been designed and built with alternating layers of tungsten and fused silica Cherenkov radiators.



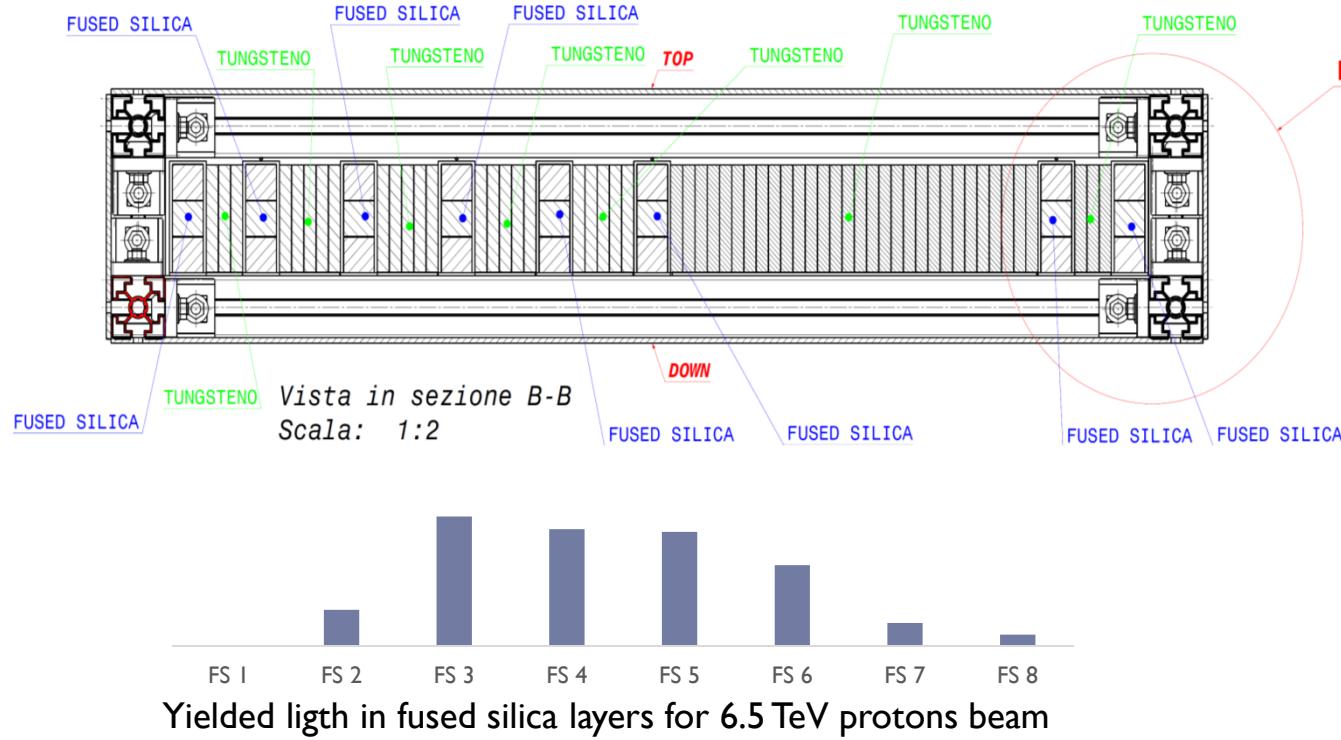
- Absorber dimensions 30 cm x 7,5 cm x 80 cm
- Cherenkov Radiators and tungsten layers for ~180 radiation length in total

The Smart Absorber



- The detectors made up of 8 layers of Cherenkov radiators , placed between layers of tungsten absorber.
- Each Cherenkov layer is composed of 3 radiators of 2.5cm x 2.5cm x 30cm.
- Every radiator has separated optical readout based on PMT (24 channels in total).

The Smart Absorber



The layers of fused silica Cherenkov radiators and the thickness of the tungsten layers are optimized, thanks to GEANT 4 simulation, to best follow the hadron shower evolution.

Optical readout

Hamamatsu R7378A

Figure 3: Dimensional Outline and Basing Diagram (Unit: mm)

