

Plan

Intoduction. Crystal field strength

Diversity of orientation effects in crystals

- synchrotron type radiation and pair production, intensity growth and saturation. Shift to zero incidence angle
- polarization and spin effects, charm and beauty hyperon magnetic and electric dipole moments
- gamma-telescopes, gamma-background suppression (CLEVER)
- crystal undulators (PEARL), Ferrara experiments in u-short bent crystals

Crystal assisted collimation *Crustal cut*. *Multiple volume reflection*. *Miscut problem*. *Scattering by atomic strings*.

Simulation of high-energy electron radiation in crystals.

- the need of detailed simulations (CFA insufficiency)
- incoherent processes in crystals
- radiation under multiple volume reflection
- crystalline scintillator performance (PRIN)
- medium energy electron radiation in W <111>

The uniqueness of crystal field

Moving in oriented crystals, particles come under the action of the practically inter-atomic-scale *effective crystal field*



Planar channeling









When channeling radiation becomes synchrotron-like radiation:

$$\vartheta \approx \sqrt{\frac{2V_0}{\varepsilon}} > \frac{m}{\varepsilon} \Rightarrow \varepsilon \gg \frac{m^2}{2V_0} \sim 1 \div 10 \, \text{GeV}$$

Dipole radiation



Synchrotron-like radiation





Field amplification in the particle rest frame



Lorentz amplification of field can exceed $\gamma \sim 10^5$ times

Critical energy for hard synchrotron-like radiation and pair production by gamma-quanta

$$\mathbf{E}_{0} = \frac{m^{2}c^{3}}{e\hbar} \approx 1.32 \times 10^{16} eV$$

$$\mathbf{E}_{cryst} \sim 10^{10} \div 10^{12} \frac{Volt}{cm} \equiv 3 \times 10^3 \div 3 \times 10^5 tesla$$

$$\varepsilon_{\chi=1} = \hbar \omega_{\kappa=1} = \frac{E_0}{E_{cryst}} mc^2 \sim 10 \div 1000 GeV$$

13.6 GeV for <111> W 293 K

Synchrotron-like radiation quantum parameter χ

Invariant parameter:
$$\chi = \frac{\sqrt{\left(F_{\mu\nu}k^{\nu}\right)^{2}}}{m^{3}} \rightarrow \frac{E}{(m^{2}c^{3}/e\hbar)}\frac{\varepsilon_{e^{\pm}}}{mc^{2}} = \frac{E\gamma}{E_{0}} = \frac{E_{com}}{E_{0}}$$

Critical field:

$$\Xi_0 = \frac{m^2 c^3}{e\hbar} \approx 1.32 \times 10^{16} eV$$



"Quantum" synchrotron-like radiation is observable in crystals:

Channeling radiation



Channeling radiation



A virtual pair conversion to a real one in electric field $E \sim E_0$





Electron-positron pair production by gamma-quanta in the uniform field

Invariant parameter:
$$\kappa = \frac{\sqrt{\left(F_{\mu\nu}k^{\nu}\right)^{2}}}{m^{3}} \rightarrow \frac{E}{\left(m^{2}c^{3}/e\hbar\right)}\frac{\hbar\omega}{mc^{2}} = \frac{E\gamma}{E_{0}} = \frac{E_{com}}{E_{0}}$$

Critical electric field:

$$\mathbf{E}_0 = \frac{m^2 c^3}{e\hbar} \approx 1.32 \times 10^{16} eV$$

PP in vacuum (Schwinger)

$$\frac{dP_{e^+e^-}}{d\vec{r}} = \frac{\alpha E^2}{\pi^2} \sum_{n=1}^{\infty} \frac{1}{n^2} \exp\left(-\frac{E_0}{E_{com}}\right)$$

this process is also observable in crystals:

Critical electron and photon energies Baryshevsky, Tikhomirov UFN, 1989

ele- ment	Z	plane / axis	E _{max} (GV/cm)	H _{eff} (kilotesla)	$\frac{\hbar\omega_{cr}}{(GeV)} = \varepsilon_{cr}$
Si	14	<i>plane</i> (110)	5.7	1.9	1200
Ge	32	<i>axis</i> <110>100K	144	48	47
W	74	axis <111>	500	167	13.6

