

Efficient numerical methods for solving evolution equations in QCD

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Outline

Collinear factorisation:

-  definition of parton distributions and their evolution.

Lagrange interpolation:

-  basic concepts,

-  a showcase test,

-  (aside: derivation and integration).

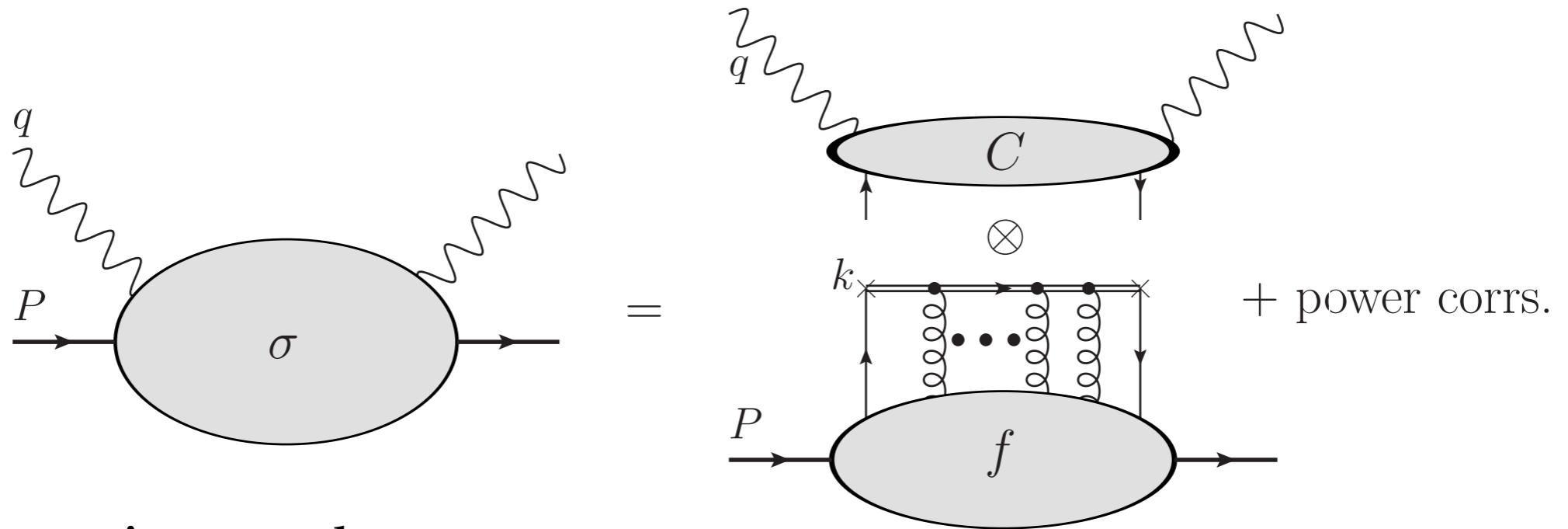
Evolution on the grid:

-  technicalities,

-  numerical results.

Collinear factorisation in QCD

- Typical examples where *collinear* factorisation in QCD applies are the cross section for deep-inelastic scattering (cut diagram) and deeply-virtual Compton scattering (uncut diagram) at large $Q^2 = -q^2$:



- The cross section reads:

$$d\sigma = C \otimes f^{(0)} + \mathcal{O}(\Lambda_{\text{QCD}}^n / Q^n)$$

- C is the “partonic” cross section:

- computable in perturbation theory (coll. divergences may need to be subtracted),

- $f^{(0)}$ is the *bare* partonic distribution:

- non-perturbative**, *i.e.* perturbation theory inapplicable,

- UV divergent: needs to be **renormalised!**

Renormalisation and evolution

🍏 Renormalisation of the a partonic distributions proceeds as usual:

$$f^{(0)}(\epsilon) = Z(\mu, \epsilon) \otimes f(\mu)$$

🍏 so that:

$$\frac{df(\mu)}{d \ln \mu} = P(\mu) \otimes f(\mu)$$

🍏 with:

$$P(\mu) = - \lim_{\epsilon \rightarrow 0} \frac{d \ln Z(\mu, \epsilon)}{d \ln \mu}$$

🍏 P is finite and computable in perturbation theory (splitting function).

🍏 Prominent example: the DGLAP equations that in their full glory read:

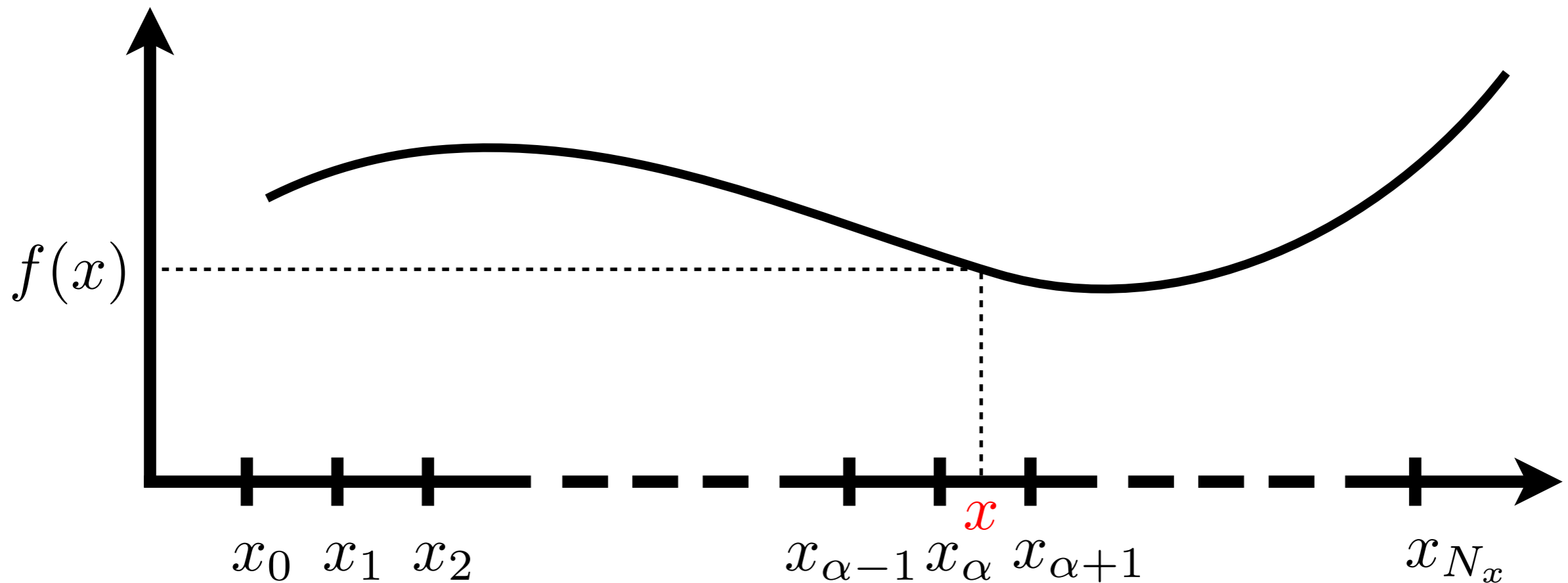
$$\frac{df_i(x, \mu)}{d \ln \mu^2} = \sum_j \int_x^1 \frac{dy}{y} P_{ij}(y, \alpha_s(\mu)) f_j\left(\frac{y}{x}, \mu\right) \quad i, j = g, q$$

🍏 A set of **coupled integro-differential equations**.

🍏 The focus of this talk is the **efficient numerical solution** of equations of this type.

Lagrange interpolation

🍏 Goal: interpolating a 1D function knowing its values on a grid.



🍏 “Moving” polynomial interpolation of degree k :

$$f(x) = \sum_{\alpha=0}^{N_x} w_{\alpha}^{(k)}(x) f(x_{\alpha}) + \mathcal{O}\left(\frac{f^{(k+1)}(x)}{(k+1)!} h^{k+1}\right)$$

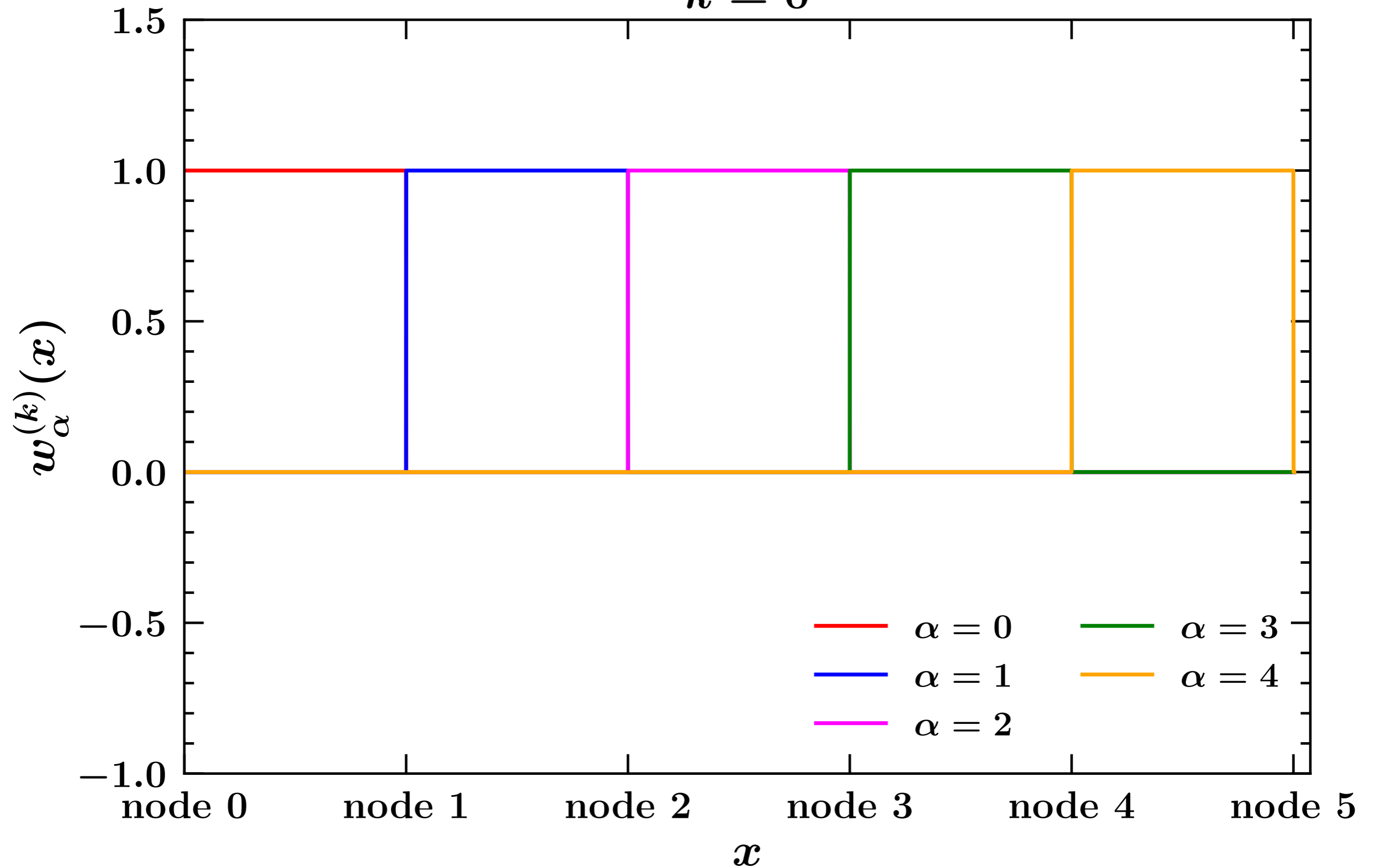
🍏 assuming that $x_{\alpha} < x \leq x_{\alpha+1}$:

$$w_{\alpha}^{(k)}(x) = \sum_{i=0, i \leq \alpha}^k \theta(x - x_{\alpha-i}) \theta(x_{\alpha-i+1} - x) \prod_{m=0, m \neq i}^k \frac{x - x_{\alpha-i+m}}{x_{\alpha} - x_{\alpha-i+m}}$$