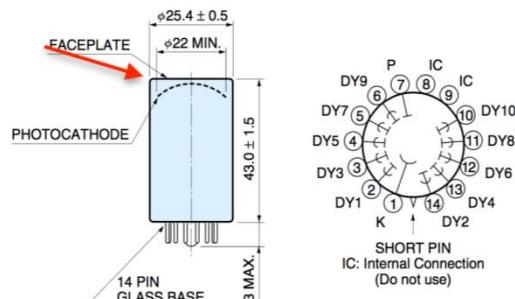
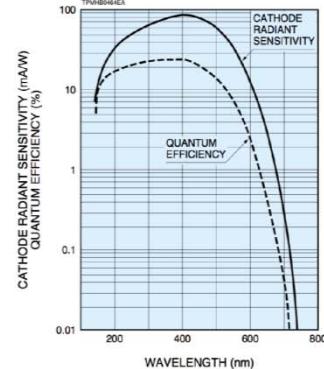
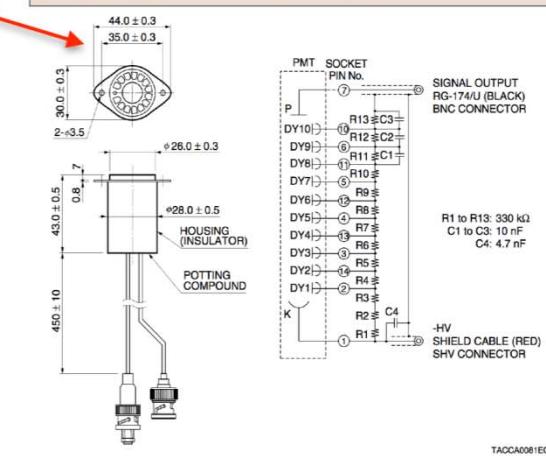


Figure 1: Typical Spectral Response



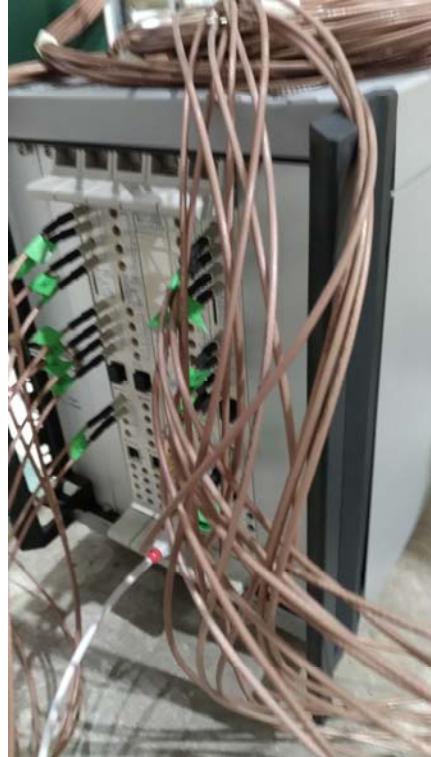
E2924-500 HV divider



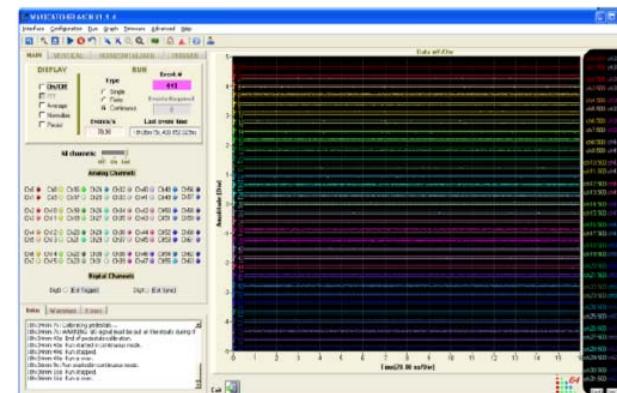
TACCA008/EC

- Good spectral response for Cherenkov (200nm)
- Fast response (anode pulse rise time ~1.5ns)
- Already tested for on-beam instrumentation (UA9 CpFM)

Readout system



- 64 channels Wavecatcher System, usb connected with external PC
 - ▶ A waveform digitizer
 - ▶ 3.2 Gs/s sampling rate
 - ▶ 12 bits
 - ▶ < 10ps rms sampling time precision

