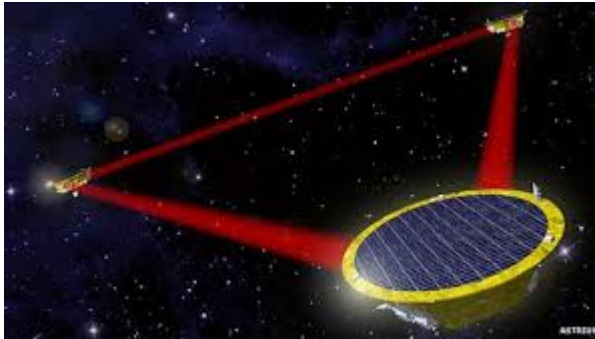


# Round table : roadmaps

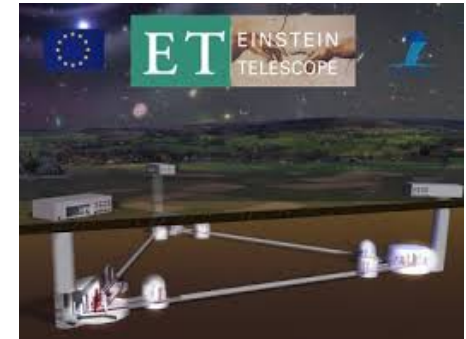
today



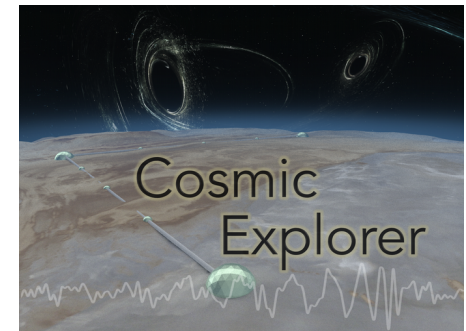
today



tomorrow



- GWIC roadmap
- European roadmaps (ESFRI, APPEC)
- National roadmap (TGIR)