Lagrange polynomials

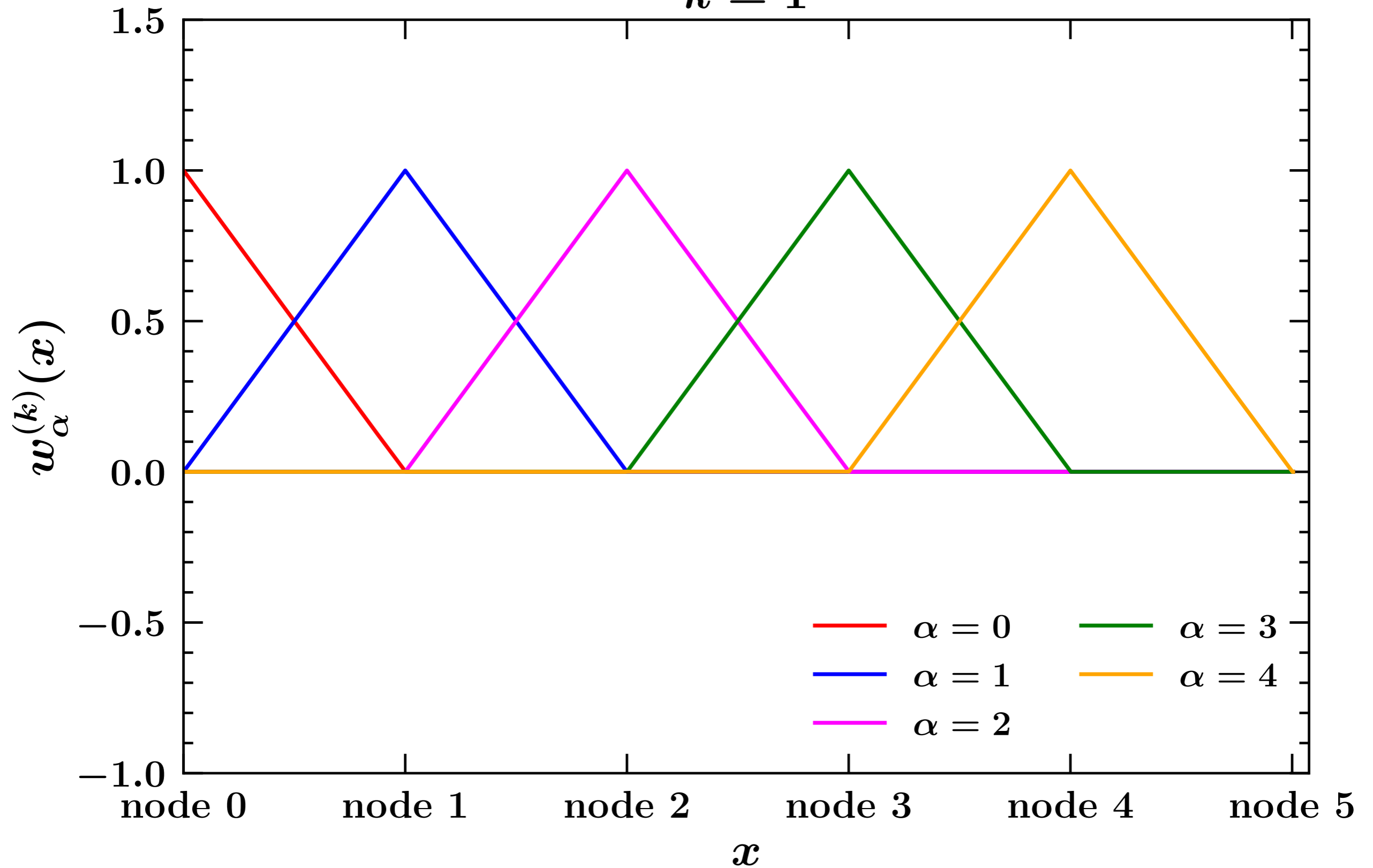
The interpolating functions

$k = 0$



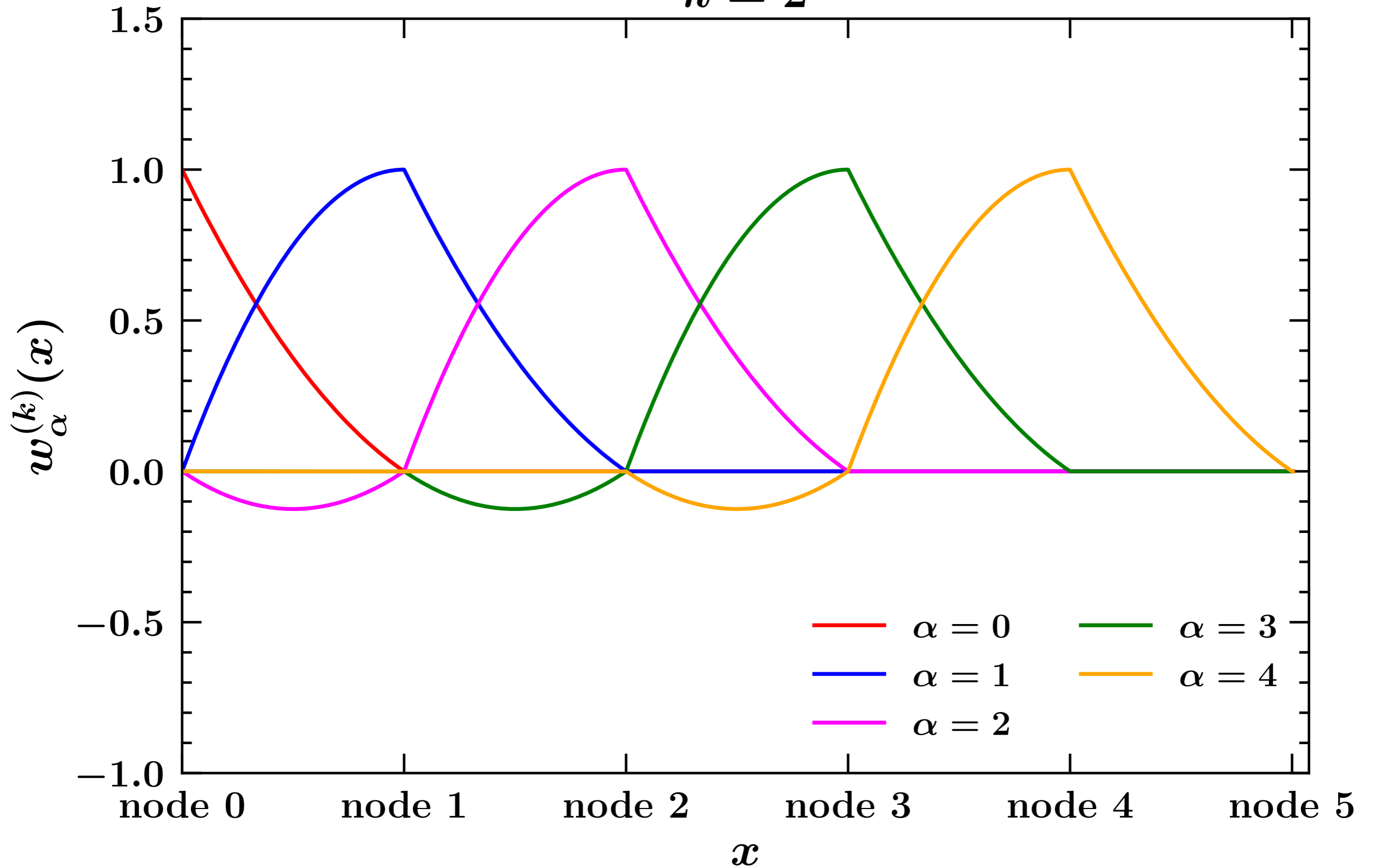
The interpolating functions

$k = 1$



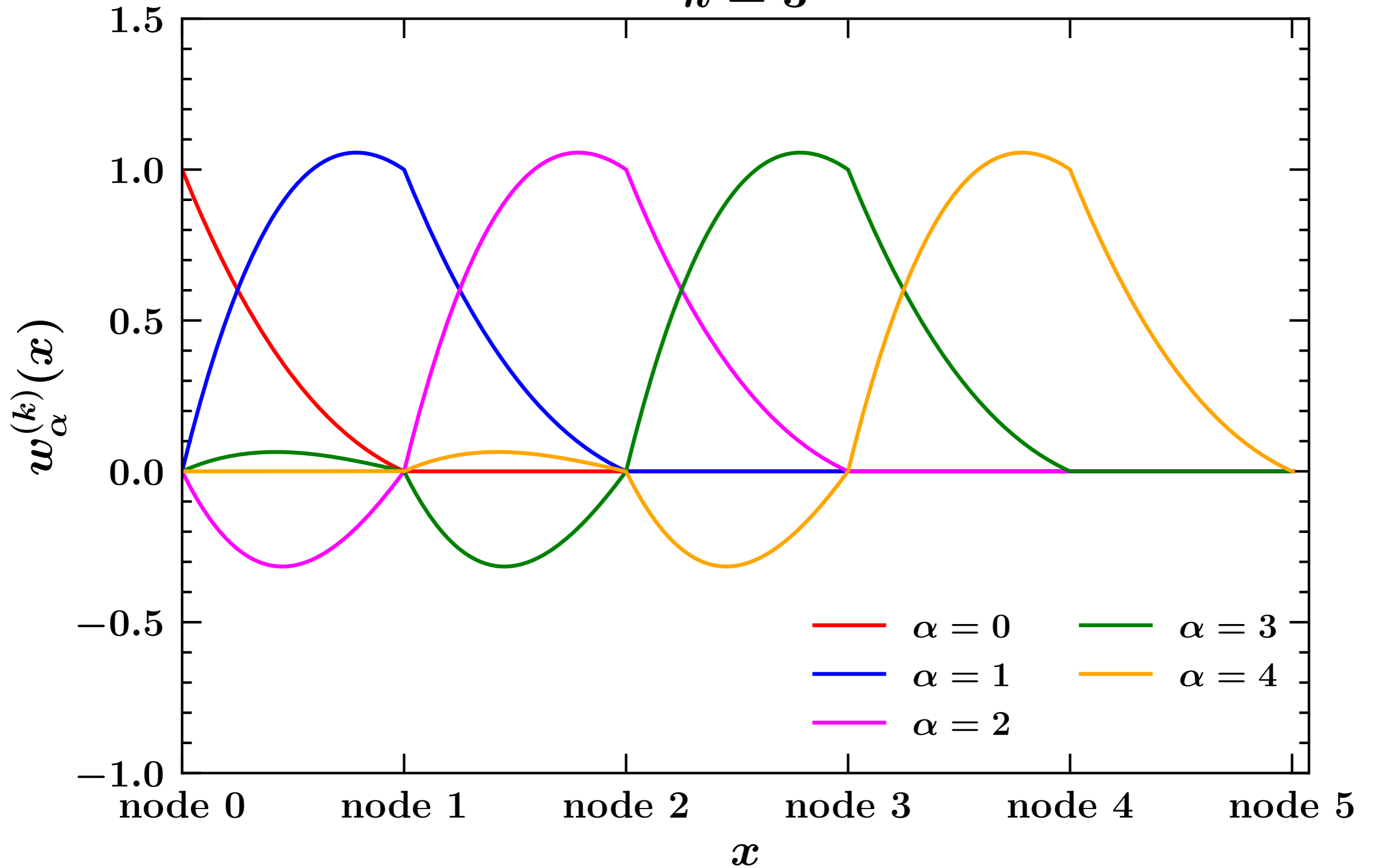
The interpolating functions

$k = 2$



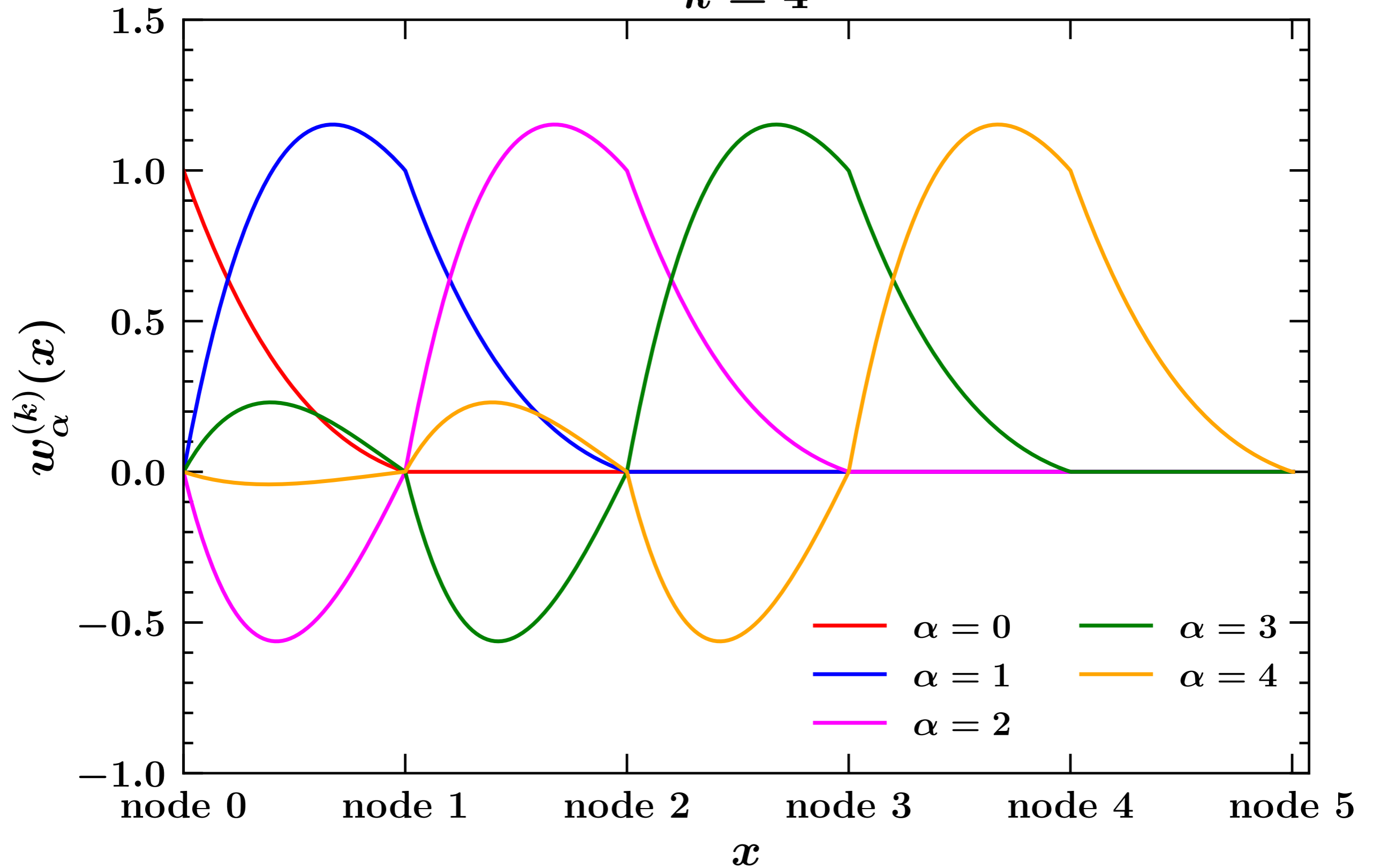
The interpolating functions