Typical angle V_0/m for synchrotron-like radiation and PP processes in crystals

$$\mathcal{G} \Box \frac{\left| \delta \vec{p}_{\perp} \right|}{\varepsilon} = \frac{1}{\varepsilon} \left| \int e \vec{E}(z) dz \right| = \frac{1}{\varepsilon} \left| \int e \vec{E}(z) \frac{d \vec{\rho}}{\psi} \right| \approx \frac{V_0}{\varepsilon \psi} \approx \frac{m}{\varepsilon} \Longrightarrow$$

$$\psi \approx \frac{V_0}{m} \gg \sqrt{\frac{2V_0}{\varepsilon}}$$

At high energies both radiation and PP processes acquire synchrotron nature



25/14-time increase of radiative energy losses of 150 GeV electrons in 0.4/1.4 mm Ge<110> 100K *A. Belkacem, PRL 1985*



Electron radiative cooling

Baryshevskii V. G., Dubovskaya I. Ya. // Phys. Lett. 1977. Vol. A62. P. 45.
Belkacem A. et al. // Phys. Lett. 1986. Vol. B177. P. 211.
Tikhomirov V. V. // Phys. Lett. 1987. Vol. A125. P. 411.
Tikhomirov V. V. // Nucl. Instr. Meth. 1989. Vol. B36. P. 282.



All plans for **PeV energies** and higher are based on **radiative cooling** in crystlas

Источник Рентгеновские лазеры 0,1-1	Кристаллическая Кристаллически ускоритель 1,0 ПэВ "воронка"



8-time increase of PP probability in Ge<110> 100K at 150 GeV *A. Belkacem, PRL 1987*



INCIDENT PHOTON ENERGY (GeV)

Crystal dichroism and birefringence



Spin effects in bent crystal

V. G. Baryshevsky. Pis'ma Zh. Tekh. Fiz. 5(1979)182; 5(1979)1529.



Channeled e^+ and e^- move or are produced by gamma-quanta in bent crystals in the regions with **dominating direction** of the planar electric field. which represents itself an origin of a number of **spin effects**.

Spin effects in bent crystal



Spin rotation in bent crystals and magnetic and electric moment measurement

 $\Omega = \gamma_S B = \frac{g \,\mu_B B}{\hbar}$ Ŕ \vec{P}_{R} **S** \vec{P}_B



E761 Collaboration, FERMILAB

"First observation of spin precession of polarized hyperons channeled in bent crystals", LNPI Research Reports (1990-1991) 129.

Energy of : 200 – 300 GeV

D. Chen "First Observation of Magnetic Moment Precession of Channeled Particles in Bent Crystals", Phys. Rev. Lett. 69 (1992) 3286.





FIG. 3. Measured polarizations and uncertainties (1 σ statistical errors) after spins have been precessed by the two crystals. The dashed arrows show the expected precessions.

Baryshevsky V.G., The possibility to measure the magnetic moments of short-lived particles (**charm and beauty baryons**) at LHC and FCC energies using the phenomenon of spin rotation in crystals, Physics Letters B, V. 757, 2016, pp 426–429.



1.0

Crystal-based angular-sensitive gamma-telescope

V.A. Baskov, V.A. Khablo, V.V. Kim et al., NIM B 122(1997)194.

V.N. Baier, V.M. Katkov, V.M. Strakhovenko, *Electromagnetic Processes at High Energies in Oriented Single Crystals*, World Scientific, Singapore, 1998.

CsI scintillators in the **Fermi** Large Aperture Telescope



GLAST Observatory after the integration of the Large Area Space Telescope. Picture taken at the General Dynamic on December 2006.





Schematic view of the GLAST imaging calorimeter.

The GAMMA-400 project



Crystals in the GAMMA-400 telescope



To the crystal gamma-telescope development



Pair production probability by 300Gev. 1 Tev and 3 TeV gamma-quanta vs the angles of incidence w.r.t. <110> Si axis.

The probability is measured in units of Beth-Heitler PP probability WBH ≈ 0.083 /cm.