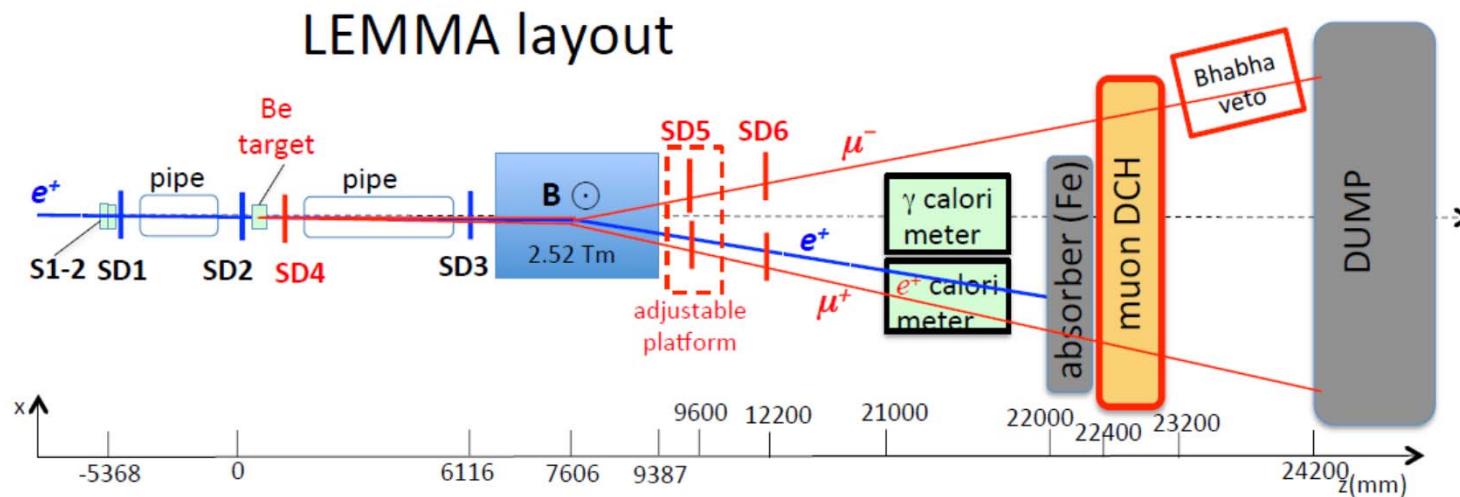
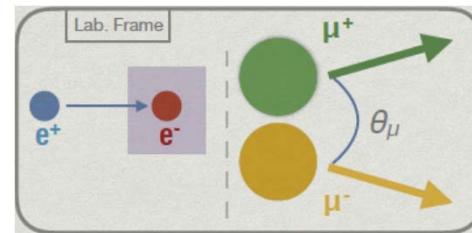


D. Breton, PhotoDet2012

The first test on LEMMA experiment

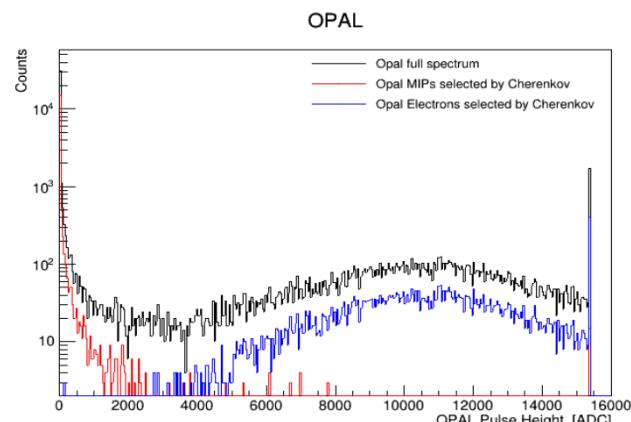
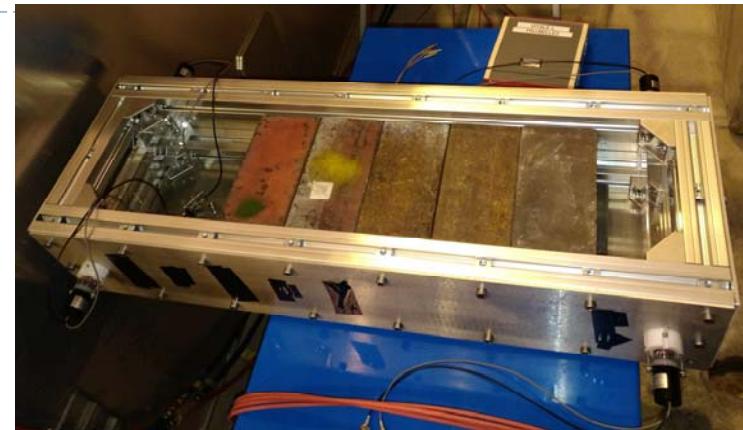
A first version of Smart absorber, with only 4 channels and with iron instead of tungsten , was used as Bhabha veto during LEMMA (Low EMittance Muon Accelerator) test beam at H4 beam line, CERN, North experimental area.

LEMMA: direct μ -pair production from $e^+e^- \rightarrow \mu^+\mu^-$
 just above production threshold ($\sqrt{s}=212\text{MeV}$), by
 using a beam of $\sim 45\text{ GeV}$ e^+ on a thin target



Lemma test beam

- only 2 layers of fused silica ,(the first and the last one) and 50 cm of iron absorber (~30 radiation length , ~ 3 pion interaction length)
- Used as veto to separate μ and e
- Only 4 channels (2 front and 2 back) readout by 8-channels WaveCatcher



Amplitude spectrum of lead glass calorimeter with end without veto

- Data was integrated on LEMMA DAQ and used for offline analysis
- Preliminary analysis shows the Smart Absorber can be used to cut Bahbah events