$k = 3$



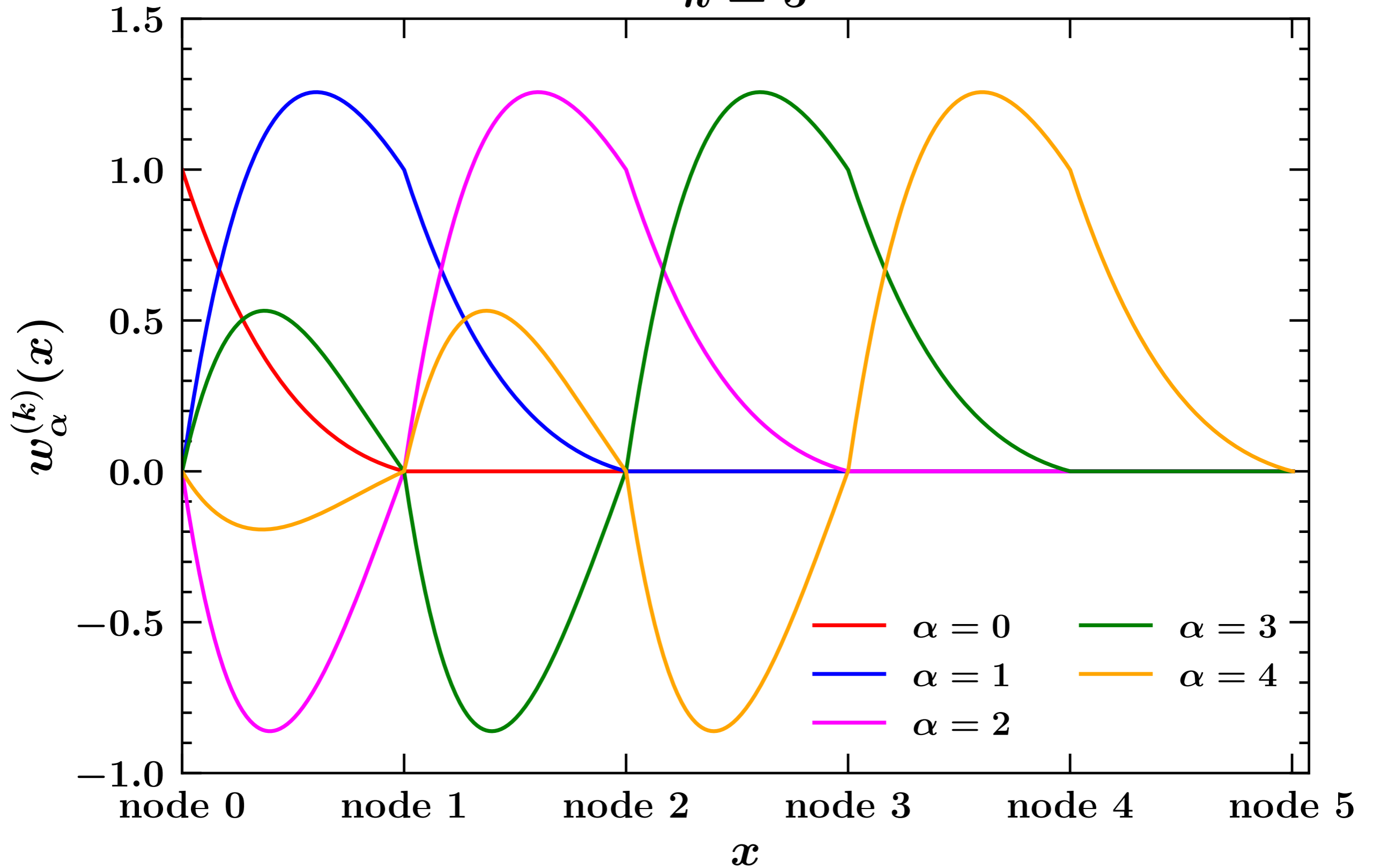
The interpolating functions

$k = 4$



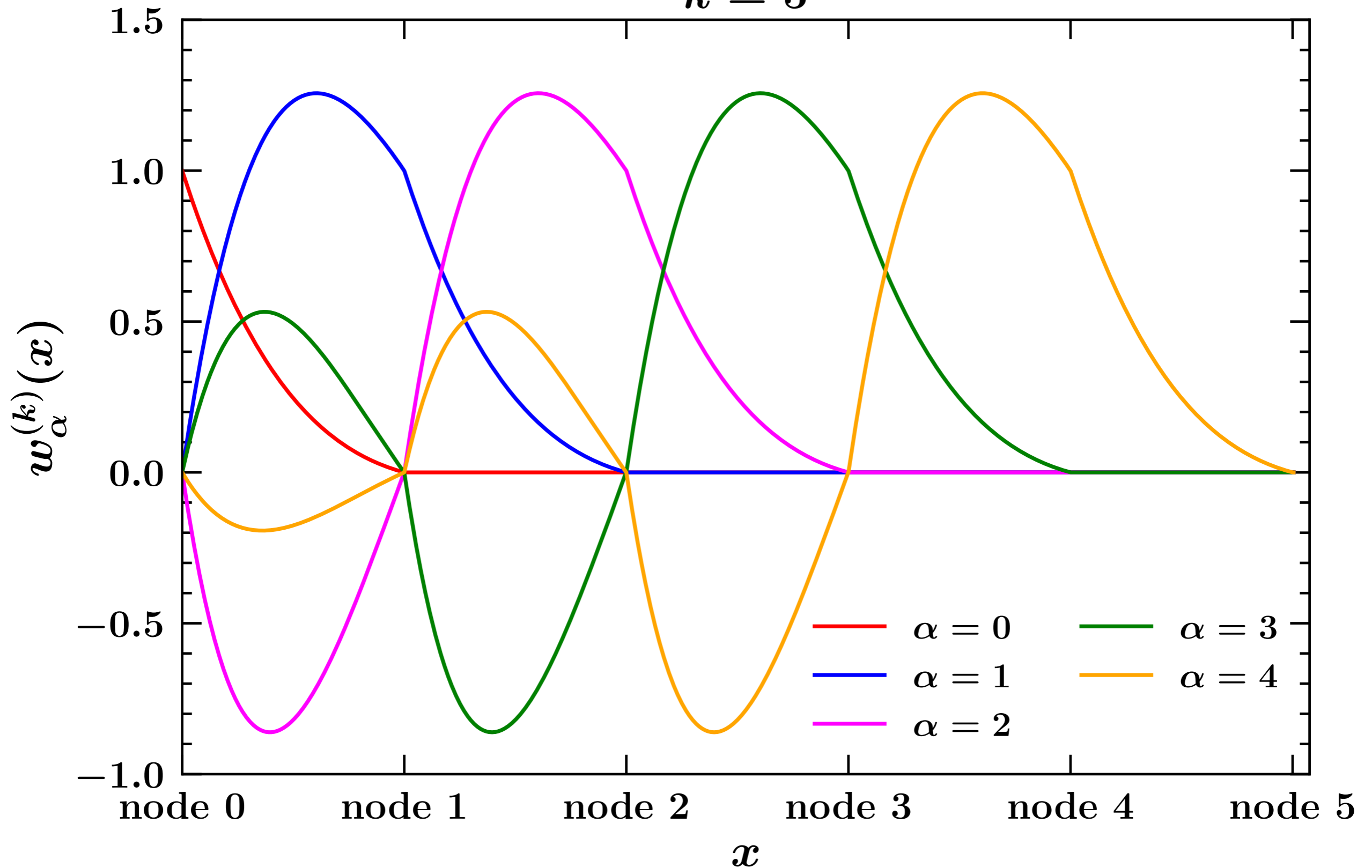
The interpolating functions

$k = 5$



The interpolating functions

$k = 5$



At any interpolation degree, the interpolating functions are such that $w_\alpha^{(k)}(x_\beta) = \delta_{\alpha\beta}$.