Any ($E_{\gamma} \ge 10$ TeV) energy can be measured with the same detector thickness





Magnetic undulator (Ginzburg, 1947; Motz, 1953)


Crystal Undulators





855 MeV electron rechanneling in "Backe CU"









A "fortunate" 500 MeV positron trajectory



















"Principles" of construction of the "optimal" positron CU

- Maximal undulator radiation **energy**
- Minimal CU radiation spectral width
- $L_{CU} \approx L_{dech}$
- $\mathbf{K} = e F_0 \lambda_U / 2\pi m \approx 1$





**



$\Delta \omega / \omega \approx 3\%,$ $\Delta N_{\gamma} / N_e \sim 10^{-3} \gamma / e^+,$ $\Delta \theta_e \leq 20 - 30 \ \mu rad$











Spectral distribution of the radiation emitted by a 1.5 GeV positron: Solid line – in the "optimal" Si (110) CU Dotted line – in a plane 0.48 mm Si (110) crystal Positron beam incidence angle equals zero, incident beam angular divergence is $\Delta \theta_e = 10 \mu rad$ and collimation semi-apex angle is $\theta_{\gamma} = 1/8\gamma = 42.6 \mu rad$.





assisted

collimation

Orientational effects in crystals allow to facilitate collimation





channeling – limited efficiency volume reflection – small deflection angle

Channeling **efficiency** can be increased by **crystal cut**

The capture probability increase by crystal cut

V.V. Tikhomirov, JINST, 2(2007)P08006



Transverse energy reduction by the cut



The cut diminishes the potential energy preserving the transverse kinetic one

Protons cease to reach the high nuclear density regions



Phase space transformation by the cut



Channeling efficiency increase by crystal cut



Deflection of nonchanneled particles can be increased by multiple volume reflection in one crystal

Volume Reflection prediction A.M.Taratin and .A.Vorobiev Phys. Lett. A119 (1987) 425, NIM B26 (1987) 512



Large acceptance, however small deflection angles

Multiple Volume Reflection in One Crystal (MVROC) V.V. Tikhomirov, PLB 655(2007)217



Axes form *many* inclined reflecting planes

In this talk only and just for short **MVR** = multiple volume reflection in one crystal/"axial VR"

VR = "one plane VR"/"planar VR"



Proton motion in comoving reference plane



Protons are reflected from *many* different crystal plane sets in *one* crystal



Reflection angles from planes of one crystal *vs* bending radius



Reflection from different crystal planes increases VR angle about *5 times*

First MVROC observation W. Scandale et al, PLB 682(2009)274



MVROC indeed increases reflection angle 5 times

multiple volume reflection in one crystal can be combined with planar channeling



MVR+channeling in comparison with axial channeling of 7 TeV protons



Uncorrelated scattering by atomic strings will enhance random scattering practically at any alignment requirements

Particle scattering by atoms constituting a string



Particle scattering by the atoms constituting a string are correlated (coherent) coherent effects in particle scattering by the atomic string will increase the "scattering power" by nearly 1000 times

Multiple uncorrelated scattering by atomic strings



$$\psi l_{string} \sim 10^{-7} cm, \quad l_{string} \sim 0.01 cm$$

$$\frac{d\langle \vartheta_{ax}^2 \rangle}{dx} \approx \frac{4\pi^2 Z^2 \alpha^2 n}{\varepsilon^2 \psi} \frac{R - \widetilde{u}}{d_{inat}} \approx \frac{\pi}{4\ln(190/Z^{1/3})} \frac{R - \widetilde{u}}{d_{inat}} \frac{d\langle \vartheta_{am}^2 \rangle}{dx}$$

Uncorrelated scattering amplification by atomic strings in 1 cm W target



The "miscut problem"

V.V. Tikhomirov, A.I. Sytov, *The miscut angle influence on the future LHC crystal based collimation system*. Problems of Atomic Science and Technology № 1 2012. 88-92. arXiv:1109.5051.



No big difference in fact!

Need of

comprehensive simulations

of high-energy electron radiation in crystals

Electron radiative cooling

Baryshevskii V. G., Dubovskaya I. Ya. // Phys. Lett. 1977. Vol. A62. P. 45.
Belkacem A. et al. // Phys. Lett. 1986. Vol. B177. P. 211.
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Tikhomirov V. V. // Nucl. Instr. Meth. 1989. Vol. B36. P. 282.