WaveCatcher



WaveCatcher Pro & Cons



Pro:

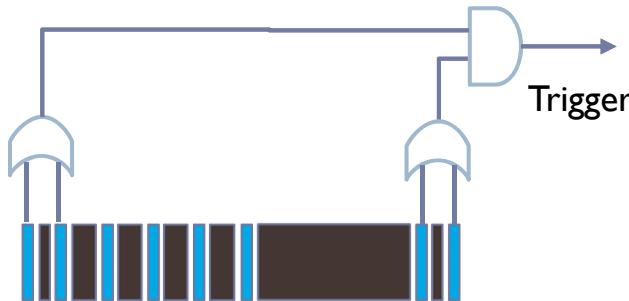
- High performance waveform digitizer with very low cost per channel
- Good time precision
- High configurable trigger logic
- Useful measurements tools
- Ready-to-use GUI interface
- UA9 already uses this electronics for some instruments (**CpFM**)

Cons:

- Long dead-time between acquisition (is not a critical issue for our application)
- Some PC interface instability (I guess due to USB driver)

WaveCatcher possible improvement

- The logic trigger features are very useful and enable self-triggered application:
 - More complex trigger logic , like coincidence of more channel in OR



- Majority trigger function even in >8 channels board
- Zero-suppression functions

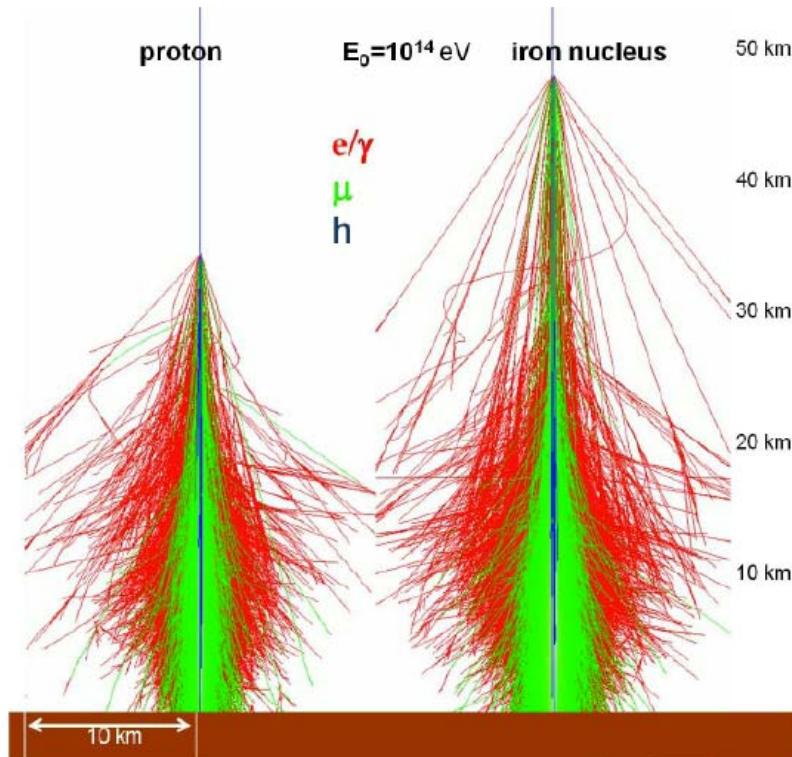
☺ A minor improvement :

- The front panel holes for MCX connectors is too small for my connector. A larger front panel holes would increase the compatibility



SPARE

Ultra High Energy Cosmic Rays



-When UHECR enters the Earth's atmosphere it produces a extensive cosmic ray showers. It is possible by analyzing these showers to discover many of the traits of the original CR.

→ Observations of Comsic Ray at ground level (i.e. Pierre Auger Observatory , HiRes..

Cosmic ray experiment observation depends on detailed **MonteCarlo code** to disentangle primary ray



Requirements and methods of CRYSBEM



ELEMENT	REQUIREMENTS	METHODS
Crystal Kicker	Orient lattice planes parallel to the beam within $\theta_c(7\text{ TeV}) \approx 2 \mu\text{rad}$ inside the vacuum beam pipe.	Finely polished, low miscut silicon crystals. Goniometer for ultra-high vacuum.
Cherenkov Screen	For deflected particles measure flux at 5%, timing <1ns, beam-spot $\approx 250 \mu\text{m}$	Fused silica slab transverse to beam. Cherenkov light internally reflected through optical micro-guide
Smart Absorber	Measure particle cross sections on nucleus A $\sigma_{\text{tot}}(\text{p-A})$ and $\sigma_{\text{tot}}(\text{Pb-A})$ as in Cosmic rays collisions	Several Cherenkov radiator layers to follow the hadron shower evolution.