Interpolation test

🍏 Interpolate the function:

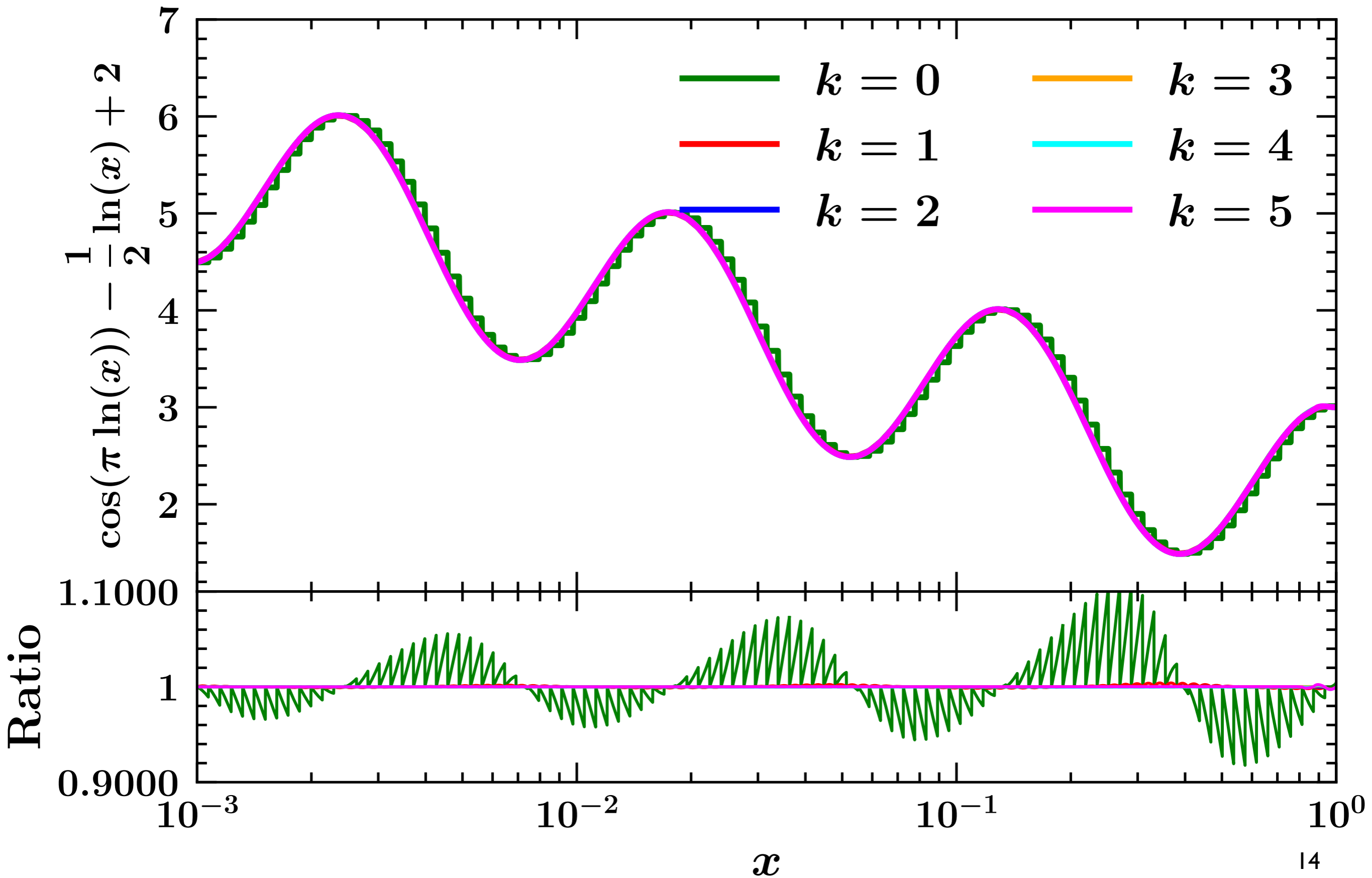
$$f(x) = \cos(\pi \ln(x)) - \frac{1}{2} \ln(x) + 2$$

🍏 on grid with 100 nodes logarithmically distributed between 10^{-3} and 1.

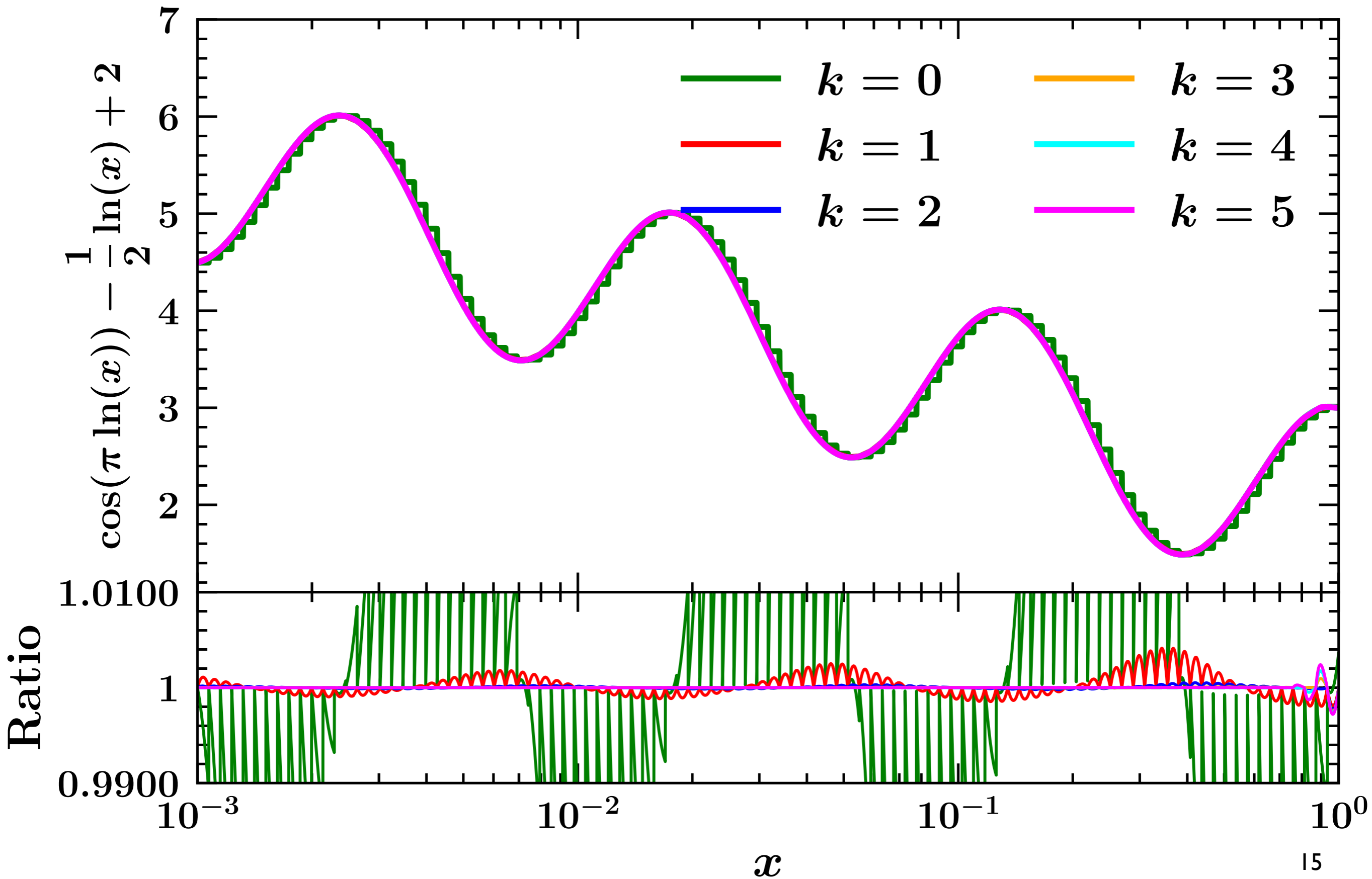
🍏 Test different degrees of the Lagrange interpolants from 0 to 5:

🍏 check the accuracy against the original function.

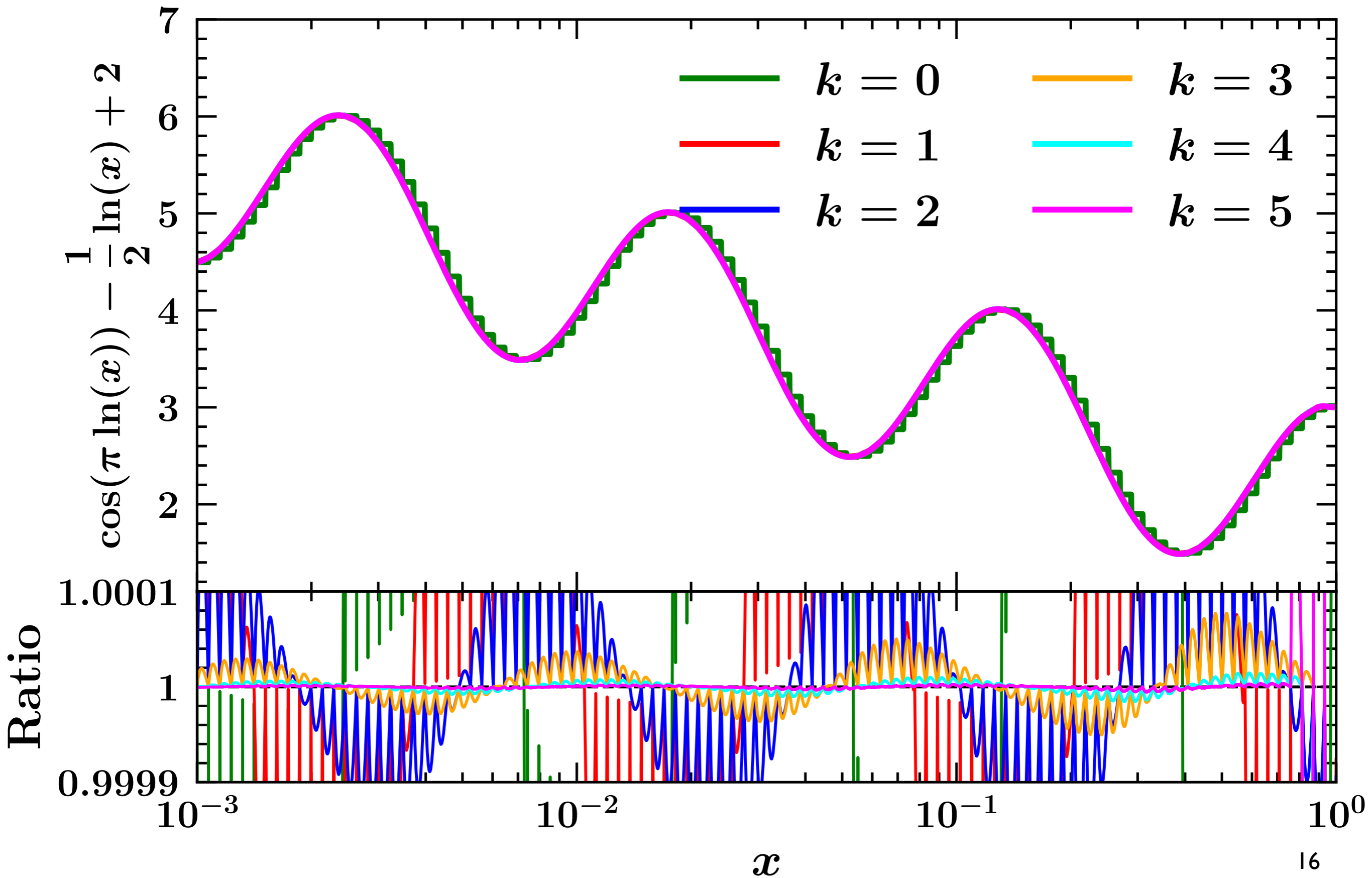
Interpolation test



Interpolation test



Interpolation test



Aside: deriving and integrating

🍏 If:

$$f(x) = \sum_{\alpha=1}^{N_x} w_{\alpha}^{(k)}(x) f(x_{\alpha})$$

🍏 then:

$$\frac{df}{dx} = \sum_{\alpha=1}^{N_x} \frac{dw_{\alpha}^{(k)}}{dx} f(x_{\alpha}) \quad \text{and} \quad \int_a^b f(x) dx = \sum_{\alpha=1}^{N_x} \left[\int_a^b w_{\alpha}^{(k)}(x) dx \right] f(x_{\alpha})$$

🍏 The derivative and the indefinite integral of the interpolation functions can be computed **analytically** for any interpolation degree.