A SIMULATION CODE FOR CHANNELING RADIATION BY ULTRARELATIVISTIC ELECTRONS OR POSITRONS

Xavier ARTRU *

Nuclear Instruments and Methods in Physics Research B48 (1990) 278-282

Laboratoire de Physique Théorique et Hautes Energies **, Université de Paris-XI, 91405 Orsay, France

2.2. Multiple scattering

After instruction (2a), we insert

$$\boldsymbol{p}_l = \boldsymbol{p}_l + \boldsymbol{Q}, \tag{3}$$

where $Q = \sum_{1}^{n} q_{j}$ is the cumulated momentum transfer of *n* incoherent collisions during dt_{i} . *n* is chosen at random about $\langle n \rangle = dt_{1}\rho_{N}(x_{i})\sigma_{inco}$, where $\rho_{N} = (2\pi u_{1}^{2})^{-1} \exp(-r^{2}/2u_{1}^{2})$ is the local density of nuclei and

$$\sigma_{\rm inco} = 4(Z/137)^2 \int q^{-4} \, \mathrm{d}^2 q f_{\rm A}(q^2) f_{\rm DW}(q^2) \tag{4}$$

is the incoherent cross section. $f_A \simeq [1 + (a_{TF}q)^{-2}]^{-2}$ represents the atomic screening and $f_{DW} = 1 - \exp(-u_1^2q^2)$ is the Debye-Waller factor. For a fast calculation of Q we take

$$f_{\rm A}(q^2)f_{\rm DW}(q^2) = (q/q_0)^4 \int_0^1 h \, \exp(-hq^2/q_0^2) \, \mathrm{d}h,$$




simulation

method

Key simulation points:

Trajectory simulations in most **realistic potentials**

Simulation of **incoherent scattering** on both nuclei and electrons

Separate simulation of **single** and **multiple** scattering

Direct integration of **Baier-Katkov formula**

Infinite trajectories, density effect...

Radiation process simulations from the *"First Principles"*

The general expression for radiation intensity

$$\frac{d^2 I}{d\omega d^2 \theta} = \frac{\alpha \omega^2 d\omega}{8\pi^2 \varepsilon'^2} \times \int \int dt_1 dt_2 \left[(\varepsilon^2 + \varepsilon'^2) (\mathbf{v}_\perp(t_1) - \boldsymbol{\theta}) (\mathbf{v}_\perp(t_2) - \boldsymbol{\theta}) + \omega^2 / \gamma^2 \right]$$
$$\exp\left\{ i \frac{\omega \varepsilon}{2\varepsilon'} \left[\int_{-\infty}^{t_1} \left(\gamma^{-2} + (\mathbf{v}_\perp(t') - \boldsymbol{\theta})^2 \right) dt' + \int_{-\infty}^{t_2} \left(\gamma^{-2} + (\mathbf{v}_\perp(t'') - \boldsymbol{\theta})^2 \right) dt'' \right] \right\}$$

contains two integrals

$$A = \int \exp\left\{i\frac{\omega\varepsilon}{2\varepsilon'}\int_{-\infty}^{t} \left[\gamma^{-2} + (\mathbf{v}_{\perp}(t') - \boldsymbol{\theta})^{2}\right]dt'\right\}dt,$$

$$\mathbf{B} = \int \left(\mathbf{v}_{\perp}(t) - \boldsymbol{\theta} \right) \exp \left\{ i \frac{\omega \varepsilon}{2\varepsilon'} \int_{-\infty}^{t} \left[\gamma^{-2} + \left(\mathbf{v}_{\perp}(t') - \boldsymbol{\theta} \right)^{2} \right] dt' \right\} dt$$

and slowly decreases with radiation angle θ , complicating its numerical integration.