# Setting the scene

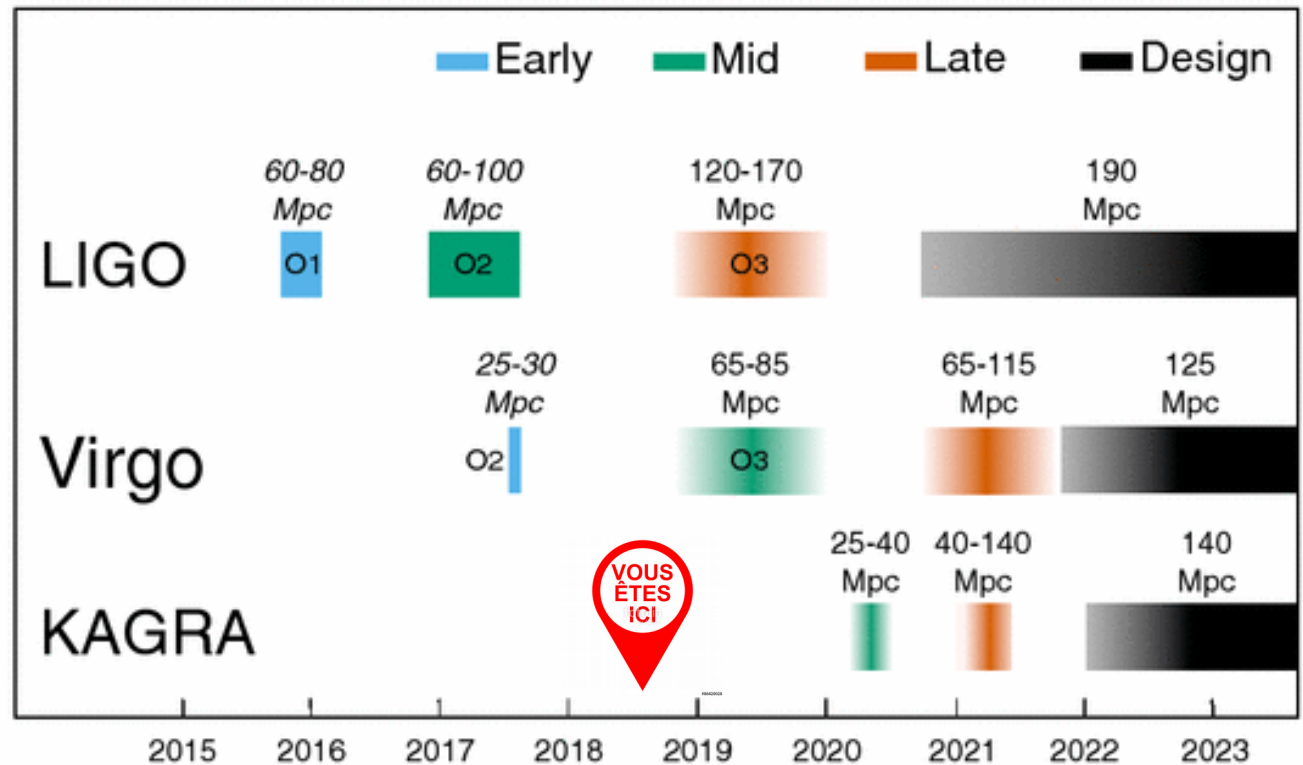
- LIGO/Virgo discoveries have modified the landscape and validated the global network approach.
- International roadmaps :
  - APPEC 2017-2026 **document** (release jan 2018): **full support to LISA & full support to ET for acquiring ESFRI status.**
  - ESFRI 2016 **roadmap**. GW detectors not in projects nor landmarks despite announced Astronomy & astroparticle science goals :
    - understand the extremes and origins of the universe
    - observe the formation of galaxies and their evolution
    - understand the formation of stars and planets
    - understand the solar system and life
    - observe gravitational waves(2018 update starting)
  - US astronomy decadal survey starting end of 2018 : **GW astronomy is on their radar.**
  - **ASTRONET** roadmap update for 2020 : any GW input required ?
  - European Particle Physics roadmap update : increase interest for GW physics ?
- GW International Committee (GWIC) 3G roadmap has started in 2016. 2 subcommittees :
  - Science case : document release dec. 2018.
  - Detectors R&D.

Before 3G : current detectors upgrade roadmap

# Current generation detectors

- Roadmap regularly updated by LVC+KAGRA : Abbott, B.P. et al. Living Rev Relativ (2018) 21: 3. <https://doi.org/10.1007/s41114-018-0012-9>
- Design sensitivity (LIGO 190Mpc / Virgo 125 Mpc / KAGRA 140 Mpc) : a bit delayed compared to current planning : 2022