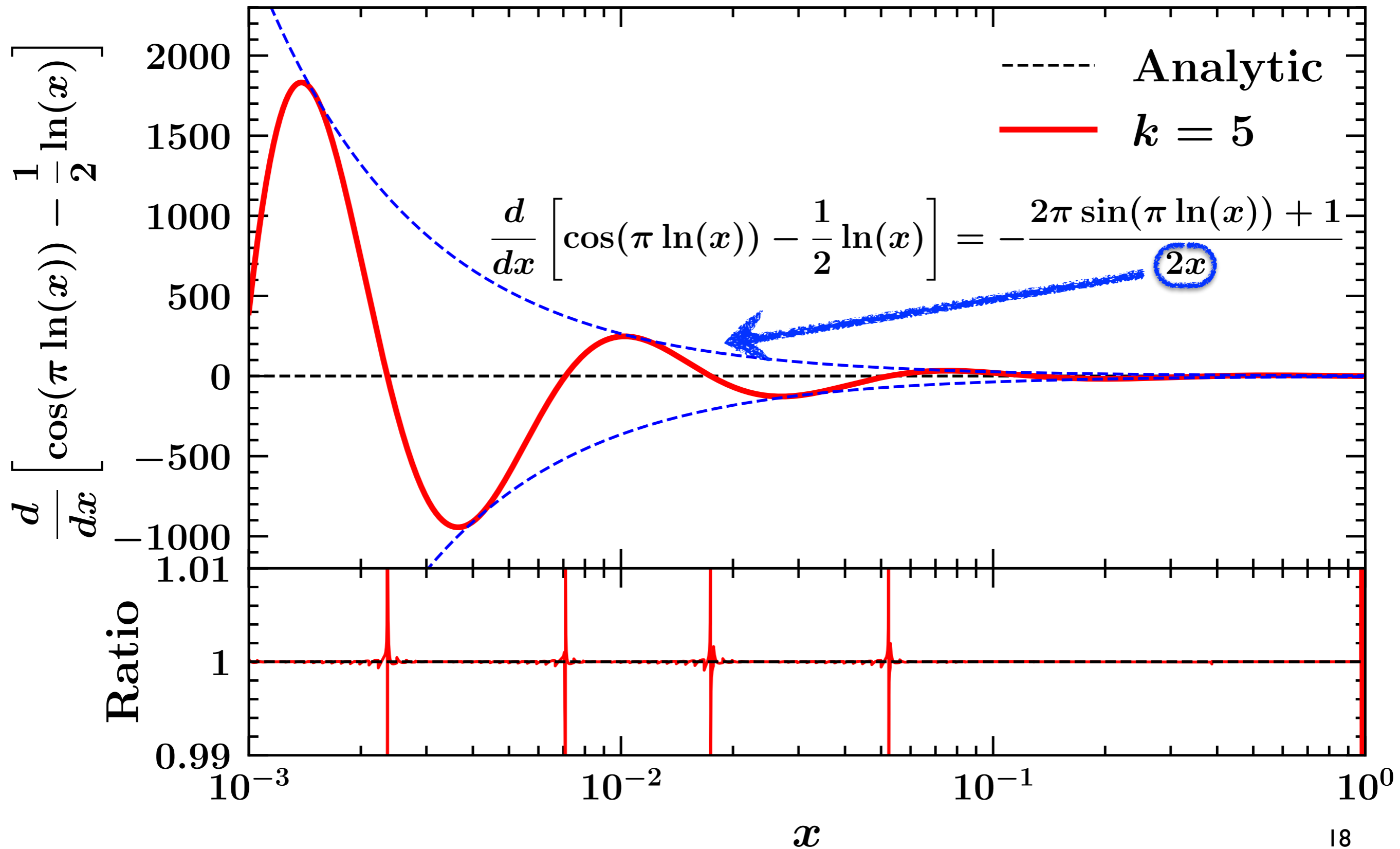
🍏 This enables one, not only to interpolate the function f , but also to compute **derivatives** and **integrals** by knowing its values on the grid.

🍏 Potential drawbacks:

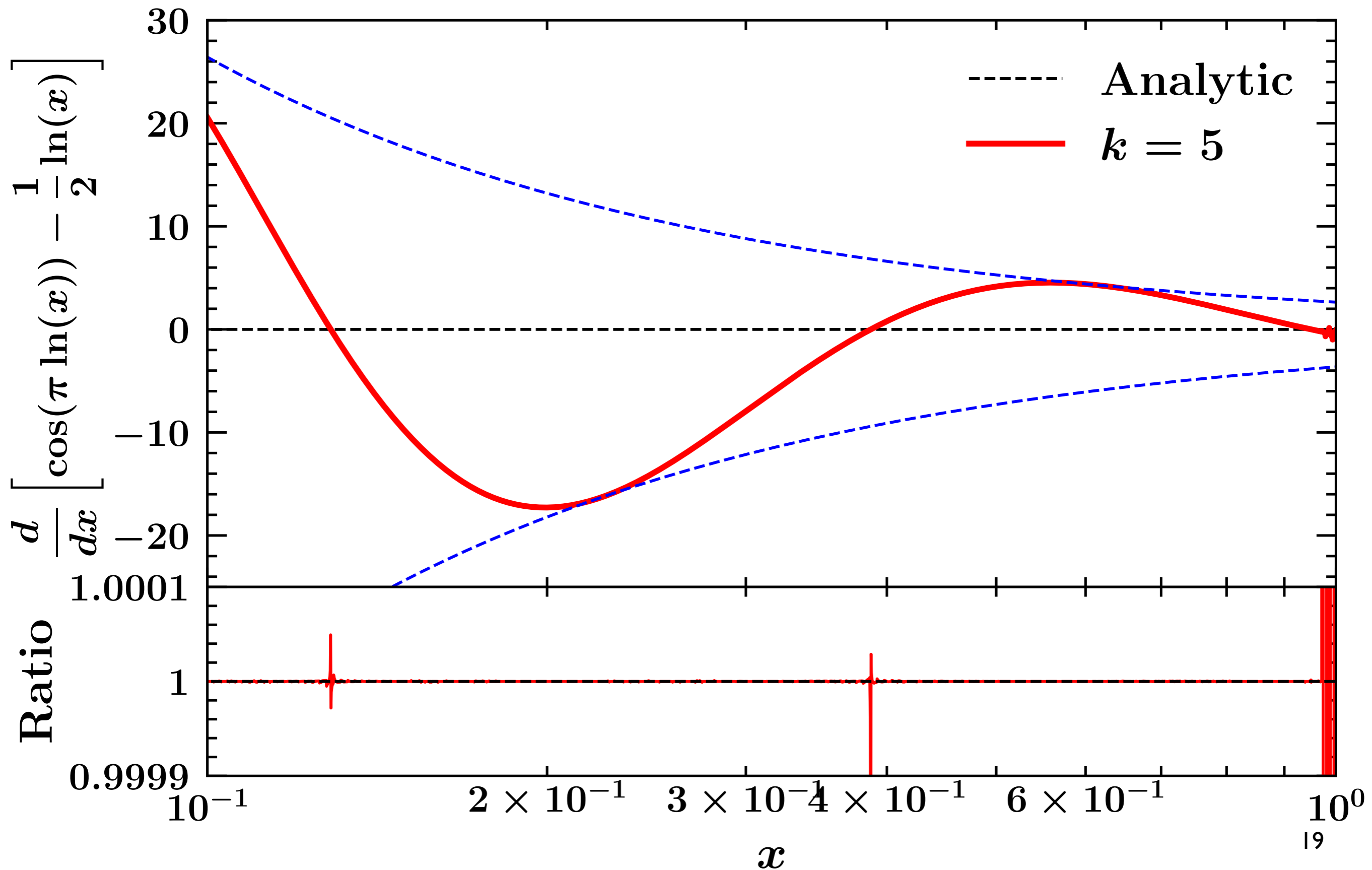
🍏 interpolation functions not smooth at the nodes, derivatives not defined there.

🍏 No direct control on the integration accuracy.

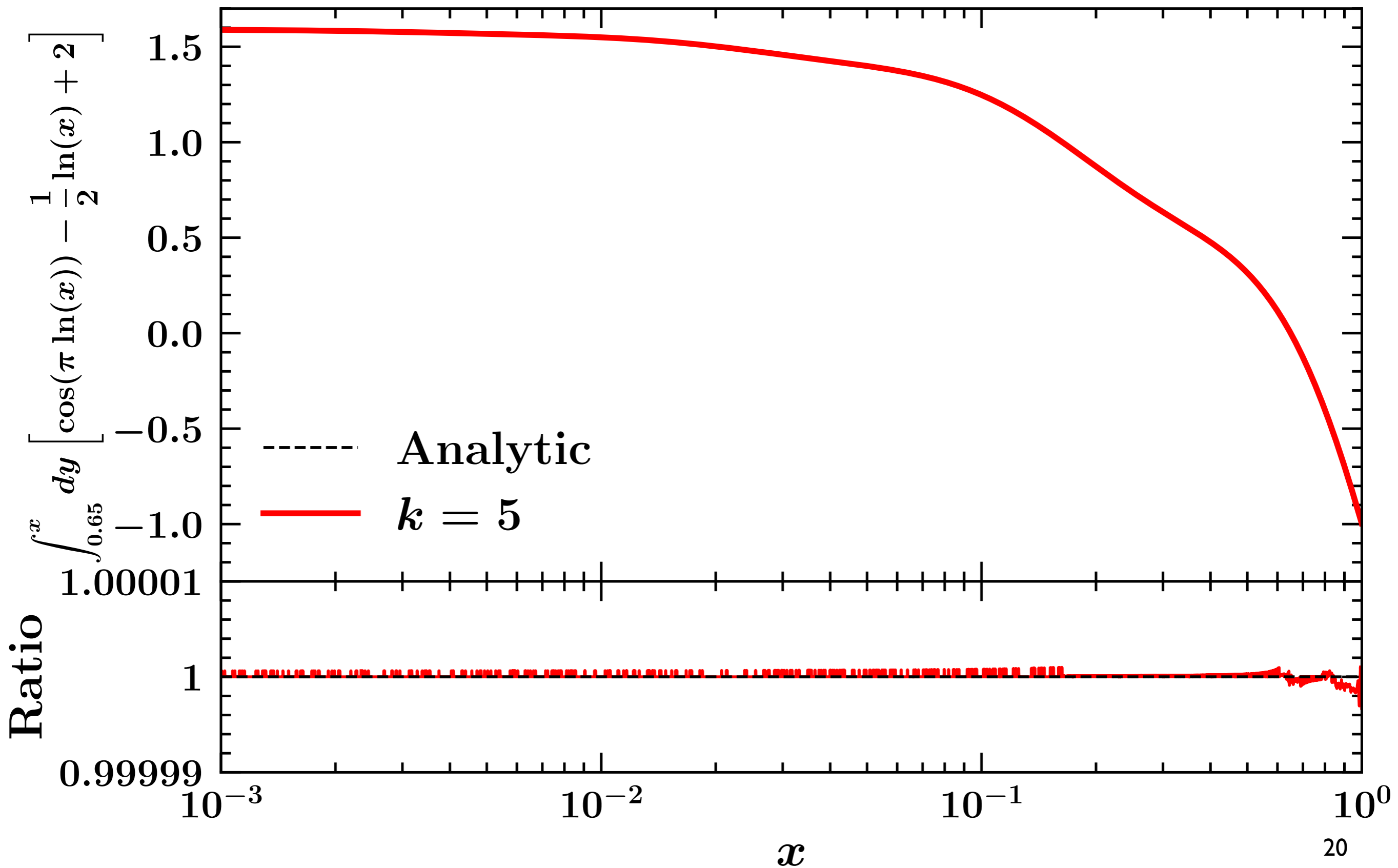
Aside: deriving and integrating



Aside: deriving and integrating



Aside: deriving and integrating



Solving the DGLAP on the grid

- 🍏 The r.h.s. of the DGLAP equations requires computing integrals of the following kind:

$$I(x, \mu) = \int_x^1 \frac{dy}{y} P(y, \alpha_s(\mu)) f\left(\frac{x}{y}, \mu\right)$$

- 🍏 Using the interpolation formula for f (the parton distribution):

$$I(\mathbf{x}_\beta, \mu) = \sum_{\alpha=0}^{N_x} \underbrace{\left[\int_{\mathbf{x}_\beta}^1 \frac{dy}{y} P(y, \alpha_s(\mu)) w_\alpha^{(k)}\left(\frac{\mathbf{x}_\beta}{y}\right) \right]}_{\Gamma_{\beta\alpha}(\mu)} f(\mathbf{x}_\alpha, \mu)$$

- 🍏 The $\Gamma_{\beta\alpha}$ don't depend on f and thus can be **precomputed** and **stored**.

- 🍏 More compactly:

$$I_\beta(\mu) = \sum_{\alpha=0}^{N_x} \Gamma_{\beta\alpha}(\mu) f_\alpha(\mu) \quad \Rightarrow \quad \mathbf{I}(\mu) = \mathbf{\Gamma}(\mu) \cdot \mathbf{f}(\mu)$$

- 🍏 This allows us to compute I on each point of the grid.