Radiation at sharp change of particle trajectory



$$I pprox rac{ic}{\omega} \left(rac{v'}{c - n \cdot v'} - rac{v}{c - n \cdot v}
ight),$$

$$\frac{d\mathcal{E}}{d\omega} = \frac{2e^2}{\pi c} \left(\frac{2\xi^2 + 1}{\xi\sqrt{\xi^2 + 1}} \ln\left(\xi + \sqrt{\xi^2 + 1}\right) - 1 \right).$$

Single scattering effects are treated separately

$$\begin{split} A &= \int_{-\infty}^{\infty} \exp\{i\varphi(t)\}dt = \frac{i}{\dot{\varphi}(+0)} - \frac{i}{\dot{\varphi}(-0)} + \\ &i \sum_{i=1}^{N} \left\{ \left[\frac{1}{\dot{\varphi}(t_i+0)} - \frac{1}{\dot{\varphi}(t_i-0)} \right] \exp i\varphi(t_i) - \frac{2\ddot{\varphi}(\bar{t}_i)}{\dot{\varphi}^3(\bar{t}_i)} \sin \left[\frac{\varphi(t_i-0)-\varphi(t_{i-1}+0)}{2} \right] \exp i\varphi(\bar{t}_i) \right\}, \\ \vec{B} &= \int_{-\infty}^{\infty} \left[\vec{v}_{\perp}(t) - \vec{\theta} \right] \exp\{i\varphi(t)\}dt = \left[\frac{i}{\dot{\varphi}(+0)} - \frac{i}{\dot{\varphi}(-0)} \right] \left(\vec{v}_{\perp}(0) - \vec{\theta} \right) + \\ &i \sum_{i=1}^{N} \left\{ \begin{bmatrix} \frac{\vec{v}_{\perp}(t_i) + \vec{\vartheta}_i - \vec{\theta}}{\dot{\varphi}(t_i+0)} - \frac{\vec{v}_{\perp}(t_i) - \vec{\theta}}{\dot{\varphi}(t_i-0)} \end{bmatrix} \exp i\varphi(t_i) - \\ & \frac{2}{\dot{\varphi}^2(\bar{t}_i)} \left[\dot{\vec{v}}_{\perp}(\bar{t}_i) - \left(\vec{v}_{\perp}(\bar{t}_i) - \vec{\theta} \right) \frac{\ddot{\varphi}(\bar{t}_i)}{\dot{\varphi}(\bar{t}_i)} \right] \sin \left[\frac{\varphi(t_i-0) - \varphi(t_{i-1}+0)}{2} \right] \exp i\varphi(\bar{t}_i) \\ & \text{where } \omega' = \varepsilon/(\varepsilon - \omega), \ \ddot{\varphi}(t) = \omega' \left(\vec{v}_{\perp}(t_i) - \vec{\theta} \right) \dot{\vec{v}}_{\perp}(t) \text{ and } \ \bar{t}_i = (t_i + t_{i-1})/2. \end{split}$$

Simulation of radiation accompanying multiple volume reflection

Radiation amplification under Multiple Volume Reflection in One Crystal (120 GeV e⁻, 2 mm Si <111>)



- soft radiation amplification by reflection
 by different planes
 - hard radiation amplification
 by axial field



120 GeV electron radiation in <001> PWO

(Dr. L. Bandiera PRIN project)



120 GeV electron energy losses on 4 mm of **amorphous** PbWO₄ (1), crystalline PbWO₄ with (3) and without (2) PP by radiated photons

Electromagnetic shower development acceleration in PWO at $E_e = 120 \text{ GeV}$



First simulations of shower development in ECAL CMS



Electromagnetic shower acceleration in PWO at 50, 100 and 1000 GeV





em shower maximum shifts

by 2 ÷ 5

radiation lengths

Electromagnetic shower acceleration in PWO can influence H boson mass measurements



Thickness of the CMS ECAL calorimeters can be made smaller



A first step to positron source simulations: electron radiation in 1 mm W <111>

A 20 GeV in <110> W field trajectory



z, cm



Transverse velocity component evolution

Mind single scattering!

Radiation amplification in <110> W at 20 GeV (photon spectra thin crystal limit)



20 GeV electron trajectory in W<111> 1 mm



20 GeV electron trajectory in W<111> 1 mm



z, cm

A part of electron trajectory in W<111>



20 GeV e⁻⁻ radiation amplification in <110> W 1 mm



20 GeV electron spectrum behind 1 mm <110> W



Possible cooperation directions

as a conclusion

- positron source
- crystal assisted collimation
- crystal undulators
- crystal scintillators
- gamma-telescopes
- radiation pair production at high energies, polarization and spin effects
 charm and beauty hyperon magnetic and electric dipole moments

Thank you for attention!