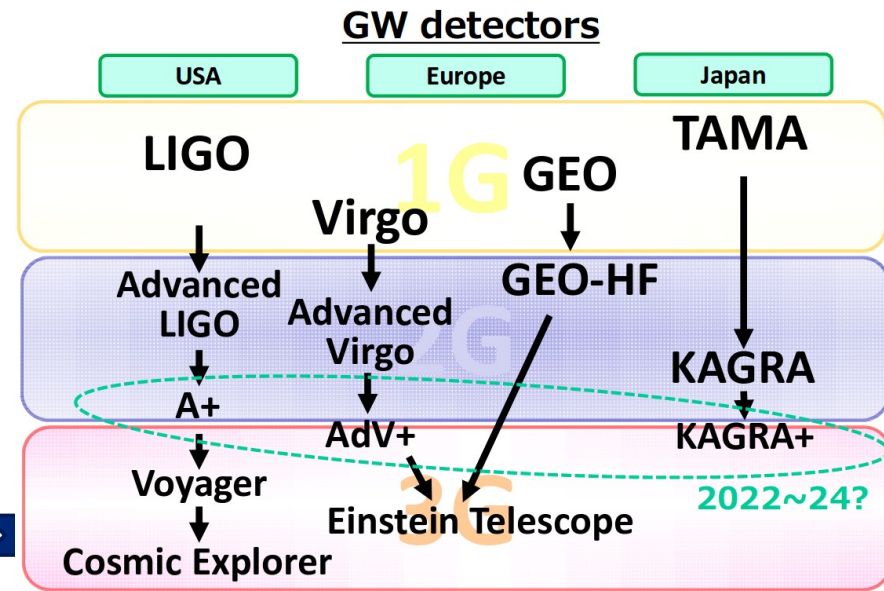
LIGO/Virgo collaboration



- **What's new ?** : KAGRA PI (Takaaki Kajita) proposed KAGRA to join end of O3 for « scientific results by end of 2019 ». Acceleration of KAGRA original roadmap.

# Current generation detectors

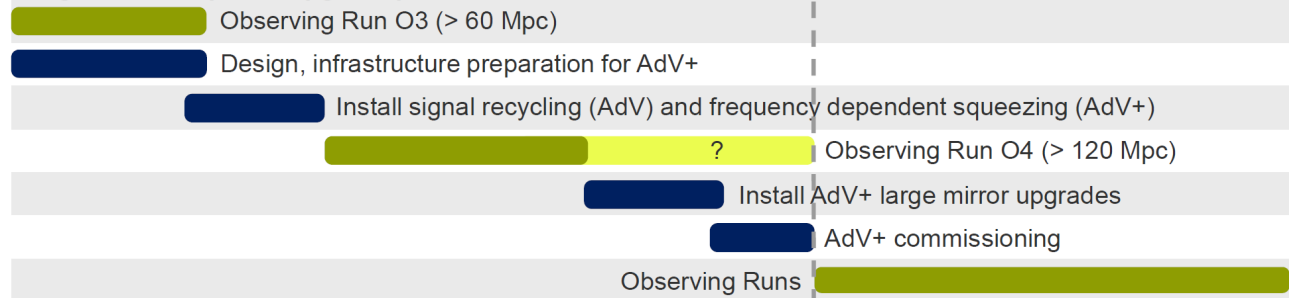
- Longer term planning : LIGO A+ and AdV+
- Longer term planning : LIGO India
- And KAGRA+ : ideas for upgrade



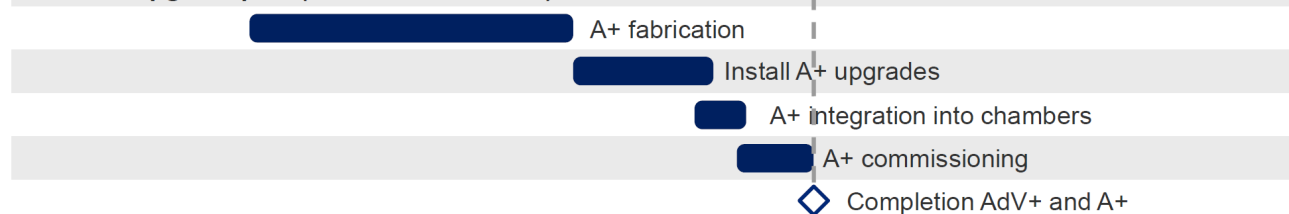
Five year plan for observational runs, commissioning and upgrades