Solving the DGLAP on the grid

- 🍏 The DGLAP equations reduce to a set of coupled **linear ordinary differential equations** (ODEs) that in vectorial notation read:

$$\frac{d\mathbf{f}(\mu)}{d \ln \mu^2} = \mathbf{\Gamma}(\mu) \cdot \mathbf{f}(\mu) \equiv \mathbf{F}(\mu, \mathbf{f})$$

- 🍏 The functions to be determined are the partonic distribution f as a function of μ on each node of the grid knowing $\mathbf{f}(\mu_0)$ on the grid.
- 🍏 Many algorithms to solve ODEs exist. A particularly popular choice is the **4th order Runge-Kutta** (RK4):

$$\mathbf{f}_{n+1} = \mathbf{f}_n + \frac{1}{6} (\mathbf{k}_1 + 2\mathbf{k}_2 + 2\mathbf{k}_3 + \mathbf{k}_4) + \mathcal{O}(h^5)$$

$$\mathbf{k}_1 = h\mathbf{F}(\mu_n, \mathbf{f}_n)$$

$$\mathbf{k}_2 = h\mathbf{F}\left(\mu_n + \frac{h}{2}, \mathbf{f}_n + \frac{\mathbf{k}_1}{2}\right)$$

$$\mathbf{k}_3 = h\mathbf{F}\left(\mu_n + \frac{h}{2}, \mathbf{f}_n + \frac{\mathbf{k}_2}{2}\right)$$

$$\mathbf{k}_4 = h\mathbf{F}(\mu_n + h, \mathbf{f}_n + \mathbf{k}_3)$$

- 🍏 Tabulate \mathbf{f} on a **grid in μ** that can successively interpolated.

- 🍏 Final result: $f(x, \mu)$ is know on a 2D grid in x and μ .

Interpolation on a log grid

🍏 In principle, one needs to compute N_x^2 integrals $\Gamma_{\beta\alpha}$.

🍏 However, if one uses a logarithmically distributed grid:

$$x_{\alpha+1} = e^{\delta x} x_{\alpha} \quad \delta x = \text{constant}$$

🍏 and the Lagrange polynomials are expressed in **powers of** $\ln(x)$:

$$w_{\alpha}^{(k)}(x) = \sum_{i=0, i \leq \alpha}^k \theta(x - x_{\alpha-i}) \theta(x_{\alpha-i+1} - x) \prod_{m=0, m \neq i}^k \frac{\ln x - \ln x_{\alpha-i+m}}{\ln x_{\alpha} - \ln x_{\alpha-i+m}}$$

🍏 one finds:

$$\Gamma_{\beta\alpha} = \begin{cases} \Gamma_{0, \alpha-\beta} & \alpha \geq \beta \\ 0 & \alpha < \beta \end{cases} \quad \Gamma_{\beta\alpha} = \begin{pmatrix} a_0 & a_1 & a_2 & \cdots & a_{N_x} \\ 0 & a_0 & a_1 & \cdots & a_{N_x-1} \\ 0 & 0 & a_0 & \cdots & a_{N_x-2} \\ \vdots & \vdots & \vdots & \ddots & \vdots \\ 0 & 0 & 0 & \cdots & a_0 \end{pmatrix}$$

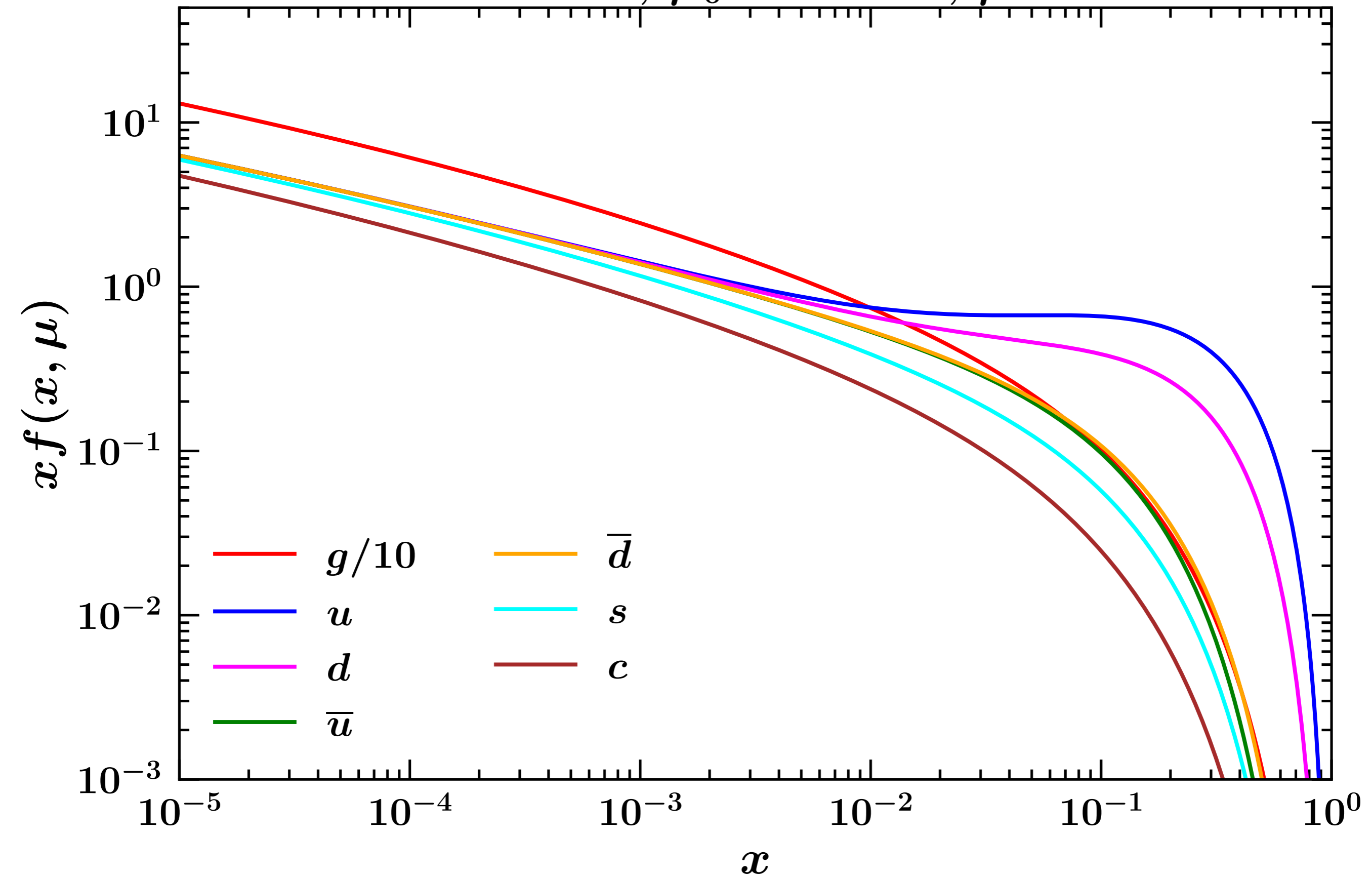
🍏 The number of integrals reduces to N_x and so thus the computation time.

🍏 Problem: a logarithmic grid becomes increasingly sparse at large x :

🍏 solution: concatenate grids with increasing density as one moves towards large x .

