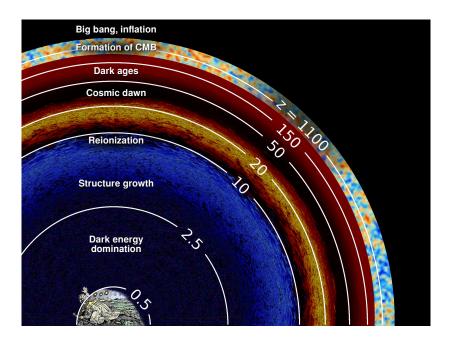
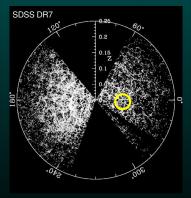
HIRAX (Hydrogen Intensity and Real-time Analysis eXperiment)

(presented by Martin Bucher APC, Université Paris 7, Paris France)

25 June 2018, LAL, Orsay, France



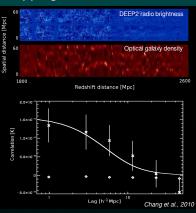
BAOs with hydrogen intensity mapping



We want large volumes (large sky, large z range) for precision cosmology

...but counting individual galaxies is hard, and getting to high redshifts is challenging

150 Mpc scale is big (degree scale)



Throw away resolution: use HI intensity mapping to measure matter distribution AND obtain redshift information.

Use BAO peak as a standard ruler to chart the universe's expansion history, probe dark energy.

Required specs for BAO intensity mapping



- Maximize sensitivity on scales of interest → Use compact array geometry
- Redshift range of interest: 0.8 < z < 2.5 to capture dark energy domination at z ~ 2 → Required frequencies: 400 - 800 MHz
- BAO 150 Mpc angular scale: 3 1.3 degrees at 0.8 < z < 2.5 → Required baseline lengths: 15 - 60 meters
- BAO scale along line of sight: 20 12 MHz at 0.8 < z < 2.5 → Required freq resolution: minimum ~100 channels, more for foregrounds and higher order peaks
- BAO signal level: ~0.1 mK → Low system temperature, large collecting area

The <u>Hydrogen Intensity and Real-time Analysis eXperiment</u>

Science goals:

Measure baryon acoustic oscillations with HI intensity mapping

Characterize dark energy Radio transient searches

Pulsar searches

Neutral hydrogen absorbers

Diffuse polarization of the Galaxy

Instrumental approach:

1024 close-packed 6-m dishes
Dishes are stationary but can be tilted

Operating frequency: 400 – 800 MHz, equivalent redshift = 0.8 – 2.5

Working closely with CHIME: channelize with FPGA ICE boards, correlation with GPUs

Location: SKA/Karoo (site agreement in progress)

The acronym:



Rock hyrax / dassie



http://www.acru.ukzn.ac.za/~hirax

Complementarity with CHIME

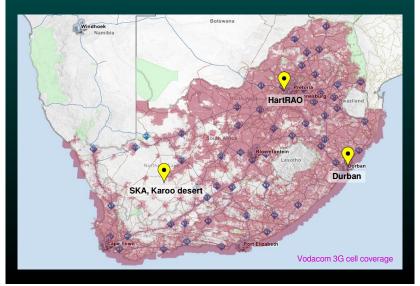


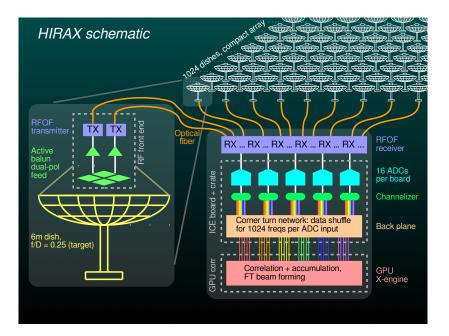


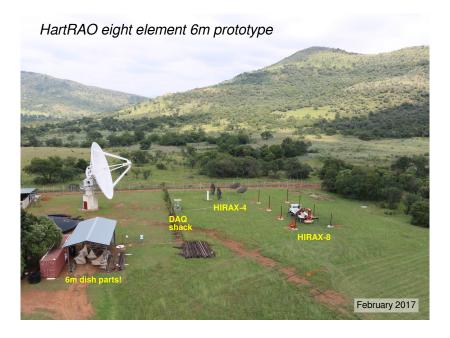
	CHIME	HIRAX
Site	DRAO, Canada	Karoo (lower RFI, no snow)
Telescope	Cylinder array	Dish array (different systematics)
Field of view	100° NS, 1°-2° EW	5° - 10° deg
Beam size	0.23° - 0.53°	0.1° - 0.2°
Collecting area	8000 m ²	28,000 m ²
Sky coverage	North	South

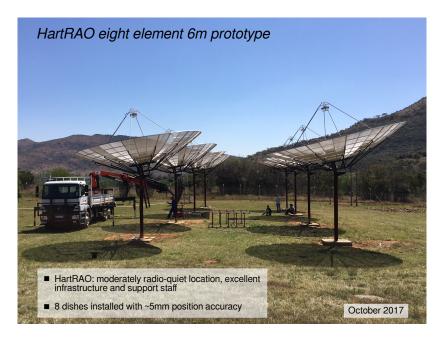
Optical surveys in the south, esp. LSST: cross-correlate for foreground mitigation and other science. More pulsars in the south.

Where we are in South Africa

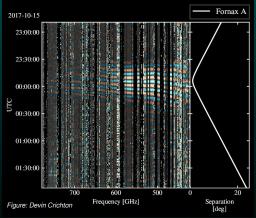






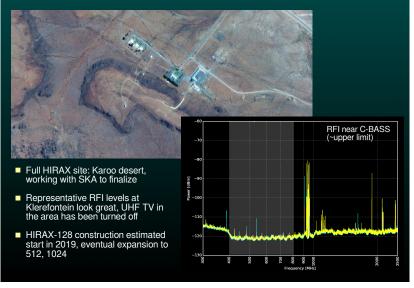


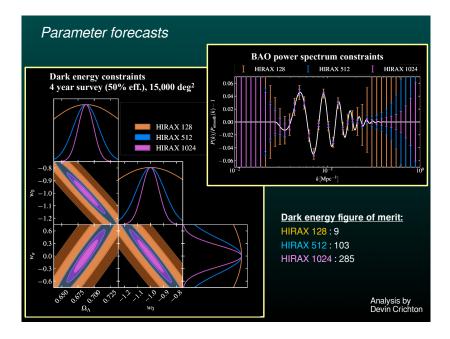
HIRAX-8 commissioning in progress



- All dishes instrumented, commissioning and troubleshooting in progress
- First tests of active-balun feeds, RFOF modules, ICE board, GPU correlator
- We've seen fringes! But lots of work ahead...
- Current plan: buildout to 32 dishes at HartRAO for further subsystem tests

Next phase: 128 element array in the Karoo





Fast radio bursts

- Fast radio bursts: short (~ms), bright (~Jy) radio transients. Distances are likely cosmological because of observed dispersion.
- Fast de-dispersion code developed for HIRAX and tested on archival GBT data: bonus detection!
- HIRAX: estimated ~10 FRB detections per day, possible coordinated observations with HERA

Outrigger stations: ~1000 km baselines for VLBI positions for HIRAX events



4 outriggers 0.5 arcsec 6 outriggers

Summary & future prospects



- \blacksquare HIRAX will do hydrogen intensity mapping to study BAOs and the universe's expansion history at 0.8 < z < 2.5
- Eight element prototype is being commissioned at HartRAO, have obtained first fringes
- Next phase: 128 elements in the Karoo, aiming for 2019
- Outrigger station site testing has begun

Possibilities for Expanded French Involvement

- International partners welcome: Canada, France (now Ken and Martin), India, Switzerland, United Kingdom, United States, Taiwan
- Synergies in studying common science/analysis/instrumental issue
- ► HIRAX Collaboration very open to new members. Financial/in kind contribution highly desirable. Invitation to discuss.