## Virgo AdV+ Proposed upgrade plan



## LIGO A+ Upgrade plan (see LIGO-G1702134)



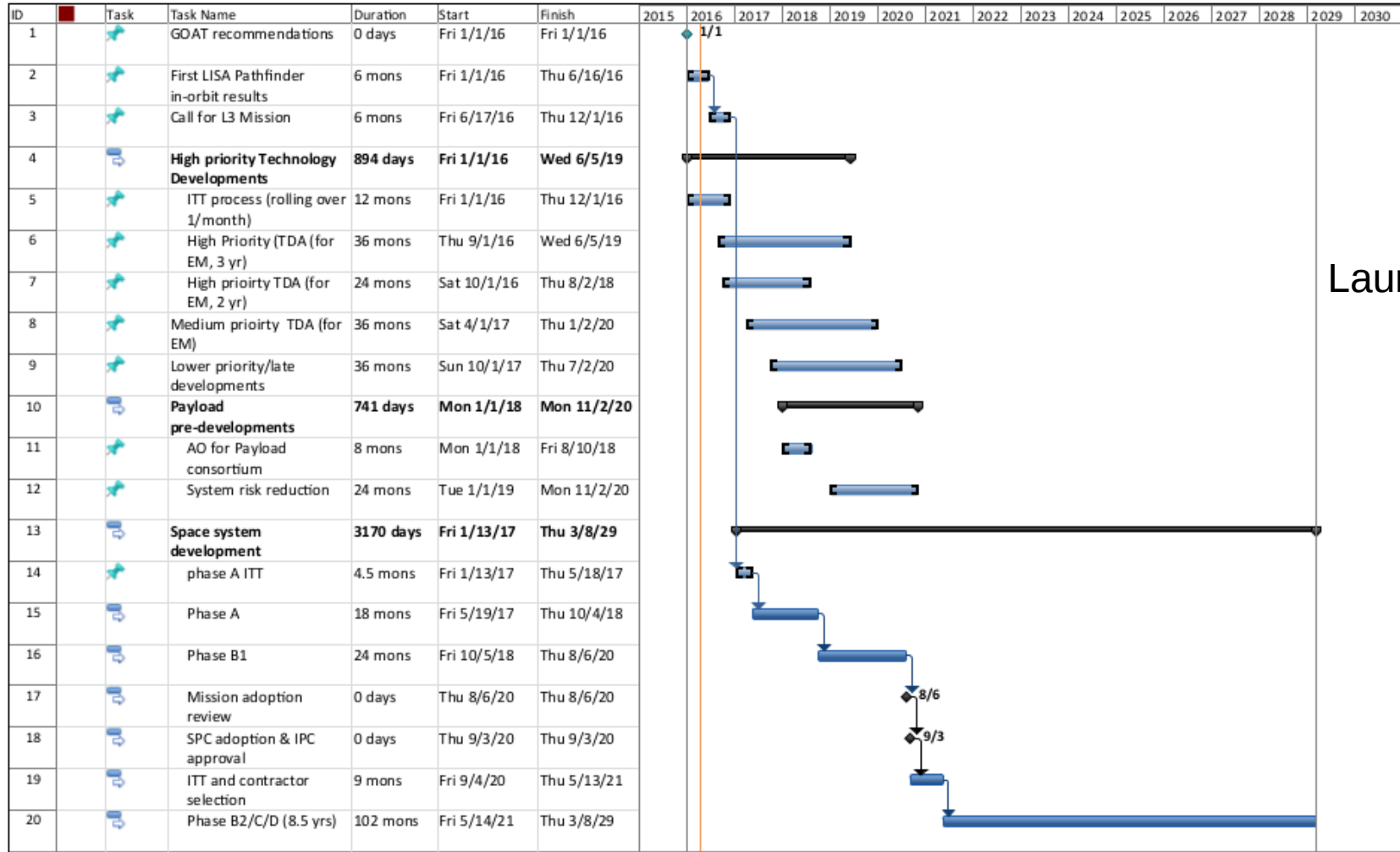
Note: duration of O4 has not been decided at this moment

Kentaro Somiya, GWADW 2018



LISA proposal [[arxiv:1702.00786](https://arxiv.org/abs/1702.00786)]

LISA consortium is born (again) in 2018



Launch 2034

Project: L3 candidate schedule Date: Fri 4/15/16	Task		External Tasks		Manual Task		Finish-only	
	Split		External Milestone		Duration-only		Deadline	
	Milestone		Inactive Task		Manual Summary Rollup		Progress	
	Summary		Inactive Milestone		Manual Summary			
	Project Summary		Inactive Summary		Start-only			

# GWIC 3G detectors roadmap

- <https://gwic.ligo.org/3Gsubcomm/>
- Charge (nov. 2016):
  - Science case → F2F Oct 2018 AEI
  - R&D coordination
  - Community networking
  - Agency interfacing
  - Investigation of Governance Structures