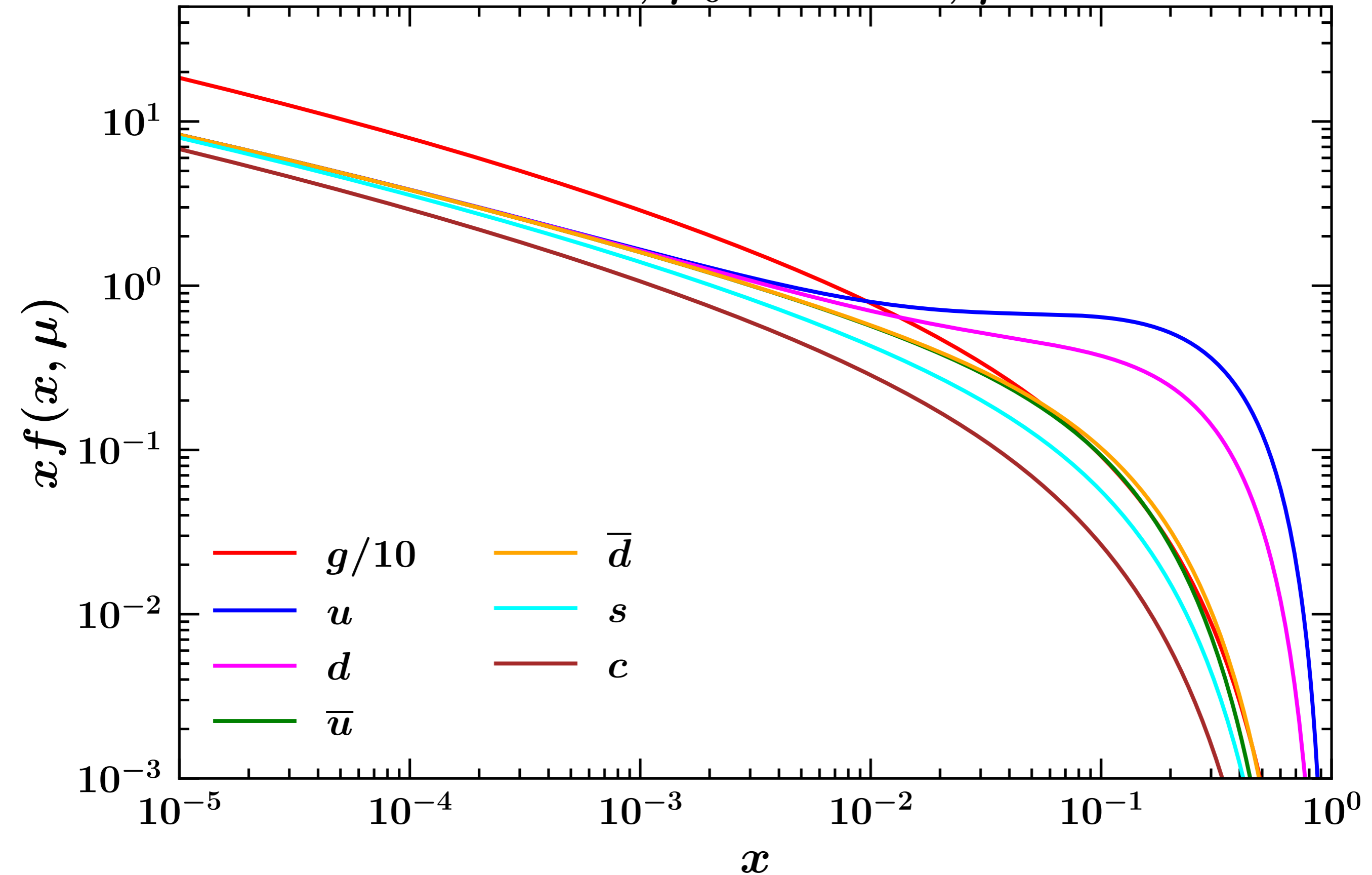
Numerical results

NNLO evolution, $\mu_0 = 1 \text{ GeV}$, $\mu = 10 \text{ GeV}$



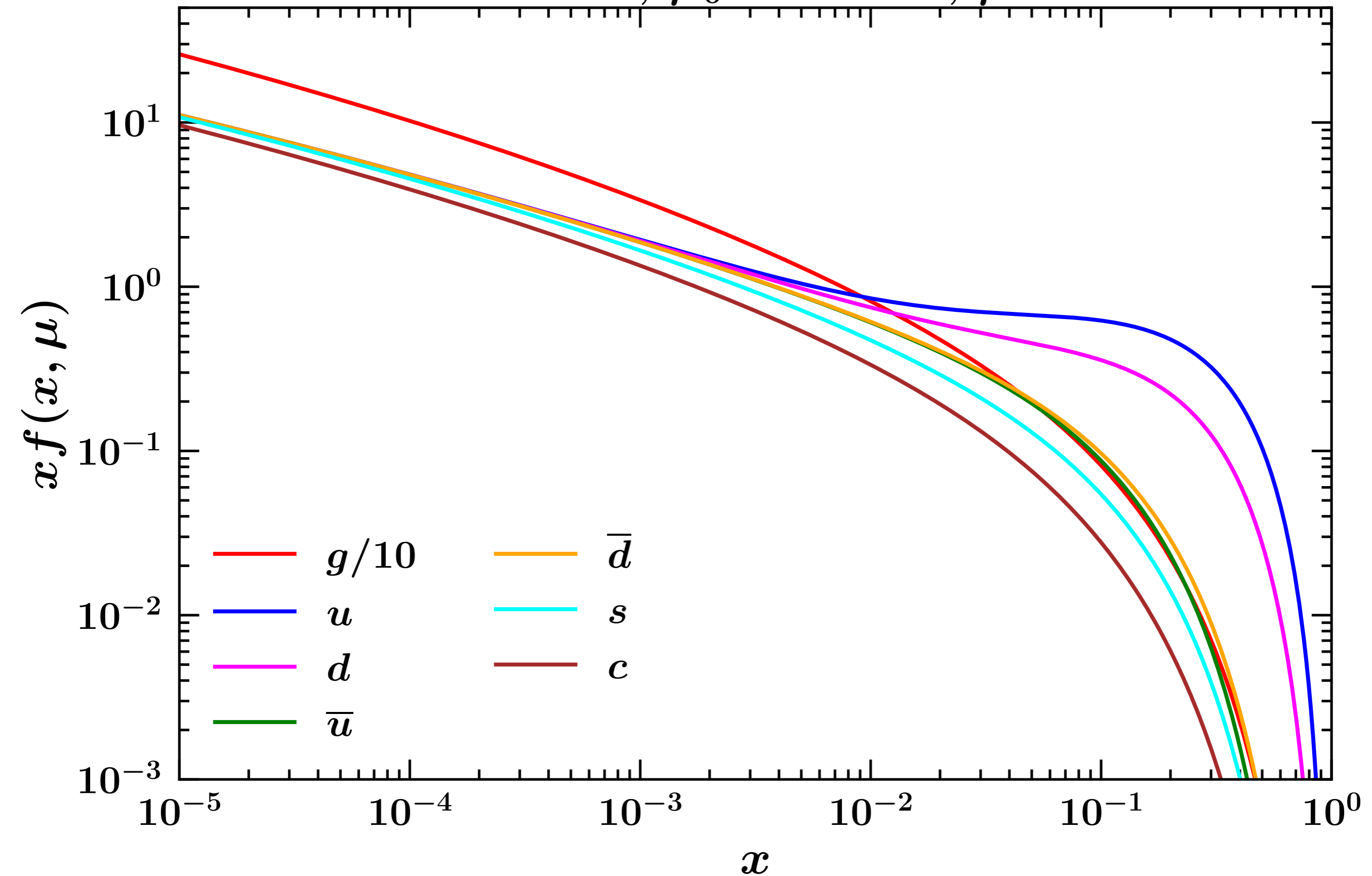
Numerical results

NNLO evolution, $\mu_0 = 1 \text{ GeV}$, $\mu = 20 \text{ GeV}$



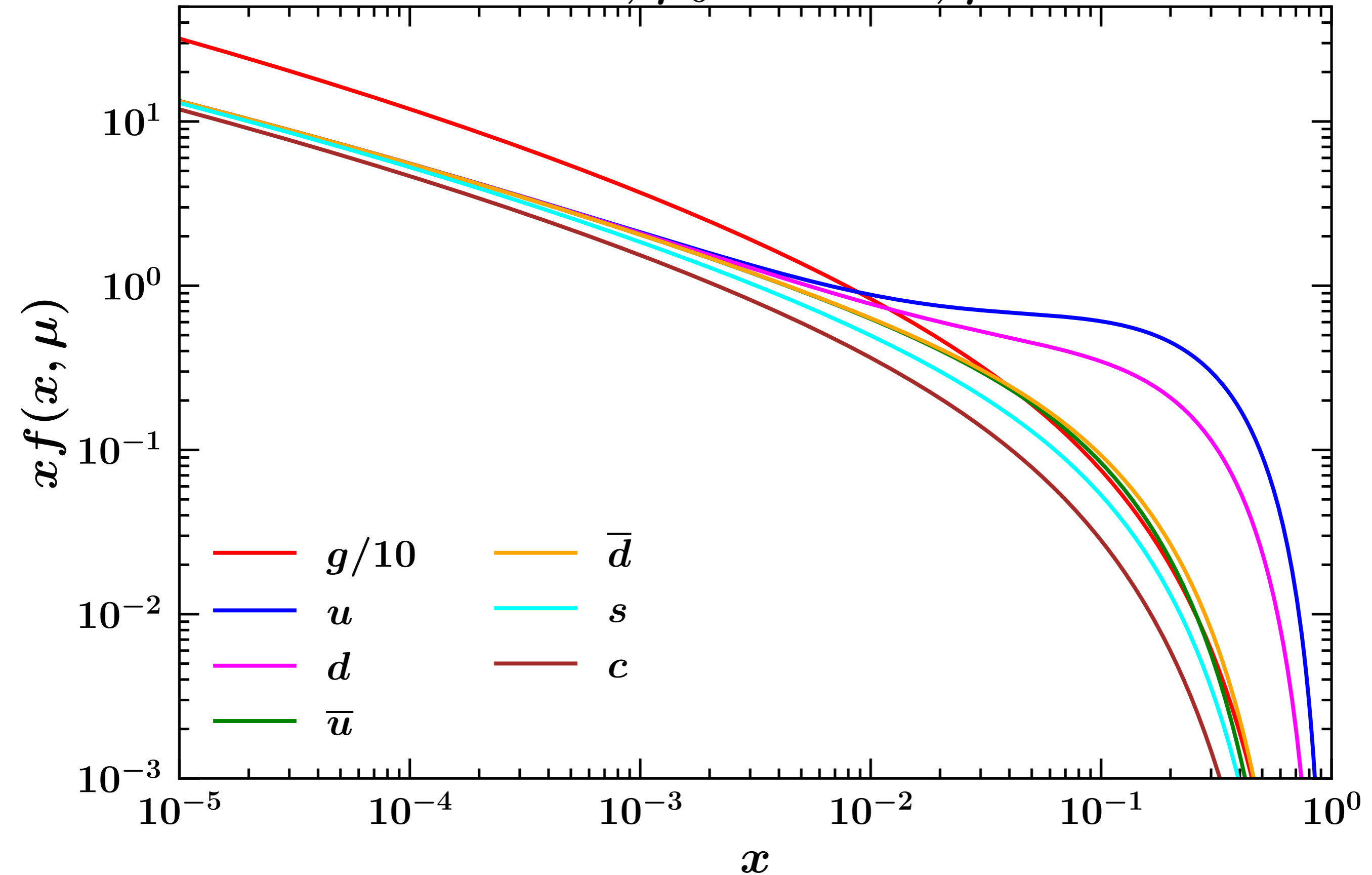
Numerical results

NNLO evolution, $\mu_0 = 1 \text{ GeV}$, $\mu = 50 \text{ GeV}$



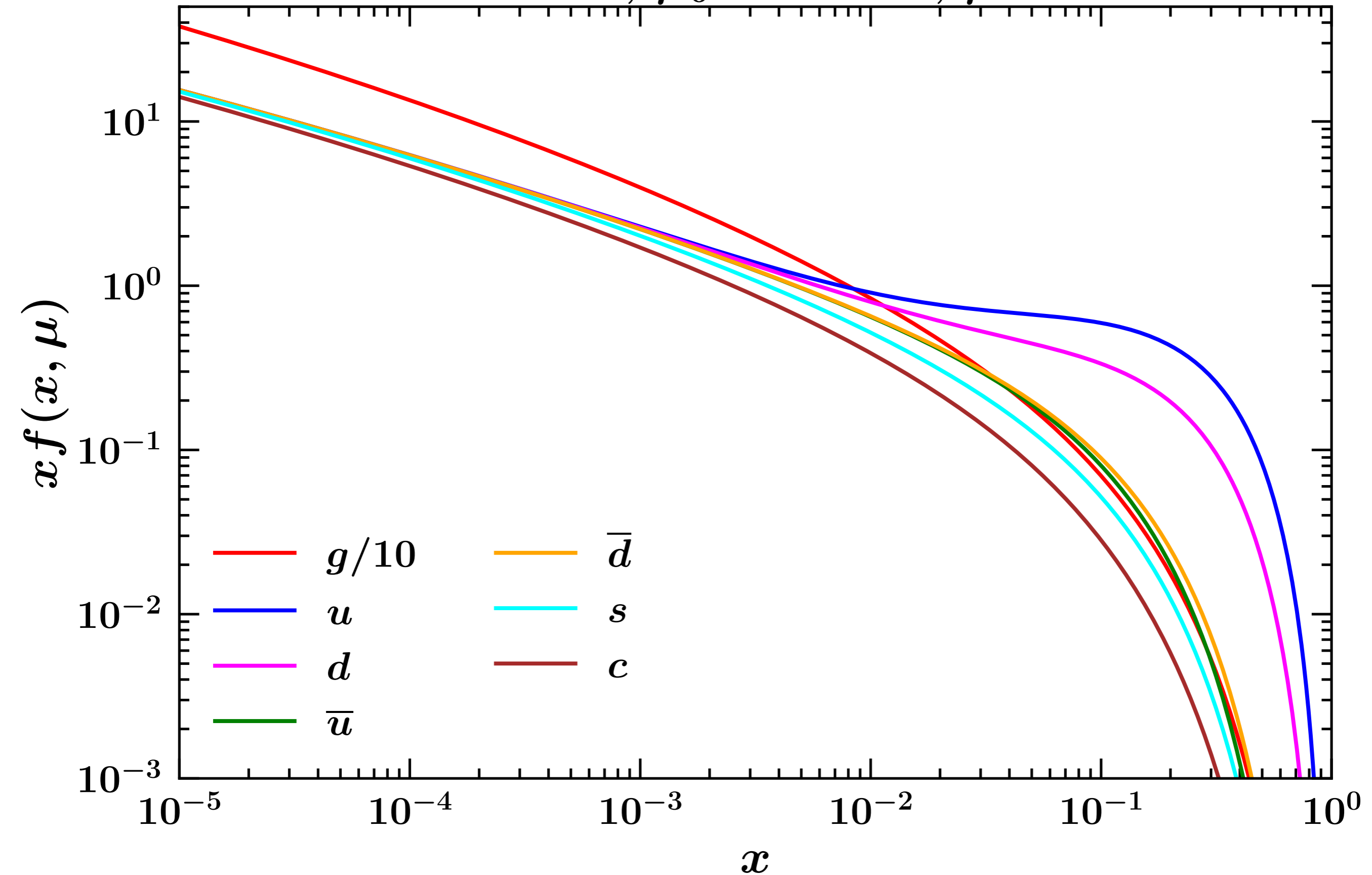
Numerical results

NNLO evolution, $\mu_0 = 1 \text{ GeV}$, $\mu = 100 \text{ GeV}$



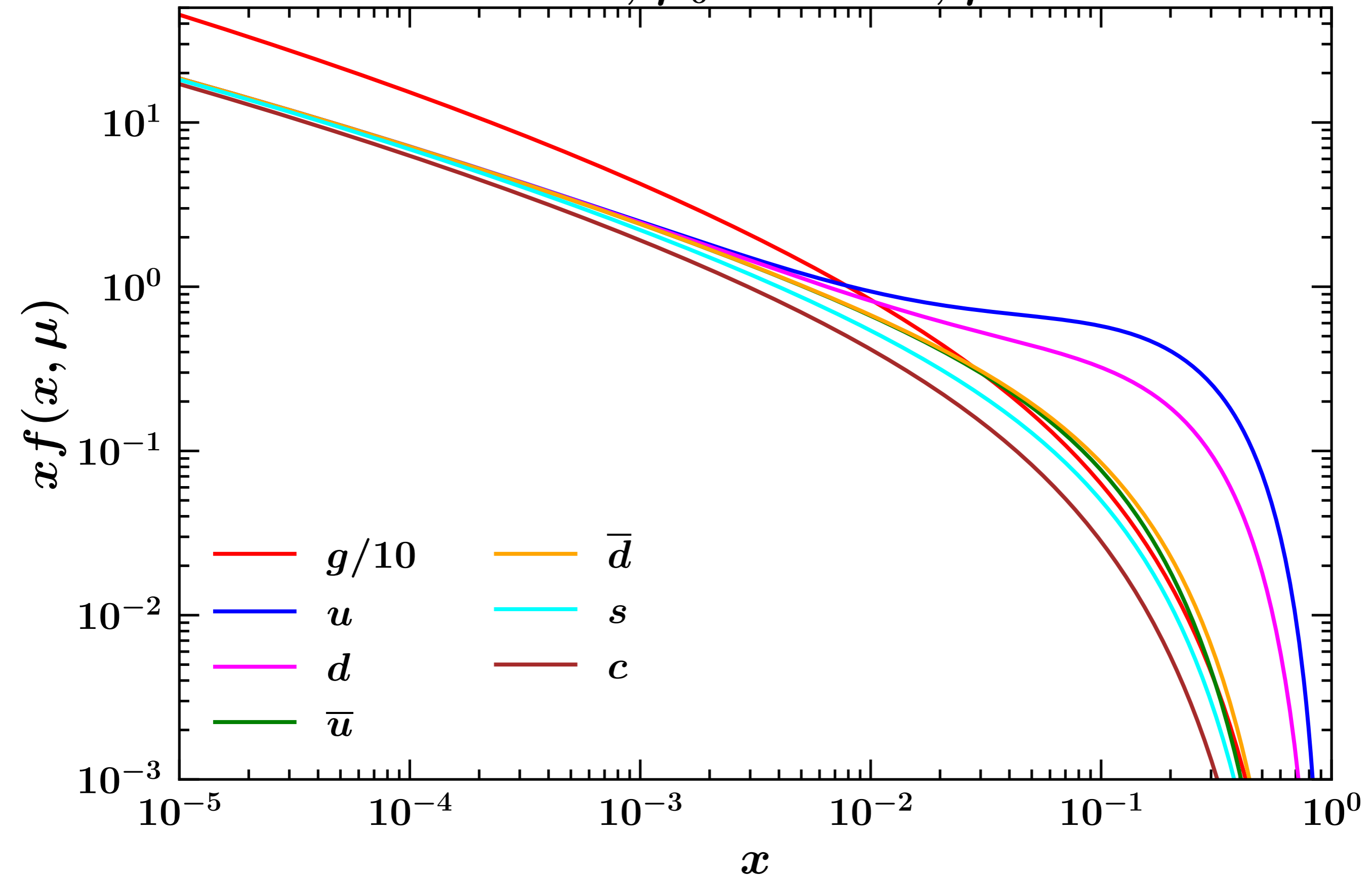
Numerical results

NNLO evolution, $\mu_0 = 1 \text{ GeV}$, $\mu = 200 \text{ GeV}$



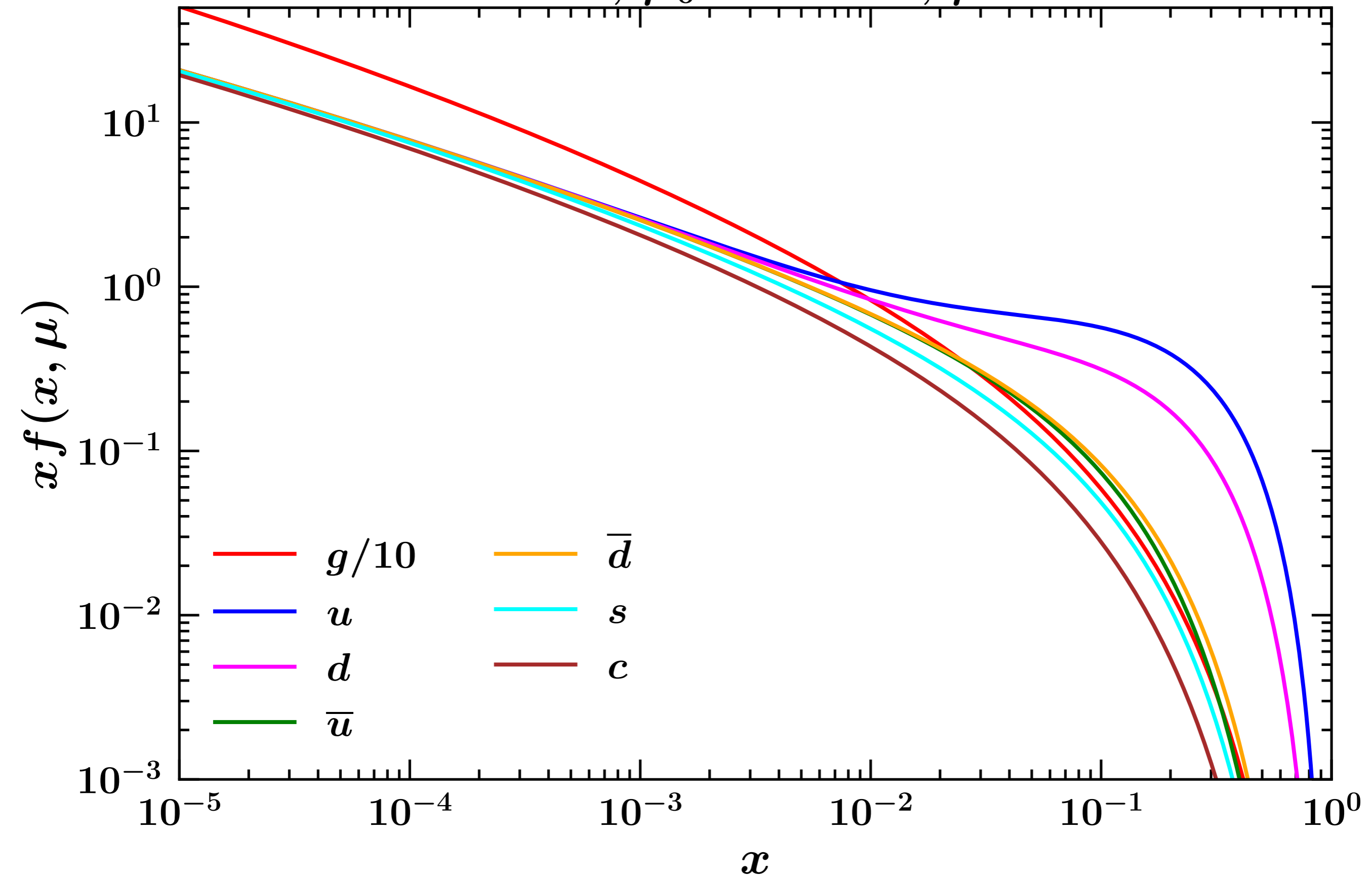
Numerical results

NNLO evolution, $\mu_0 = 1 \text{ GeV}$, $\mu = 500 \text{ GeV}$

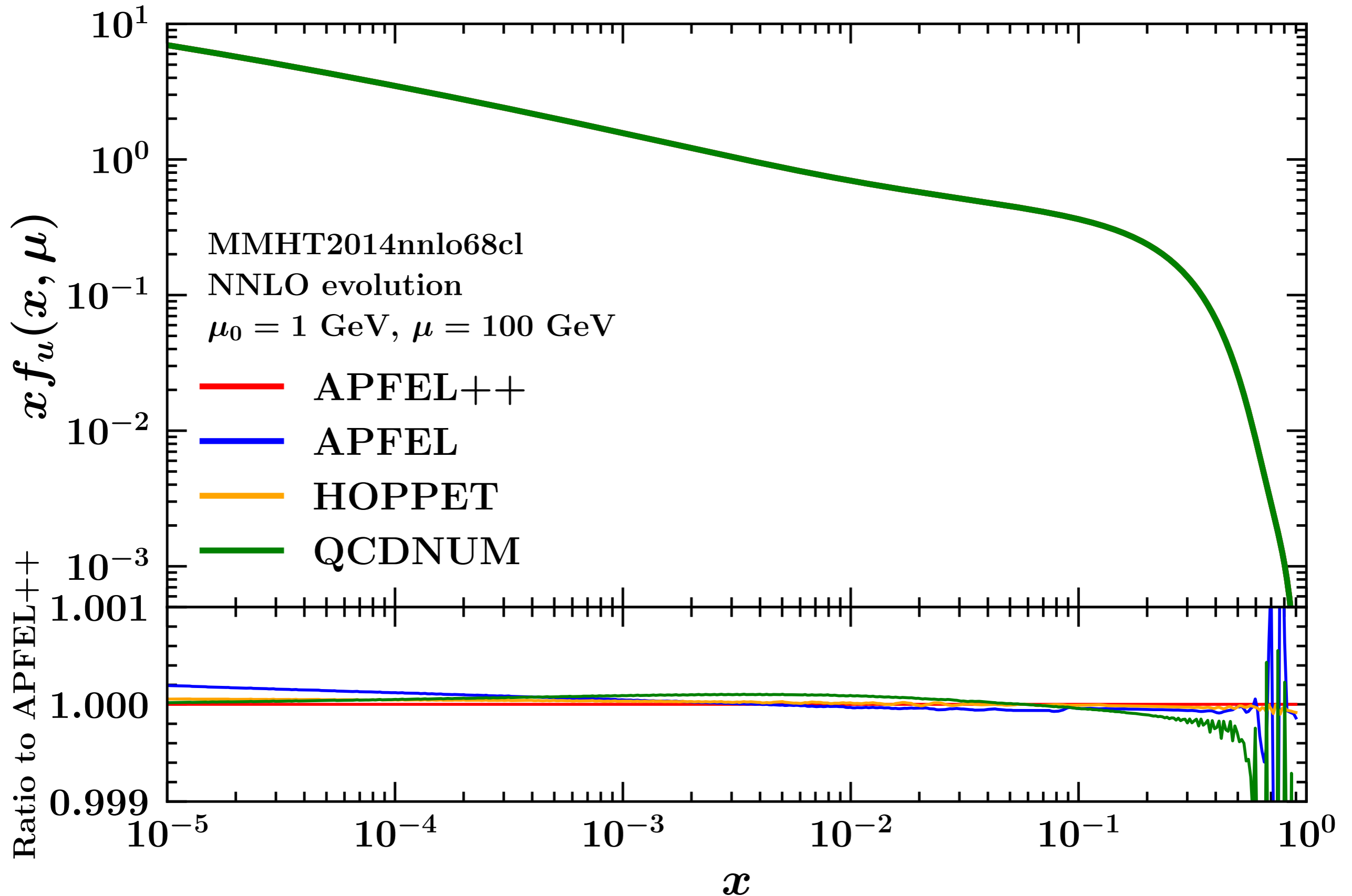


Numerical results

NNLO evolution, $\mu_0 = 1 \text{ GeV}$, $\mu = 1000 \text{ GeV}$



Comparison amongst codes



🍏 Independent implementations agree well below the per-mil level!

“Details” I didn’t discuss

- 🍏 Computing the integrals $\Gamma_{\beta\alpha}$:
 - 🍏 integrands are distributions that involve δ -functions and +-distributions.
- 🍏 Addressing the problem of inaccuracies at large x caused by a logarithmically-spaced grid.
- 🍏 The flavour dependence of parton distributions.
- 🍏 The possible presence of heavy-quark thresholds along the evolution.
- 🍏 The evolution equation for generalised parton distributions (GPDs):
 - 🍏 complications due to skewed kinematics,
 - 🍏 more complicated convolution integrals (principal-valued distributions and spurious divergences),
 - 🍏 logarithmically-spaced grid does not help reduce the number of integrals.
- 🍏 Alternative: the Mellin-moment approach.

Conclusions

- 🍏 Evolution equations in QCD:
 - 🍏 allow one to connect parton distributions at different scales,
 - 🍏 when fitting parton distributions to data, we can parametrise them at one single scale and obtain them at any other scale,
 - 🍏 fast evolution codes are required.
- 🍏 Solving evolution equations in QCD is a numerically demanding task:
 - 🍏 integro-differential equations in flavour space,
 - 🍏 discretising the problem allows one reduce the problem to a set of coupled ODEs,
 - 🍏 this however requires computing many integrals,
 - 🍏 A smart choice of the grid and of the interpolation procedure allows one to lessen the numerical burden, finally reaching very high performance and accuracy.
- 🍏 As of today, there exist several evolution codes whose mutual numerical agreement is remarkably good. Proof of the solidity of the procedure!