Beverly Berger, Rochester Institute of Technology, USA  
Federico Ferrini (co-chair), European Gravitational Observatory, Italy  
Gabriela González, Louisiana State University, USA  
James Hough, University of Glasgow, UK  
Stavros Katsanevas, Astro Particle and Cosmology Laboratory, Paris  
Ajit K. Kembhavi, Inter-University Centre for Astronomy and Astrophysics, India  
Frank Linde, NIKHEF, Netherlands  
Jay Marx (co-chair), Caltech, USA  
David McClelland, Australian National University, Australia  
Masatake Ohashi, Institute of Cosmic Ray Research, Japan  
Fulvio Ricci, Università La Sapienza, Italy  
Gary Sanders, Thirty Meter Telescope, USA  
Stan Withcomb, Caltech, USA





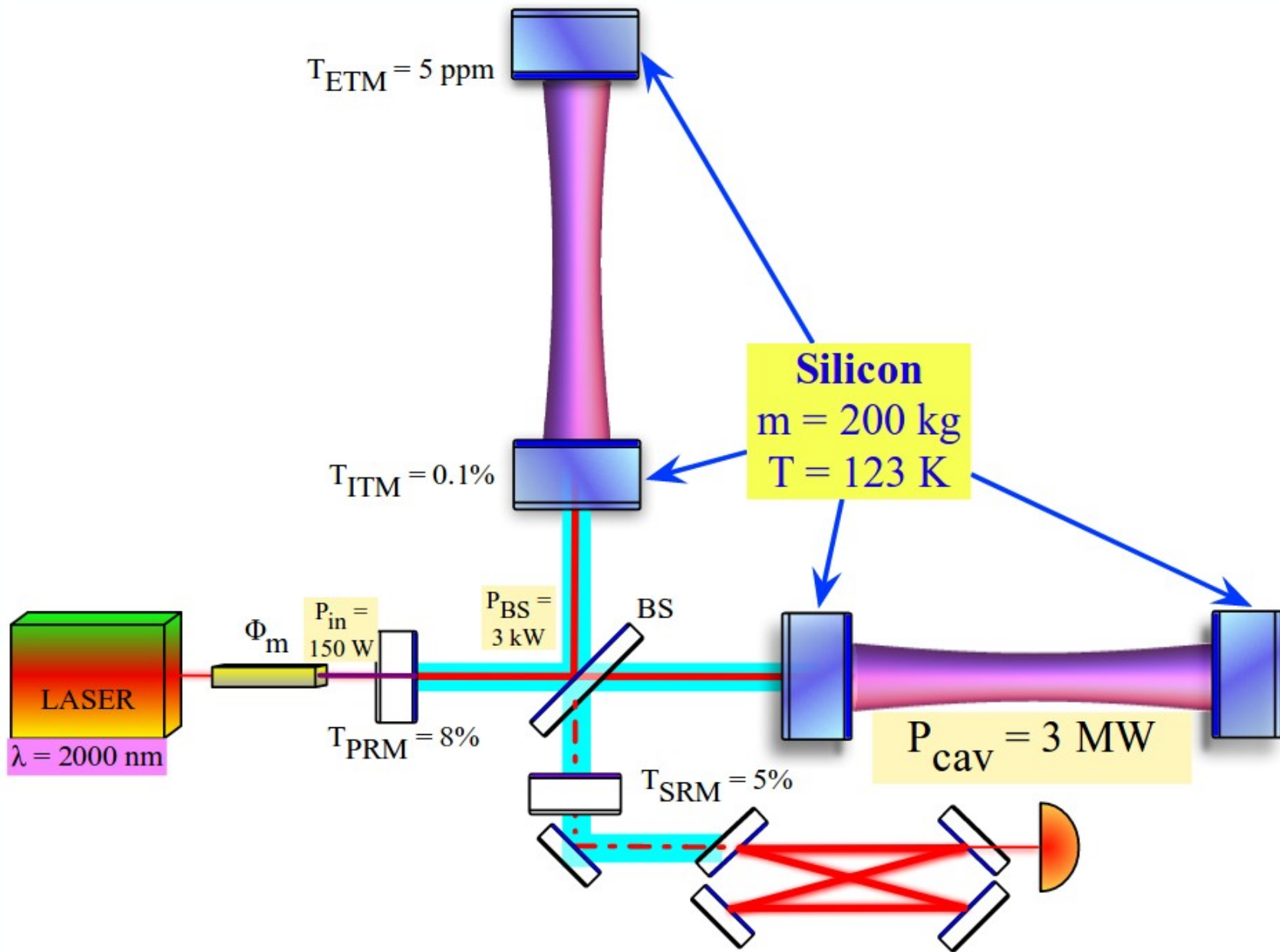
# GWIC 3G roadmap : detectors R&D

Topic	Topic coordinators	
Communication with outside	Harald Lück David McClelland	Set up wiki, make sure R&D plans&progress are communicated outside GWIC
Light sources (Lasers + squeezers)	Benno Wilke Anil Prabhakar David McClelland	Different $\lambda$ s (1064, 1550, 2100?), different powers, [not interfacing of squeezers]
Coatings	Geppo Cagnoli Marty Fejer	Requires large efforts
Low Frequencies (NN) + site requirements	Jan Harms Stefan Hild	NN subtraction; influence of geology and facility geometry on NN and seism. noise
Simulations & Controls	Andreas Freise Rana Adhikari	Error signal creation, PI, control systems
Facilities & infrastructure	Mike Zucker Fulvio Ricci	Cost saving designs (incl. Op. Costs?) maintain quietness
Cryogenics	Ando Massaki Fulvio Ricci Rana Adhikari	Different cryo regimes (4K, 20K, 124K, 300K)
Suspensions and Isolation	Norna Robertson Gabriela Gonzalez Giovanni Losurdo	Materials, sensing, actuation, coupling (for diff. temperatures)
Core optics	Geppo Cagnoli Marty Fejer	May need 'internal' x-tal growth facilities not to rely on progress of industry
Aux optics	Anil Prabhakar Matt Evans	New $\lambda$ s, lower loss, TCS
Quantum noise + Configurations	Jan Harms Stefan Hild Giovanni Losurdo Andreas Freise	Include FD squeezing,

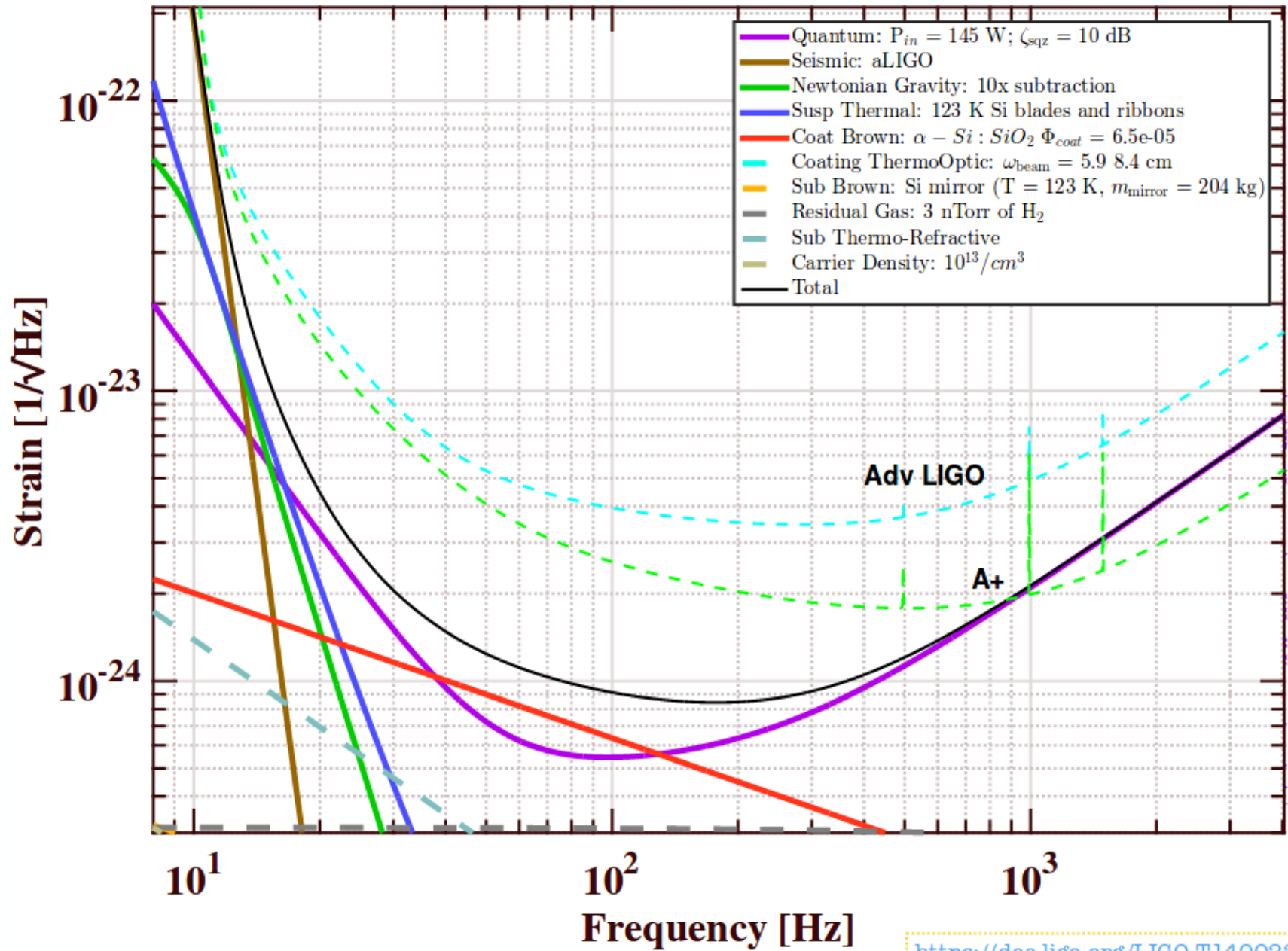




# LIGO Voyager



# LIGO Voyager



# LIGO Voyager

## Timeline

	2015-2018	2019-2022	2023-2026	2027-2030
<b>Systems</b>	Systems definition and noise budget. CDR	Continue to refine noise model based on incoming measurements of subsystem components. PDR	Deinstall A+	FDR and procure subsystems and components
<b>Cryogenics</b>	Cooling model & measurement of emissivity altering materials	Choose heat shield mat. & measure BRDF	Design cooling subsystem and Prototype at 40 meter	Install
<b>Core Optics</b>	Calc. and review noise sources in bulk Si. Study polishing process and optical abs.	Measure phase noise and Scatter in bulk Si. Develop process for polishing and do Pathfinder	Procure Substrates	Install
<b>Coatings</b>	Identify and select possible IBS materials (a-Si & Si or SiN)	Develop process for IBS coating		Procure Coatings
<b>Laser</b>	Identify optimum wavelength and best laser technology	Measure Laser noises	Laser shoot-out. FDR Procure & test 50 Watt operation	Continue working with vendor toward 150 W operation
<b>Input Optics</b>	Identify AOM, PM & Isol. materials and measure optical absorptions	Test component performance at selected wavelength	FDR & Procure Input Optics	Install
<b>Suspensions</b>	Identify ribbon fabrication technique and measure thermal noise in Si ribbons	Design and fabricate 40m prototype SUS	Test SUS at 125 K and model thermal performance	FDR & Procure SUS
<b>Squeezing</b>	Literature and modeling survey on PD, NLO materials and squeezer design	Develop design and demonstrate squeezing	FDR and develop vacuum squeezer	Fabricate Vacuum squeezer
<b>AOS</b>	Literature & modeling survey Viewports, TCS, OpLev & Baffles	Demonstrate baffles & measure BRDF, demonstrate TCS in vacuum	FDR, design and procure AOS	

Commissioning

